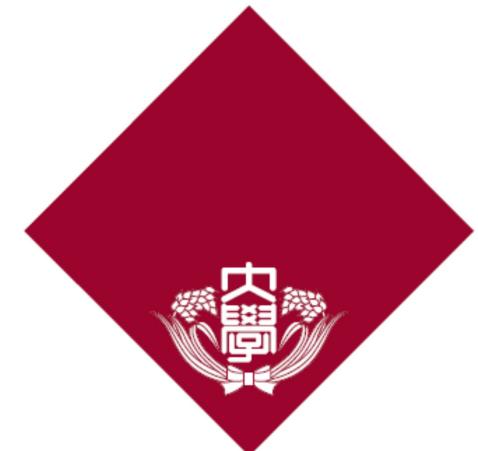
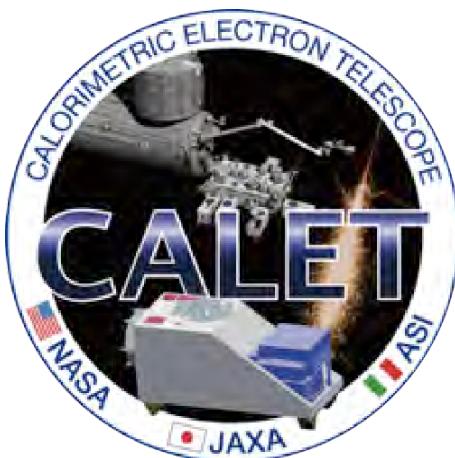


CALET5年間の観測による 炭素・酸素・鉄のエネルギースペクトルと B/C比の観測結果

早大理工総研, 東大宇宙線^B, 芝工大シ工^C,
弘前大理工^D Siena Univ.^D, INFN Pisa^E

赤池陽水, 鳥居祥二, 小林兼好,
浅岡陽一^A, 笠原克昌^B, 市村雅一^C,
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Caterina Checcia^E, Francesco Stolze^{D,E}
他 CALET チーム



CALET5年間の観測による 炭素・酸素・鉄のエネルギースペクトルと B/C比の観測結果

Publications:

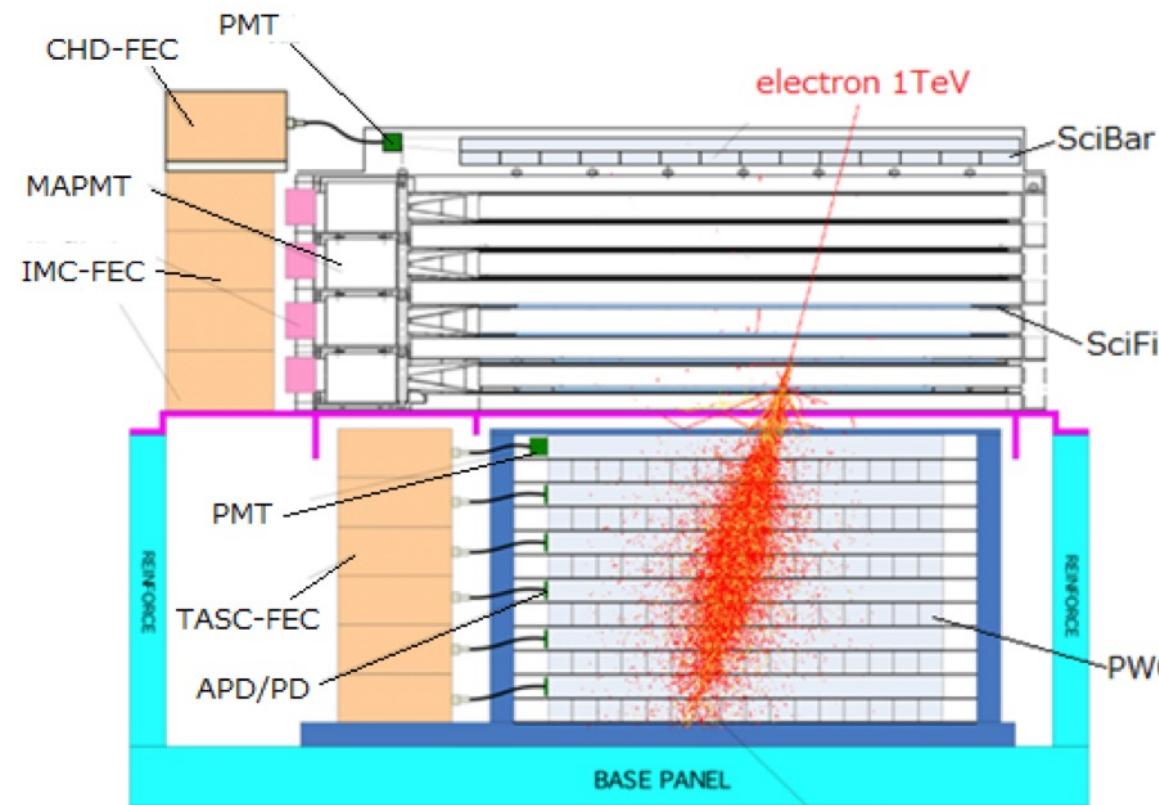
- Carbon and oxygen
 - Phys. Rev. Lett. 125, 251102 (2020)
 - PoS (ICRC2021) 093
- Iron
 - Phys. Rev. Lett. 126, 241101 (2021)
 - PoS (ICRC2021) 109
- Boron and B/C ratio
 - PoS (ICRC2021) 112



