



講演番号：17aE523-2

CALETによる陽子、ヘリウム比観測の最新結果

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早大理工総研, 小林兼好, 鳥居祥二, 赤池陽水,
他CALETチーム

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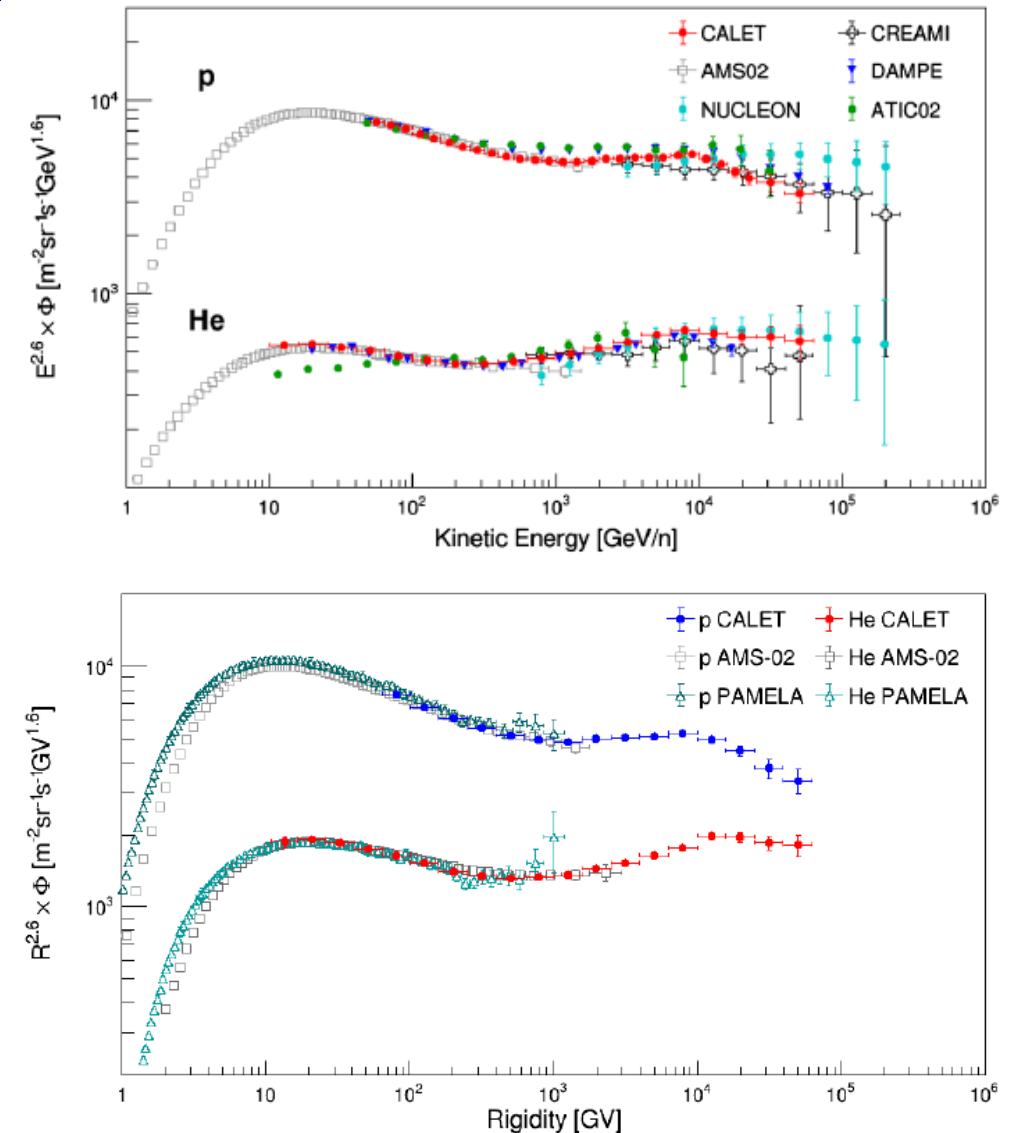




Motivation

- Proton and helium spectra and also proton/helium spectral ratio could help to understand the cosmic ray source, acceleration mechanism, and propagation.
- CALET published measurements of the proton flux ($50 \text{ GeV} < E < 60 \text{ TeV}$, PRL2022) and helium flux ($40 < E < 250 \text{ TeV}$, PRL2023) and confirmed a hardening and found softening.