Instrument of CALET

A 30-radiation length deep calorimeter designed to detect electrons and gammas to 20 TeV and cosmic rays up to 1 PeV

Since the start of operation on the ISS in October 2015,
CALET has been accumulating scientific data without any major interruption



CHD: Charge Detector

Charge measurements ($Z=1-40$)

- Plastic scintillator paddles $14 \times (X, Y)$
Unit size: $32\text{mm} \times 10\text{ mm} \times 450\text{ mm}$
 $\Delta Z/Z = 0.15$ for C, 0.30 for Fe

IMC: Imaging Calorimeter

Arrival direction, Particle ID

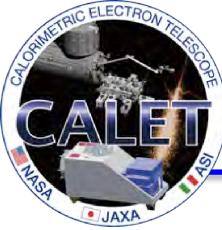
- Scintillating fiber belts 448×16 layers
Unit size: $1\text{ mm}^2 \times 448\text{ mm}$
- Tungsten plates 7 layers
 $3 X_0 (=0.2 X_0 \times 5 + 1.0 X_0 \times 2)$

ΔX at CHD = $200\mu\text{m}$, $\Delta Z/Z = 0.20$ for C

TASC: Total Absorption Calorimeter

Energy measurement, Particle ID

- PWO logs 16×12 layers
Unit size: $19\text{ mm} \times 20\text{ mm} \times 326\text{ mm}$
 $27 X_0$ for electrons
1.2 interaction length for protons
Dynamic range ; $1 - 10^6$ MIP (1GeV – 1PeV)