proton and helium flux at PRL2023



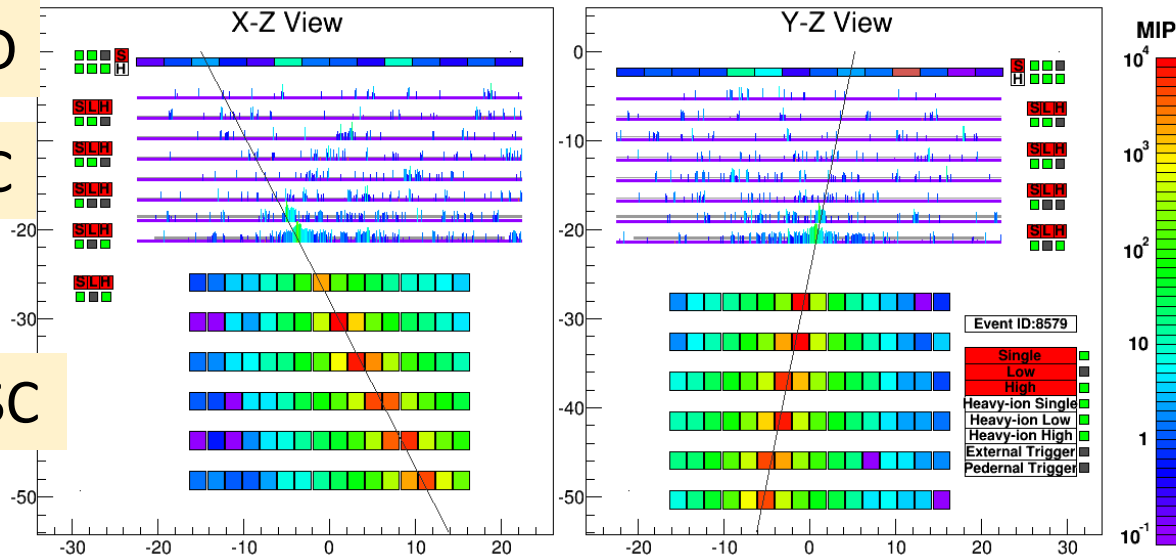
Event examples

Proton, $\Delta E=2.89$ TeV

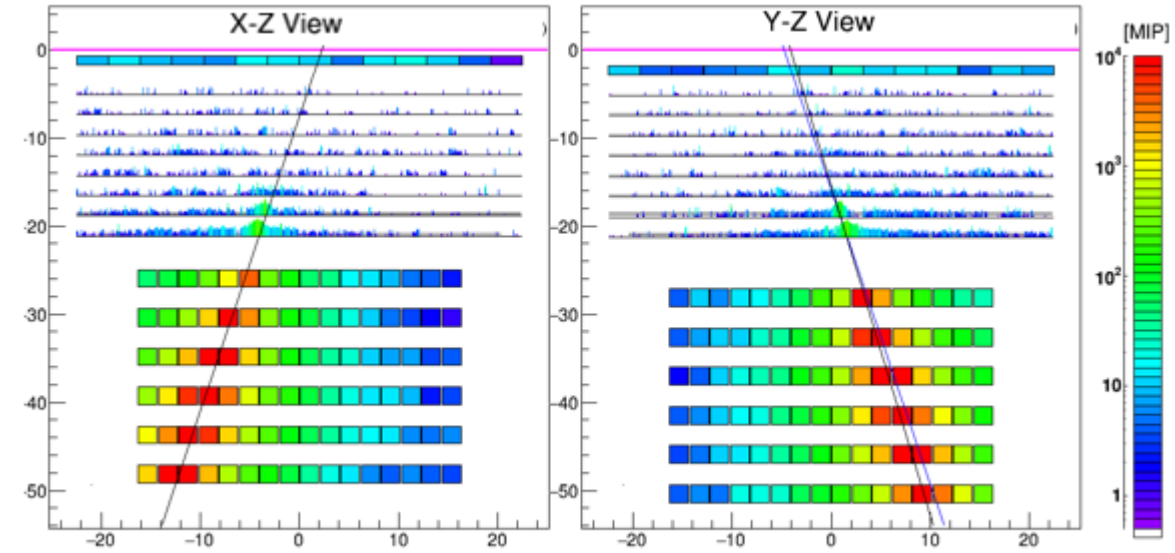
CHD

IMC

TASC



Helium, $\Delta E=11.6$ TeV



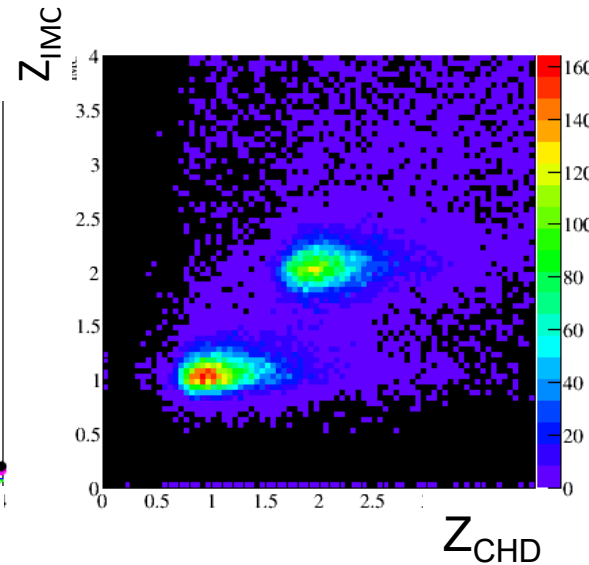
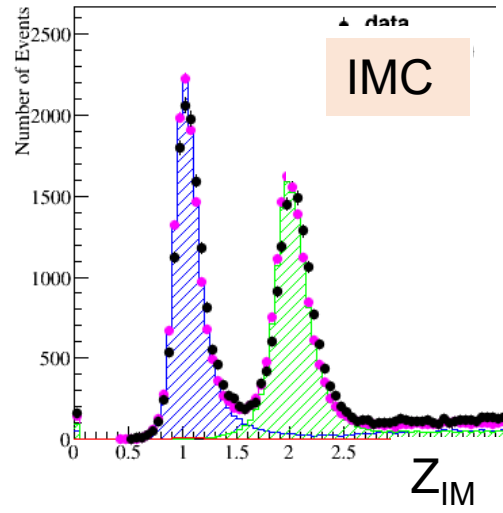
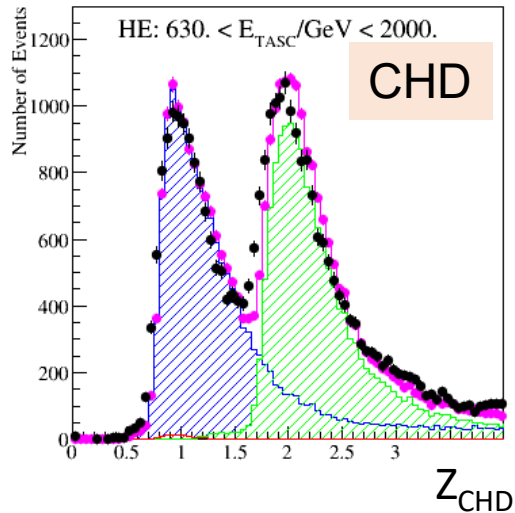


Proton Event Selection

selection	Brief description
1. Event trigger	HE trigger in $E > 300 \text{ GeV}$ and LE trigger in $E < 300 \text{ GeV}$ are required.
2. Geometrical acceptance	Track going through the detector from the top to the bottom is selected.
3. Track quality cut	Reliability of Kalman Filter tracking in IMC is checked.
4. Electron rejection	Electron events are rejected using the energy deposit within one Moliere radius along the track.
5. Off-acceptance cut	Residual events crossing the detector from the sides are rejected.
6. TASC hit consistency	In order to reject events with mis-reconstructed tracks, we require events to have energy deposits at the top X/Y layer of TASC at positions consistent with the track as reconstructed in the IMC.
7. Shower start in IMC	Shower development starting in IMC is required.
8. Charge identification in CHD and IMC	Charge identification using the energy deposit in CHD and IMC (before shower development starts) is performed to reject helium events, mainly.

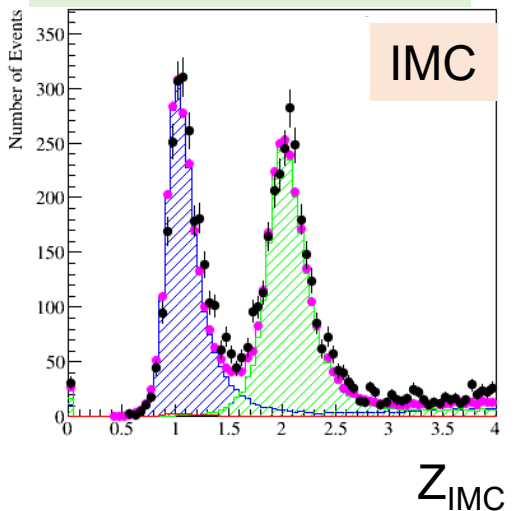
Charge Identification in CHD and IMC

HE: $630 < E < 2000 \text{ GeV}$

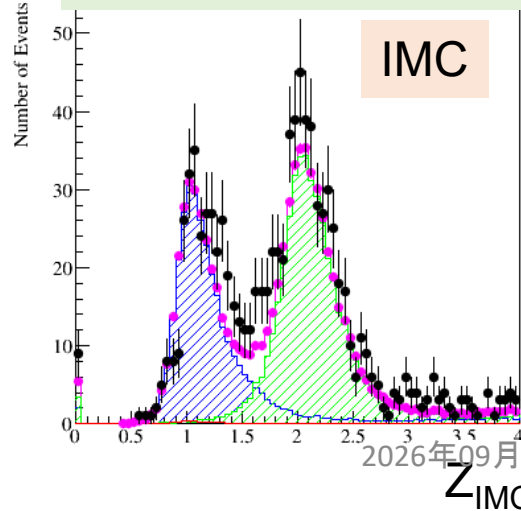


- Using the two charge identification parameters (Z_{CHD} and Z_{IMC}), proton and helium can be clearly separated.
- Total background contaminations are less than 13% in HE sample ($630 < E < 2000 \text{ GeV}$).

HE: $2 < E < 6.3 \text{ TeV}$



HE: $6.3 < E < 20 \text{ TeV}$

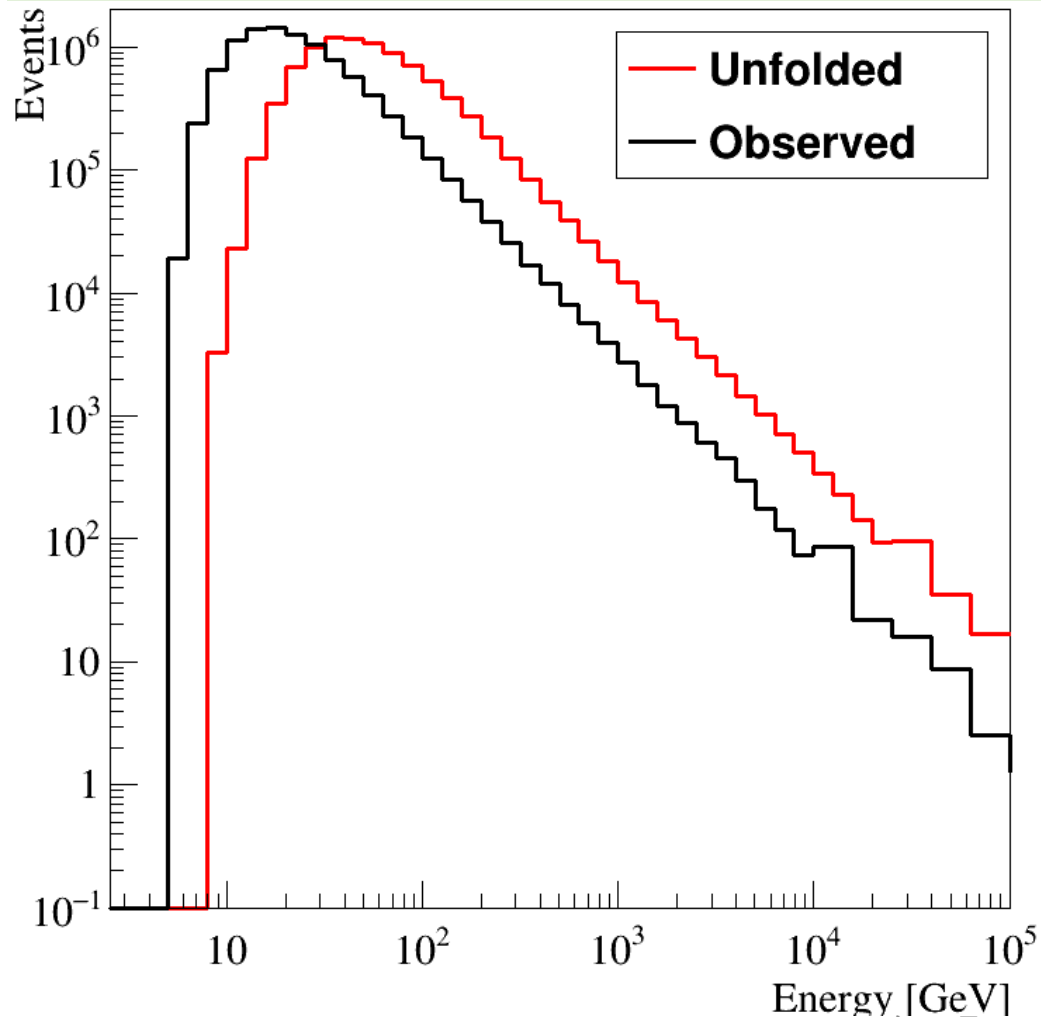


- Data
- MC all (EPICS)
- ▨ MC proton (EPICS)
- ▨ MC He (EPICS)
- ▨ MC electron (EPICS)

- Although charge identification using the CHD alone doesn't suffice for the higher energy region, identification including charge measurement in the IMC clearly separates p/He

Energy Unfolding (proton)

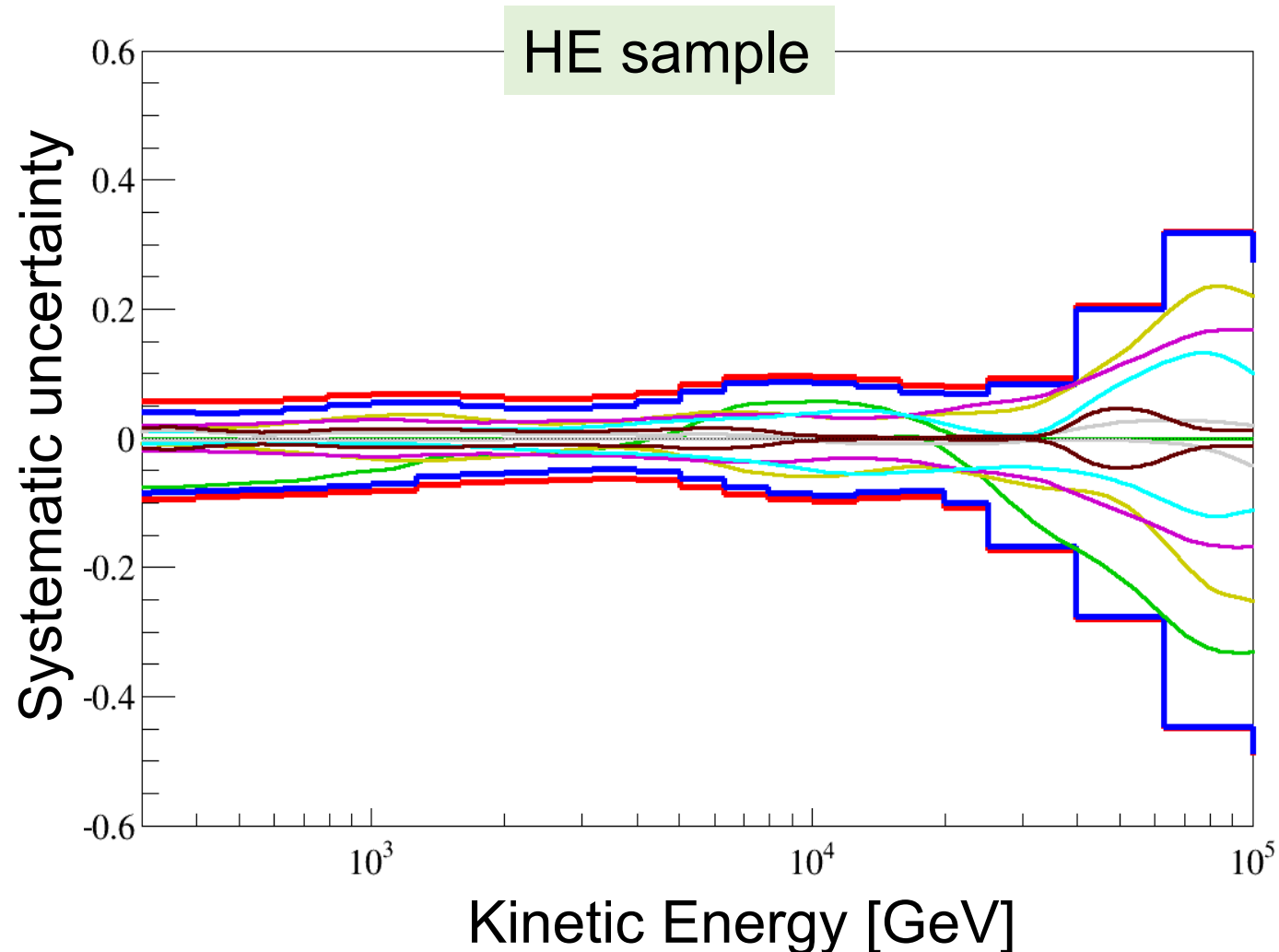
Observed/Unfolded energy spectrum



CALET's energy resolution for protons is 30-40%. Therefore, we apply Bayesian unfolding to calculate the flux.