Data analysis for nuclei

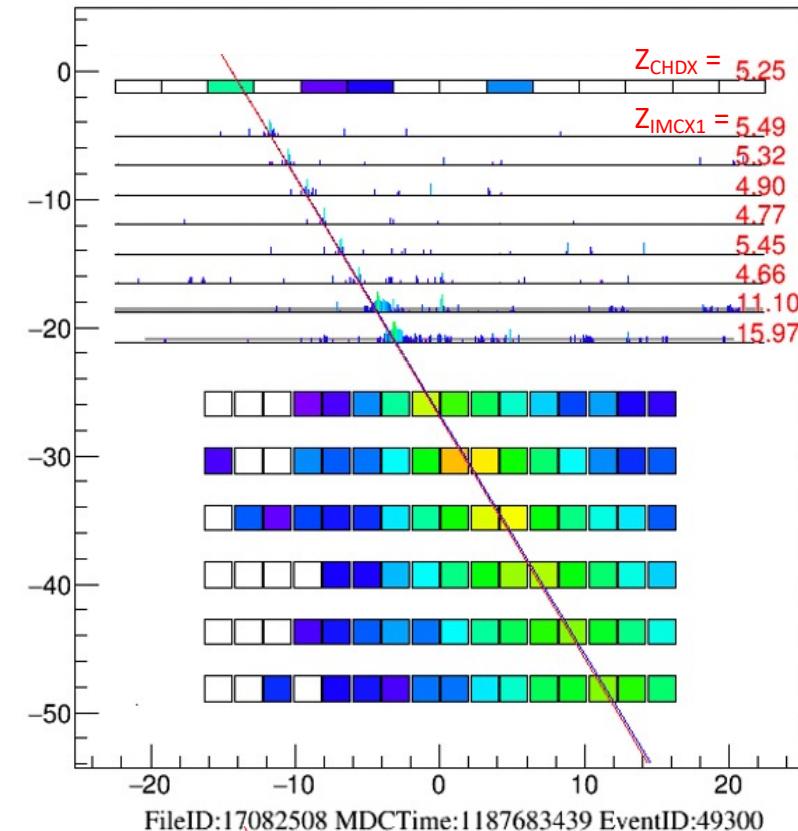
Analyzed Flight Data

1,815 days (Oct. 13, 2015 – Sep. 30, 2020)
 $T_{\text{live}} = 3.69 \times 10^4$ hours

Analysis procedure

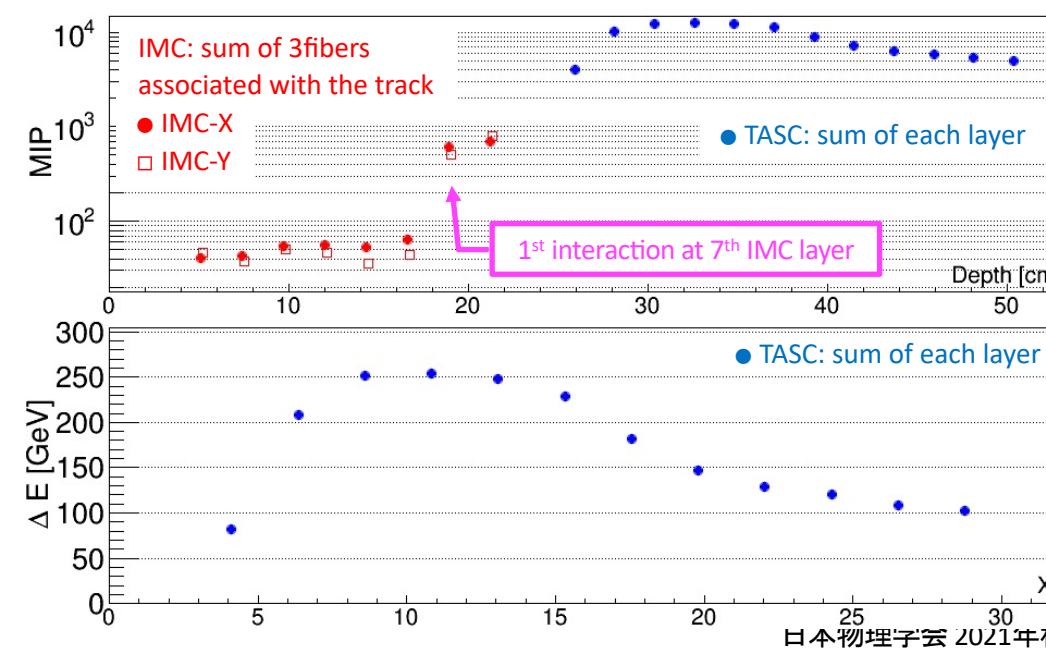
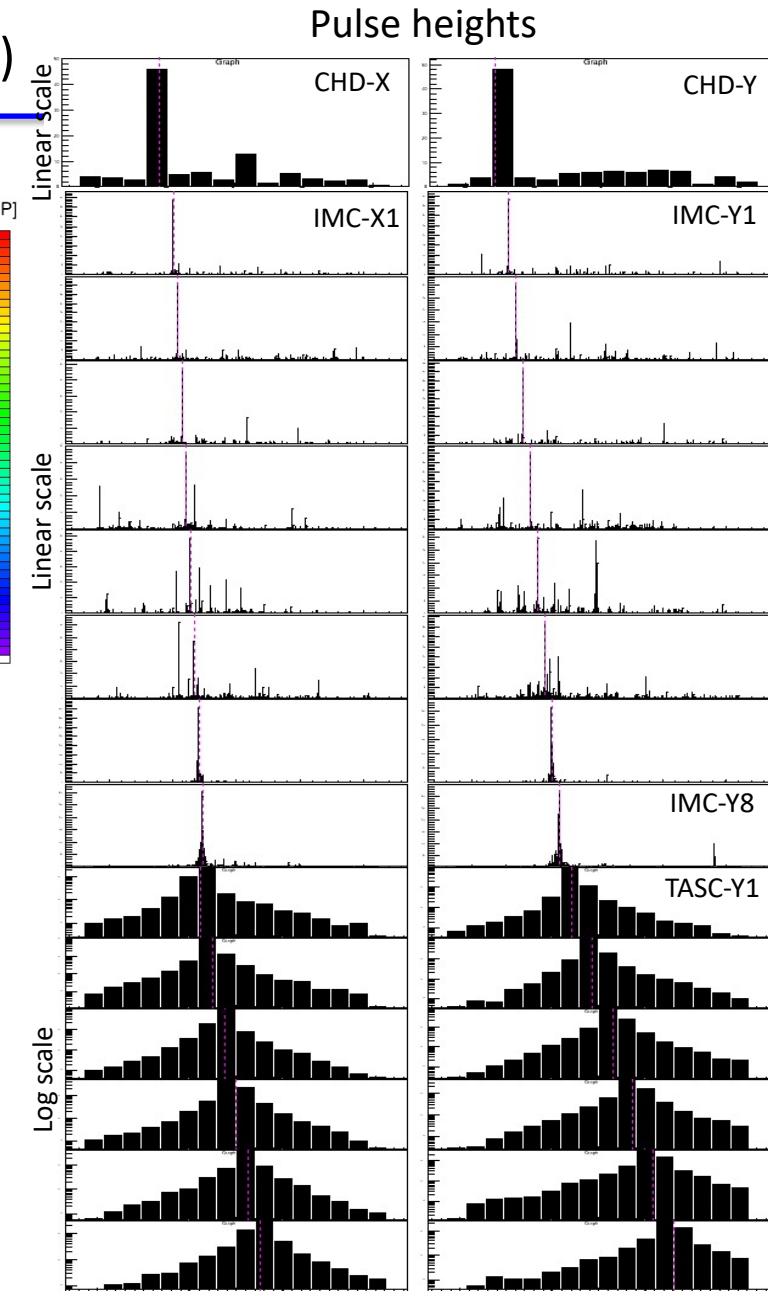
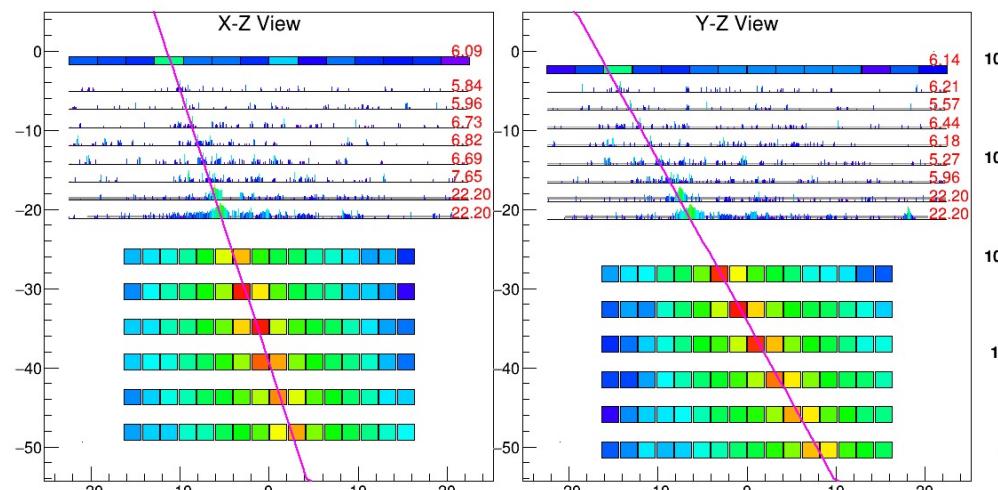
- HE + offline shower trigger
50MIP in IMC-X/Y78, 100MIP in TASC-X1
- Tracking with IMC
- Field of view cut
Remove shielded region by ISS structures
- Acceptance cut
CHD, TASC top and bottom layers
- Charge identification
Charge consistency among CHD and IMC layer
Track width selection
- Estimate efficiency and background
- Apply energy unfolding
- Calculate flux and the ratio