1. Build response matrix between true and observed energy spectra using MC simulations.
2. Apply unfolding (RooUnfold) iteratively based on Bayes theorem with helium and electron background evaluation.

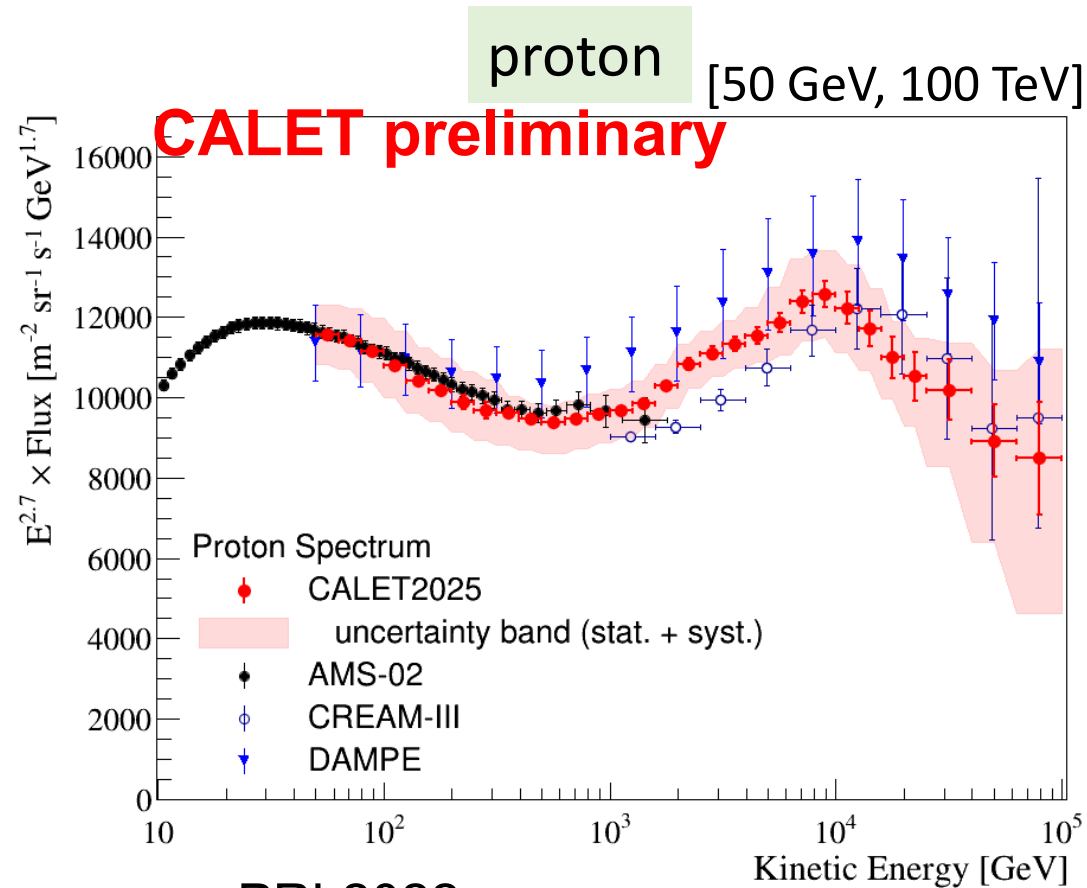
Systematic Uncertainty (proton)



- total uncertainty
- energy dependent uncertainty (sum)
- MC model dependence
- IMC Track consistency with TASC
- Shower start in IMC
- Charge identification cut
- Energy unfolding
- Beam test configuration

- Systematic uncertainty below 20 TeV is $< 10\%$.
- The uncertainty above 20 TeV comes from the MC model dependence and TASC Hit consistency, mainly.

Proton and Spectrum (kinetic energy)



1 more year data are newly analyzed
after ICRC2025

LE: same as PRL2022

HE: 3156 days of live time (Oct. 2015 – Dec. 2025)

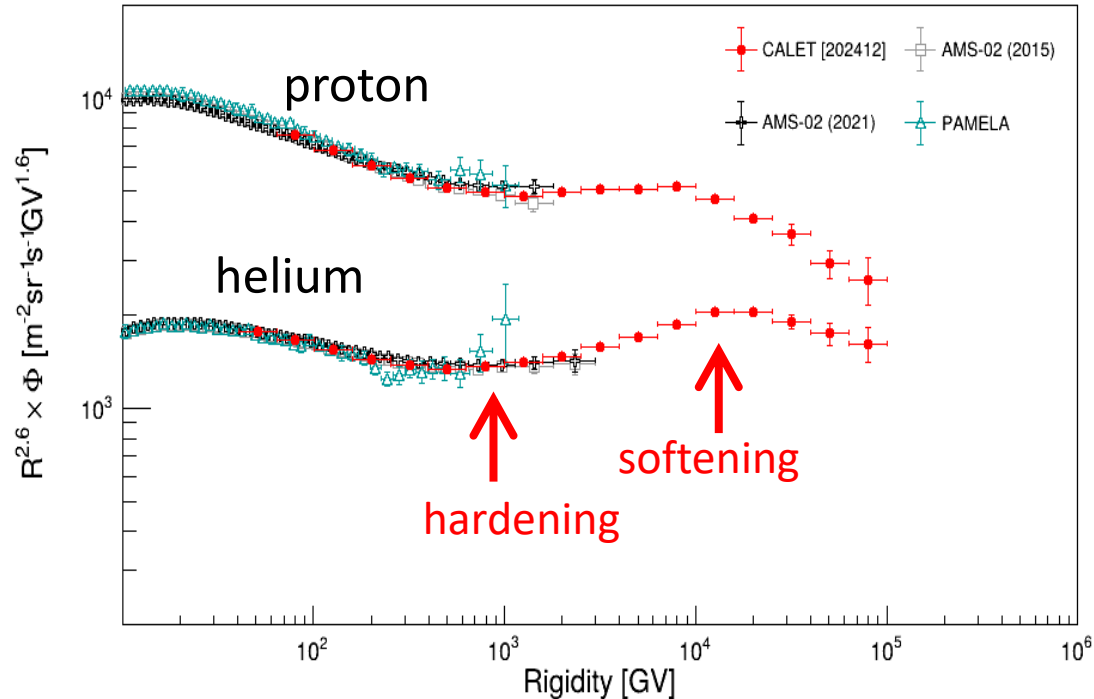
Proton and helium have same spectral signature such as hardening and softening.