An example of Boron candidate (X-Z view)
 $E_{\text{TASC}} = 475.8$ GeV



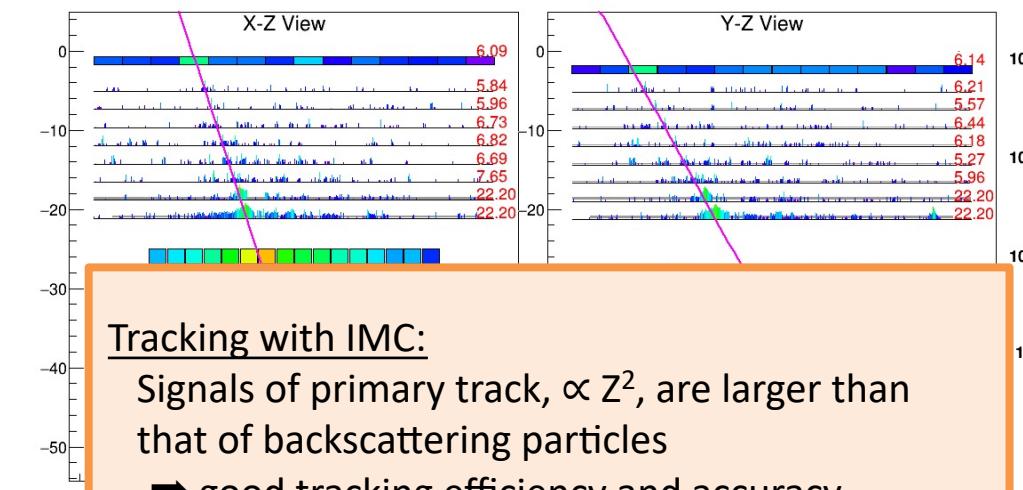


Event display: Carbon ($\Delta E_{\text{TASC}}=2.06\text{TeV}$)