Proton/He ratio (rigidity) 1/2

p and He fluxes (rigidity)

data: Oct. 2015 – Dec. 2025



If a simple leaky box model is assumed,

$$\frac{dN_p}{dR} \sim Q_{p \text{ source}}(R) T_{res}(R)$$

$$\frac{dN_{He}}{dR} \sim Q_{He \text{ source}}(R) T_{res}(R)$$

$$\Rightarrow \frac{dN_p}{dR} / \frac{dN_{He}}{dR} \sim Q_{p \text{ source}}(R) / Q_{He \text{ source}}(R)$$

$N_{p/He}$: p/He density

R : rigidity

$Q_{source}(R)$: source

$T_{res}(R)$: residual time

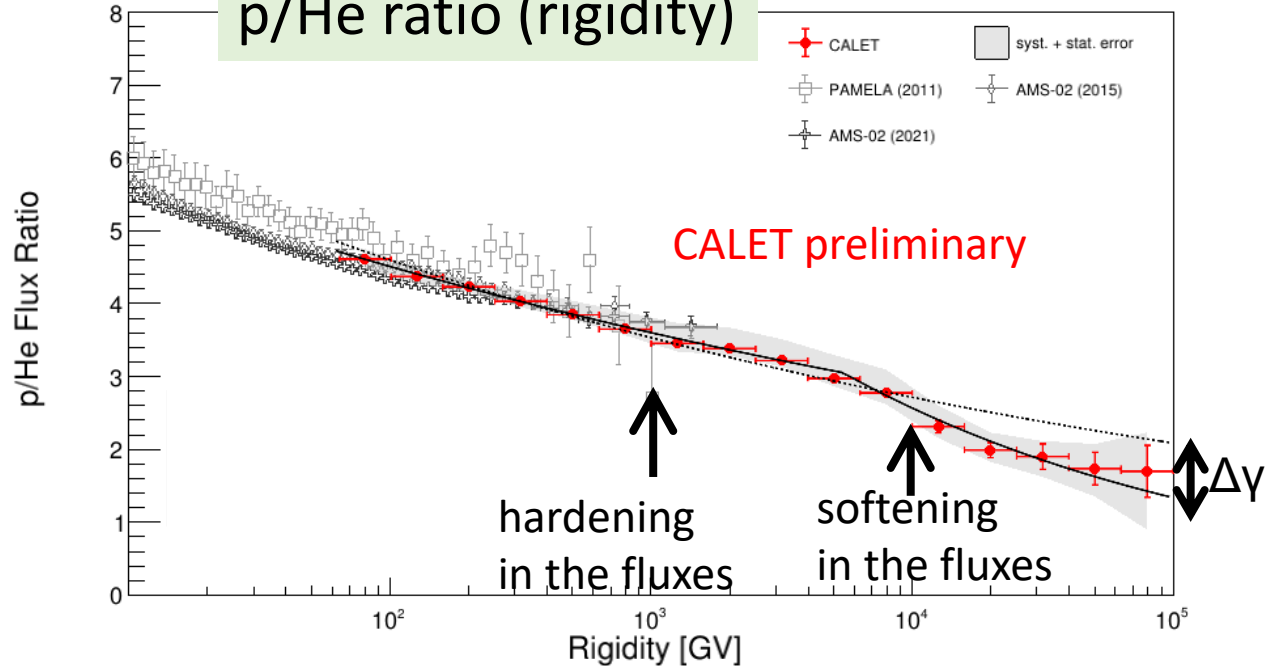
p/He ratio would provide useful probes of CR sources.

- Spectral hardening and softening in rigidity are roughly consistent between proton and helium.



Proton/He ratio (rigidity) 2/2

p/He ratio (rigidity)



$$\text{DPL: } \Phi(R) = \begin{cases} C \times \left(\frac{R}{1\text{GV}}\right)^{\gamma} & (R < R_0) \\ C \times \left(\frac{R}{1\text{GV}}\right)^{\gamma} \times \left(\frac{R}{R_0}\right)^{\Delta\gamma} & (R > R_0) \end{cases}$$

γ	-0.10 ± 0.01
R_0	$5.4 \pm 2.4 \text{ TV}$
$\Delta\gamma$	-0.19 ± 0.08

- p/He ratio from tens GV to 10 TV is well described by SPL ($\Phi(R) = A \times R^{\gamma}$, $\gamma = -0.10 \pm 0.01$).
- A spectral break is found around 10 TV, where both proton and helium spectra start softening.
- We performed three independent analyses and got the consistent result.

-> Different CR sources are needed to explain:

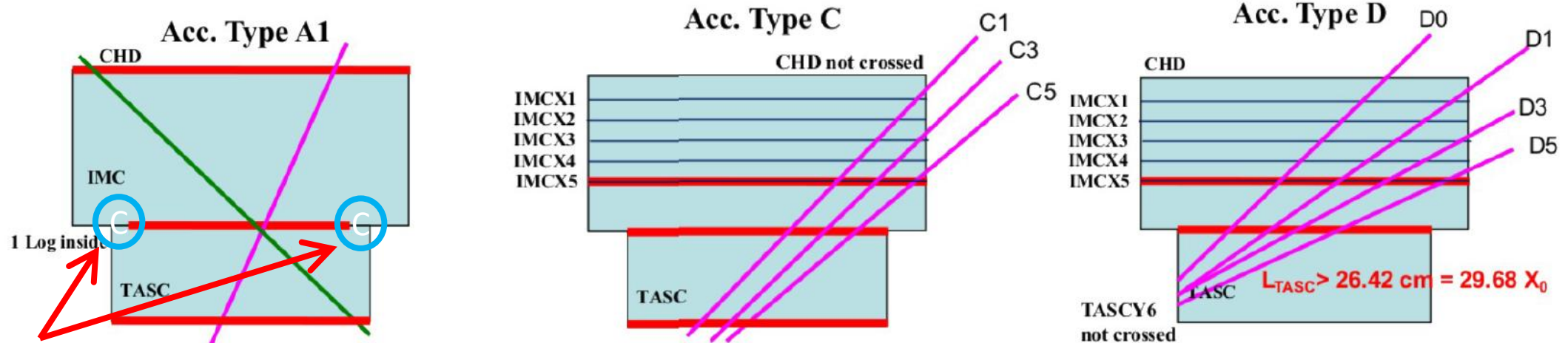
- the energy dependent (SPL up to 10 TV)
- the break around 10 TV (different break energy between proton and helium?)

Analysis improvement

We are trying to increase the efficiency and to reduce the statistics and systematic errors.

- Expand the geometrical acceptance. (See below. A1 is 20% larger and K4 which allow A1+C1+C3+C4+D0+D1+D3+D4 is 2.1 times larger.)
- Keep the same analysis method between proton and helium to reduce the systematic errors.
- We are trying to use BDT based analysis which keeps the same background rejection power with higher efficiency, especially improving the “off-acceptance cut”, “TASC hit consistency”,

ar



2cm margin in TASC is taken.

Current proton analysis have 2cm margin in the bottom side as well.