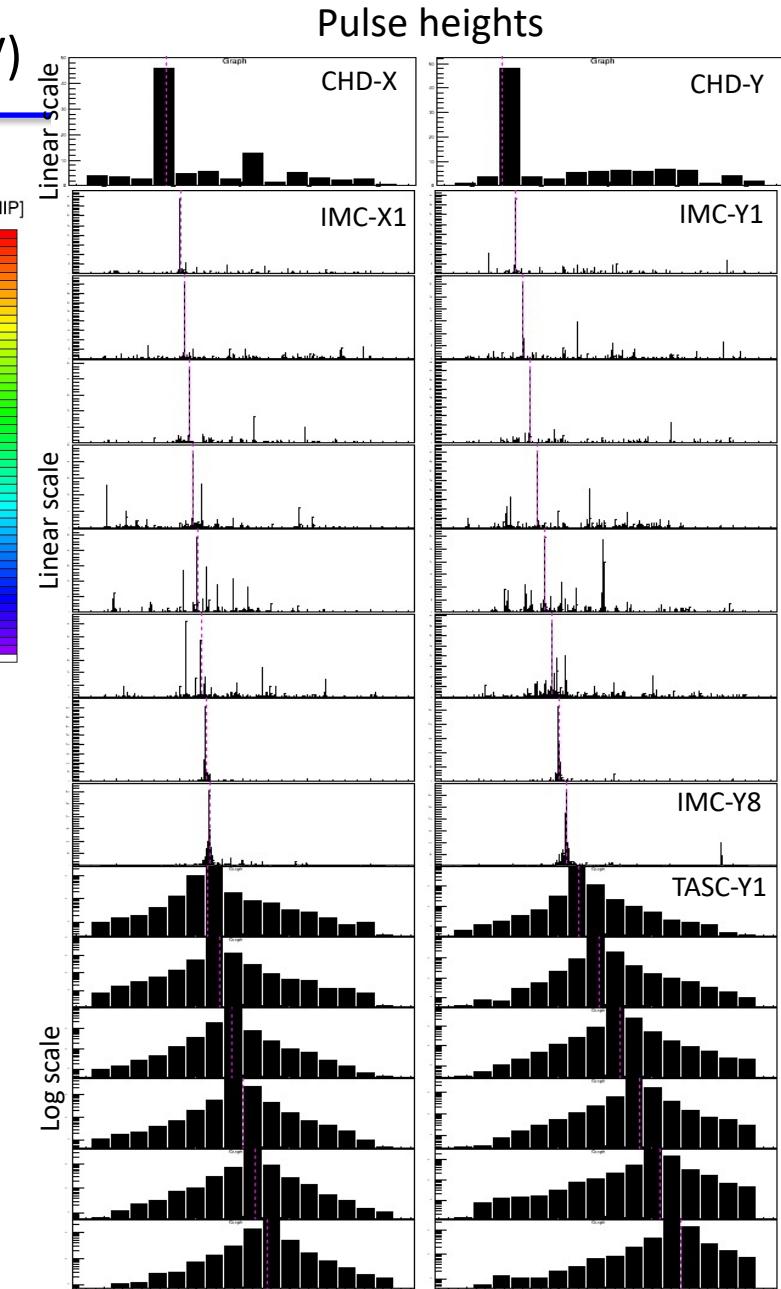
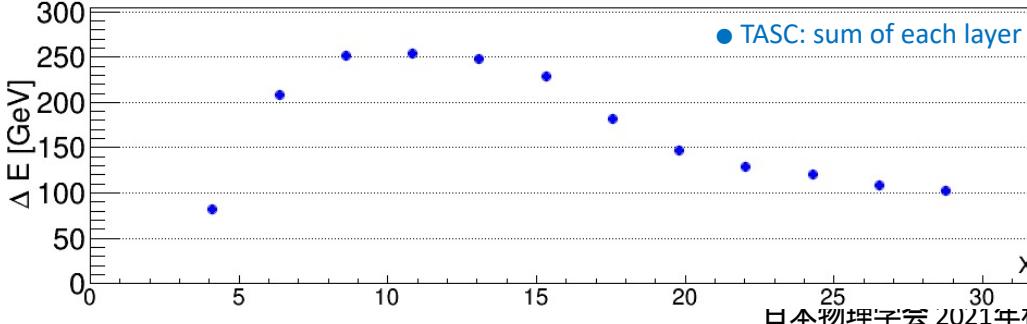
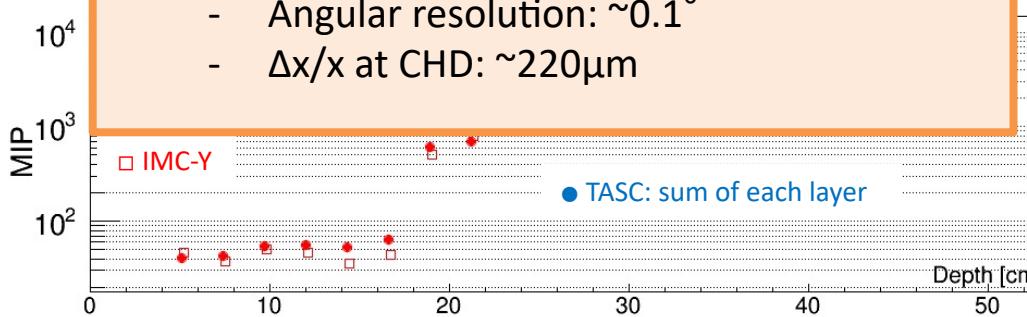
Event display: Carbon ($\Delta E_{\text{TASC}}=2.06\text{TeV}$)



Tracking with IMC:

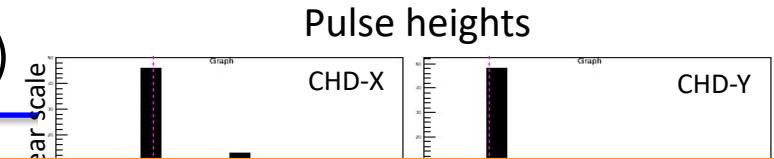
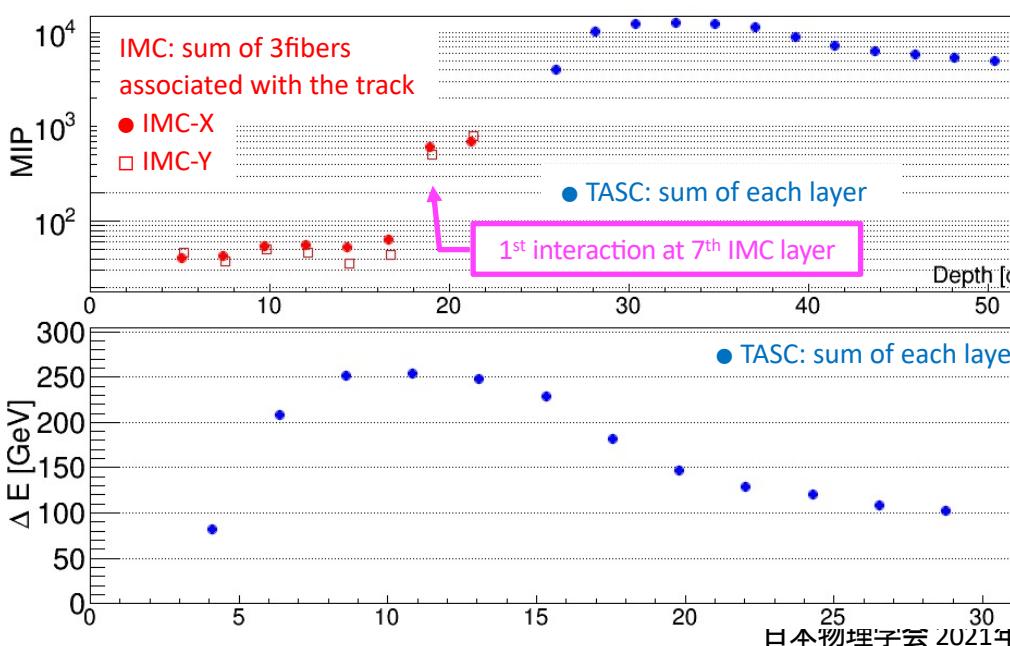
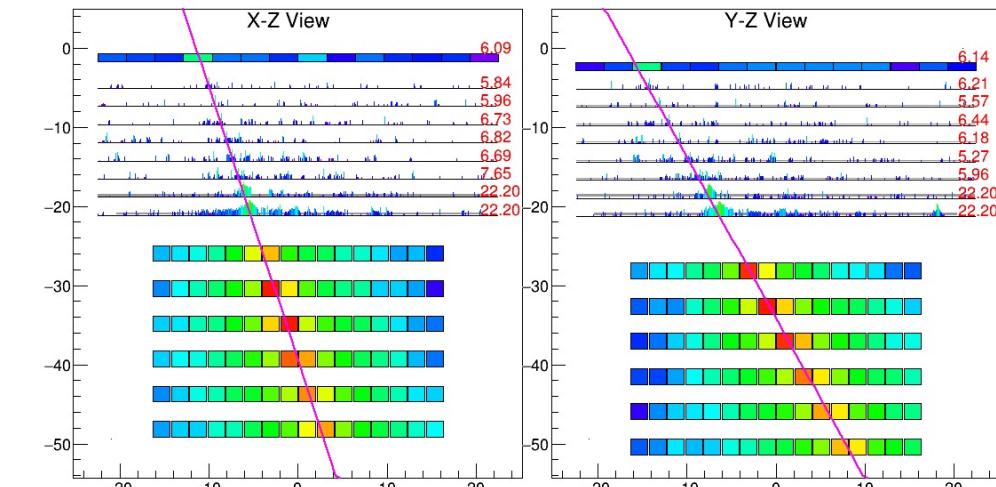
Signals of primary track, $\propto Z^2$, are larger than that of backscattering particles

- good tracking efficiency and accuracy
 - Angular resolution: $\sim 0.1^\circ$
 - $\Delta x/x$ at CHD: $\sim 220\mu\text{m}$