Summary

- CALET data taking is stably running without any serious problem more than 10 years.
- We have updated the proton/helium analysis using data until Dec. 2025. Proton and helium energy spectrum have a similar shape. Both have hardening and softening. Starting rigidity of spectral hardening are consistent between proton and helium.
- Proton/He ratio tend to be not independent of rigidity. The ratio from tens GV to 10 TV is well described by SPL, while that from 60 GV to 100 TV is not well described by SPL. Around the breaking rigidity (~ 10 TV), spectral softening of protons and helium starts at slightly different rigidity. These results may indicate that different CR sources affect the proton and helium spectra.

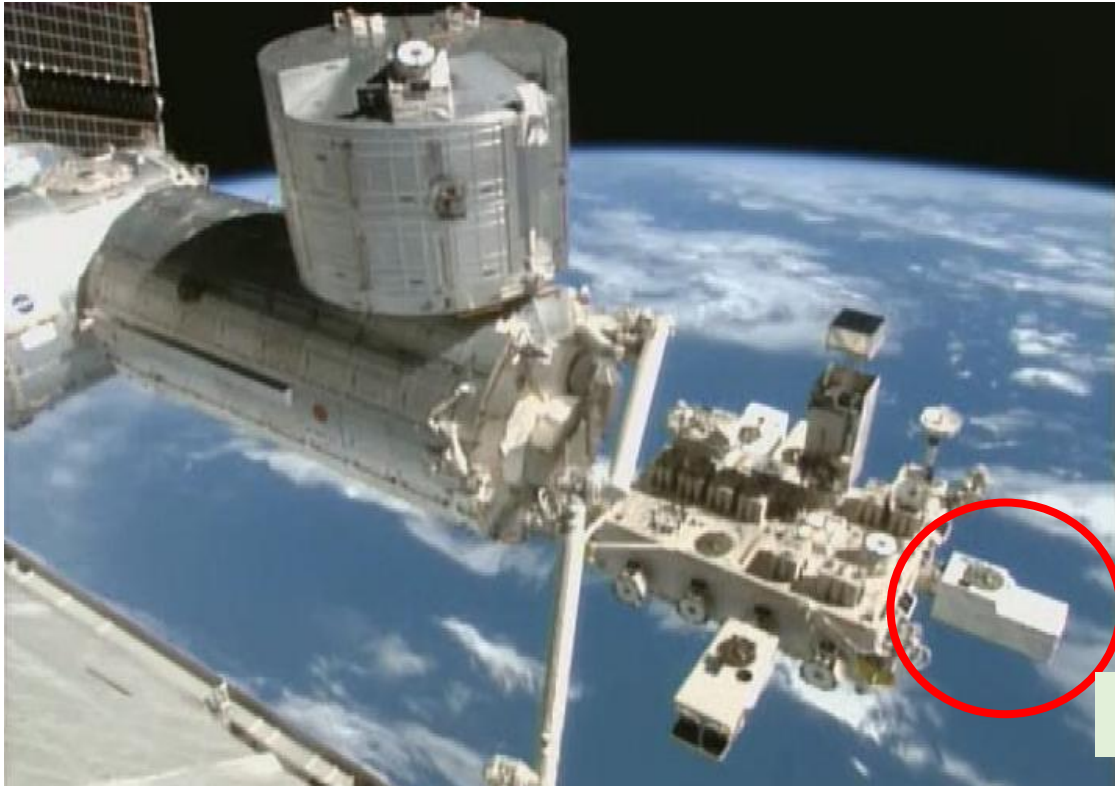


Back up

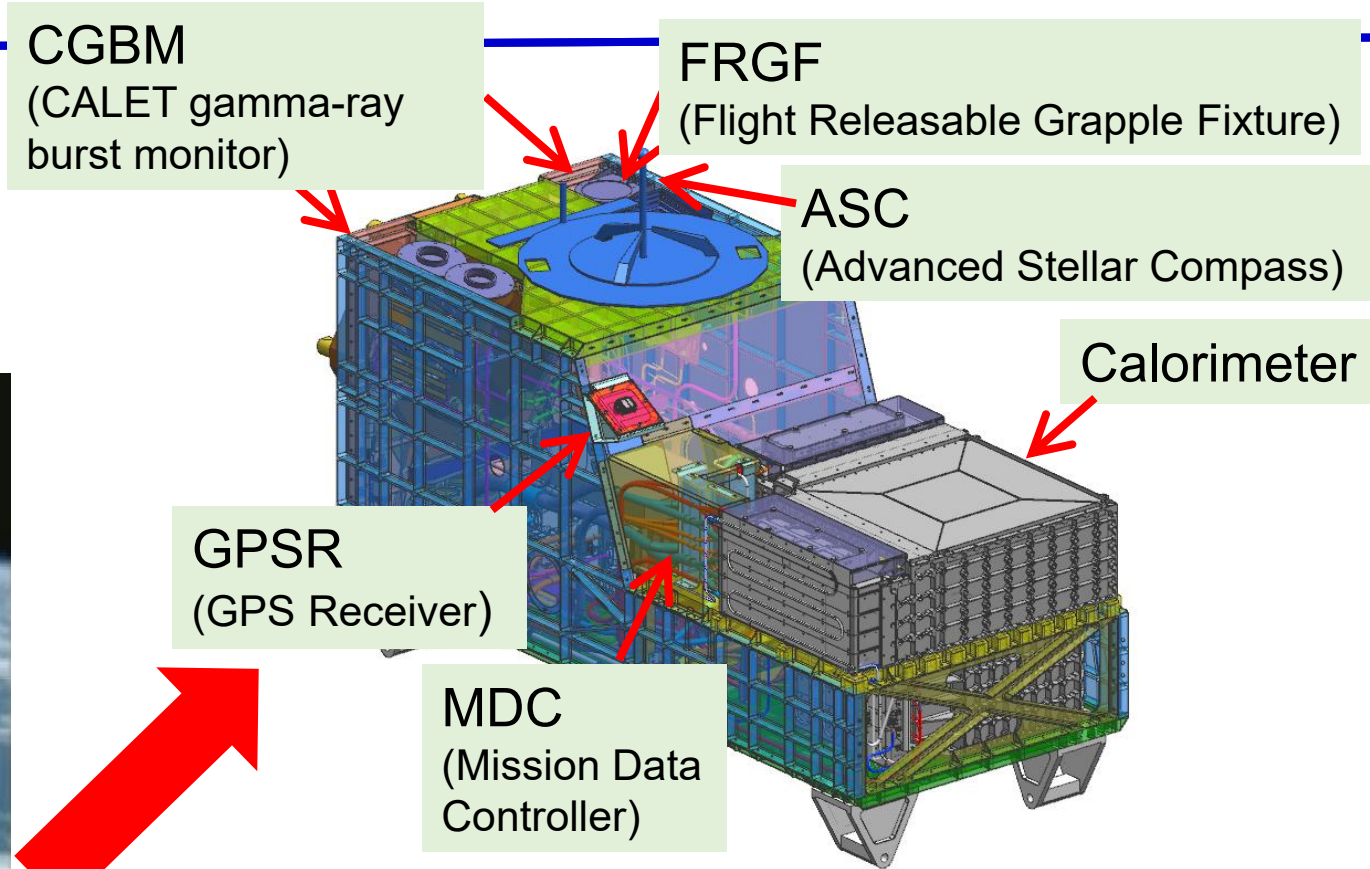


CALET Project

Aug. 2015: launched and installed on the ISS
Oct. 2015: scientific data taking began
Data taking is stably running up to now.
We plan to operate through the end of 2030.



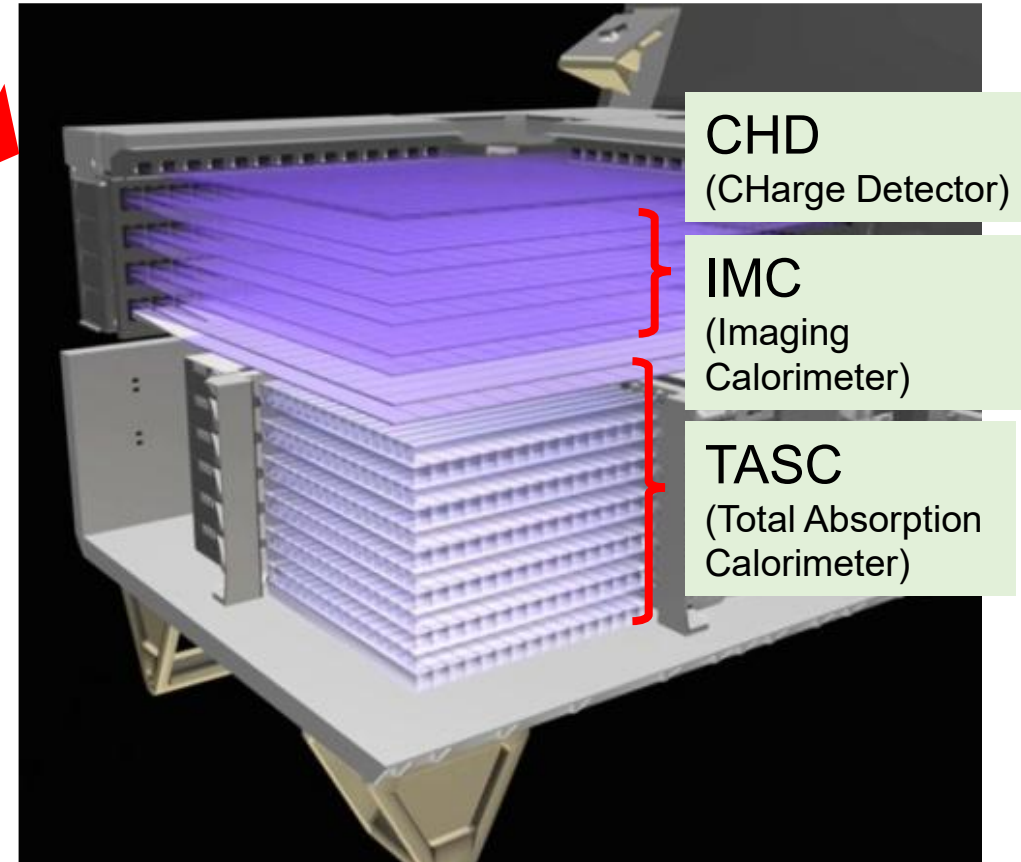
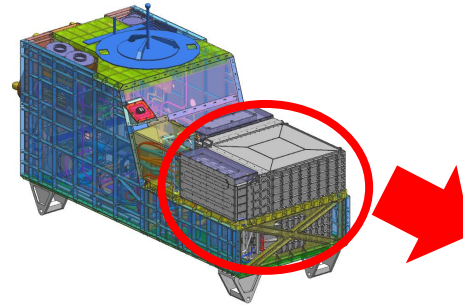
International Space Station (ISS)



JEM-EF/Port #9

Mass	612.8kg
Power	507W (max)
Telemetry	600kbps (6.5GB/day)

CALET Detector



CHD
(CHarge Detector)

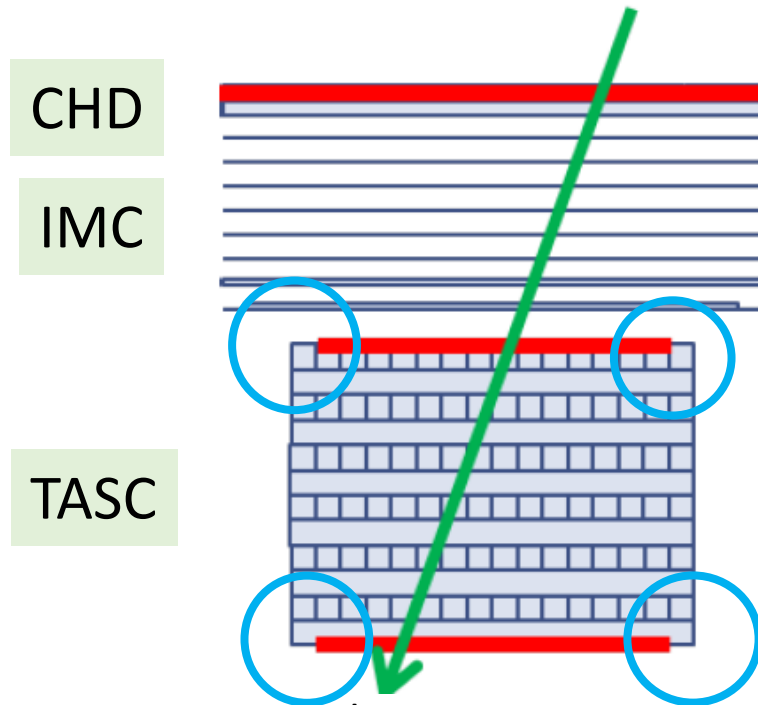
IMC
(Imaging
Calorimeter)

TASC
(Total Absorption
Calorimeter)

In total $30X_0$ thickness and 1.2λ ,
($27X_0$ in TASC + $3X_0$ in IMC)

	Material/Sensor	Function
CHD	Plastic scintillator + PMT 28 paddles (=14x2layers(x,y)) (paddle size: 32x10x450mm)	Charge ID
IMC	Scifi./W + MAPMT (64anode) 7168 Scifi. (=448x16layers(x,y)) +7 W layers (Scifi. size: 1x1x448mm)	Tracking, Charge ID
TASC	PWO scintillator + APD/PD or PMT 192 logs (=16x12layers(x,y)) (Log size: 19x20x326mm)	Energy, Particle (e/p) ID

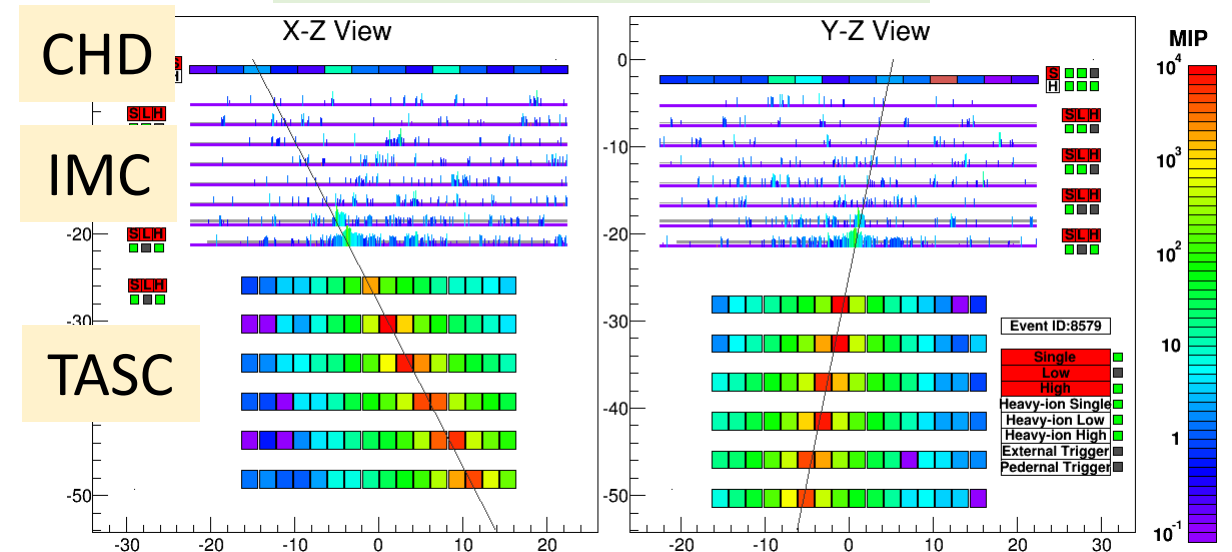
Geometrical acceptance and event example



2cm margin in TASC is taken.

- The reconstructed track is required to cross the CHD and TASC from top to bottom.

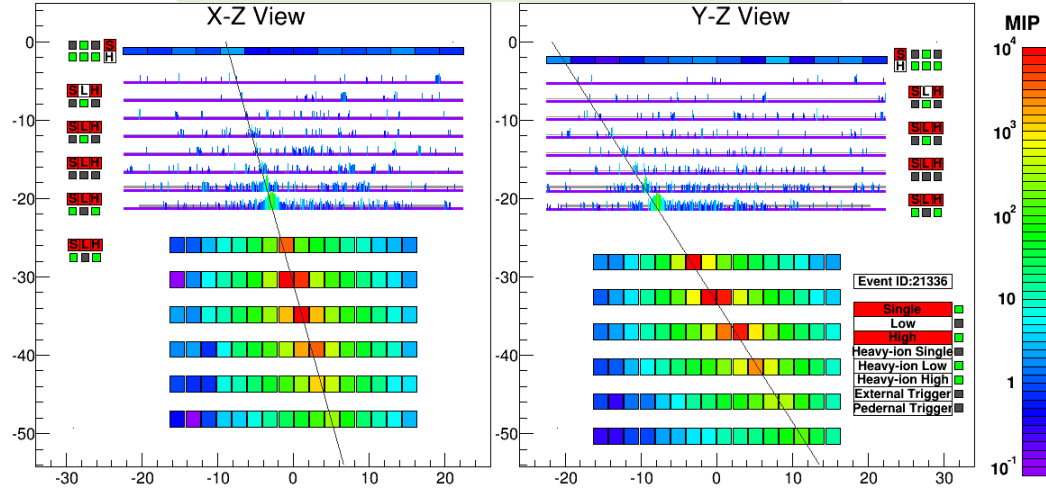
Proton, $\Delta E = 2.89$ TeV



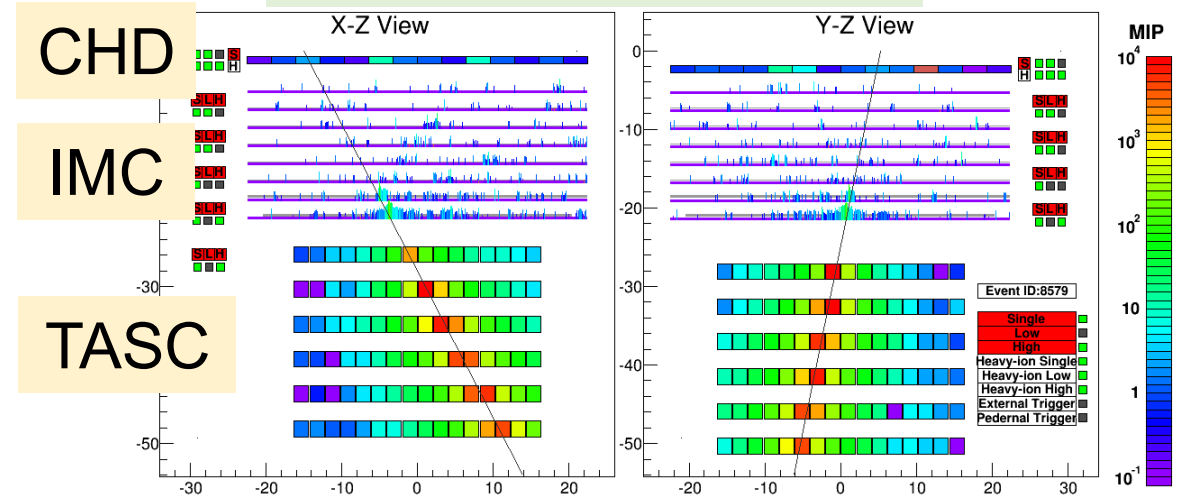


Event Examples

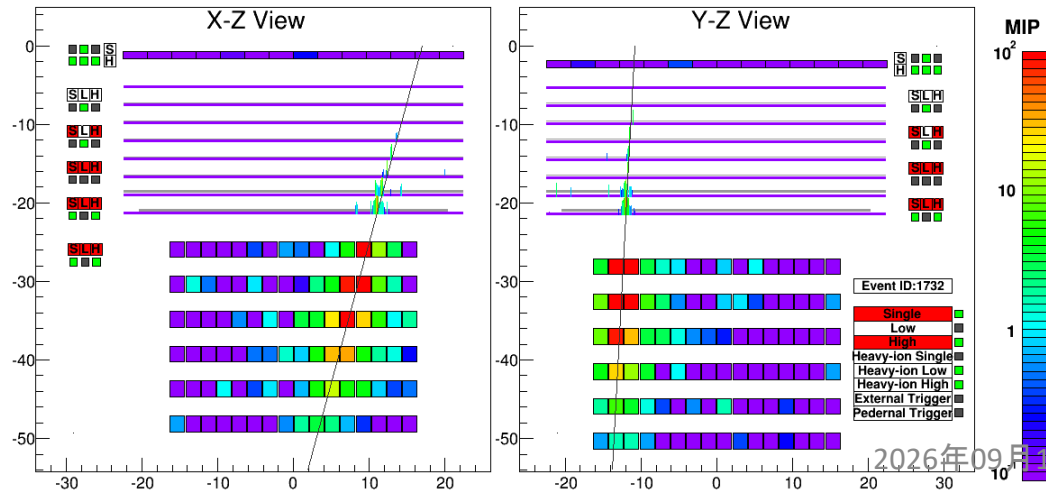
Electron, $E=3.05$ TeV



Proton, $\Delta E=2.89$ TeV



Gamma-ray, $E=44.3$ GeV



Fe, $\Delta E=9.3$ TeV