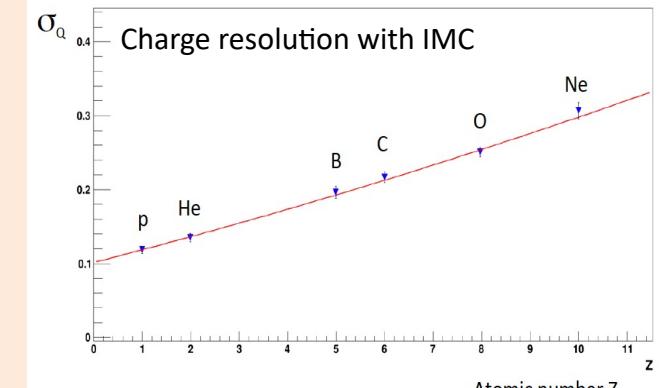
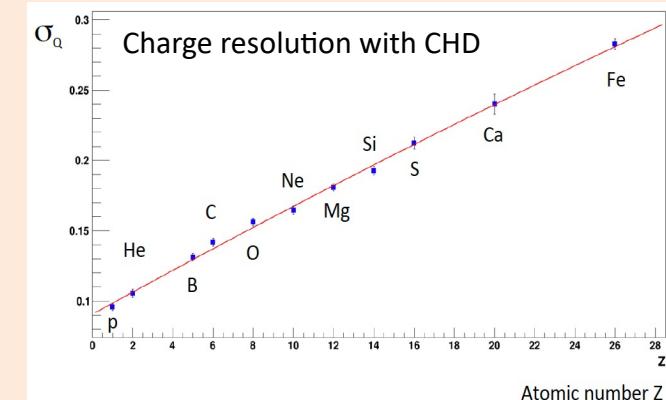


Event display: Carbon ($\Delta E_{TASC}=2.06\text{TeV}$)



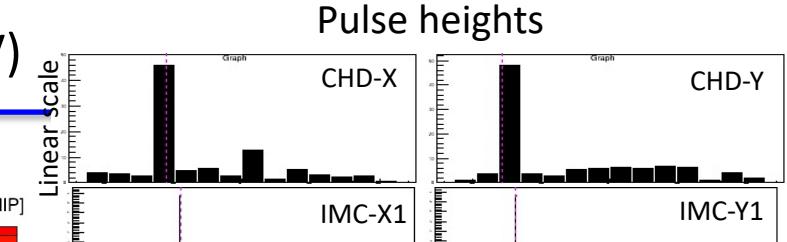
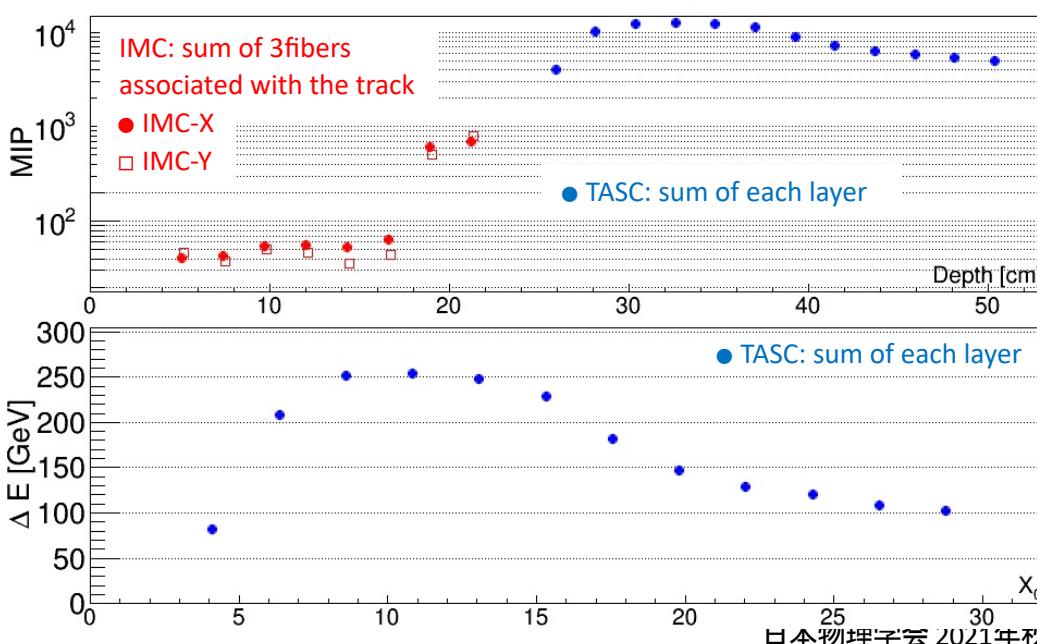
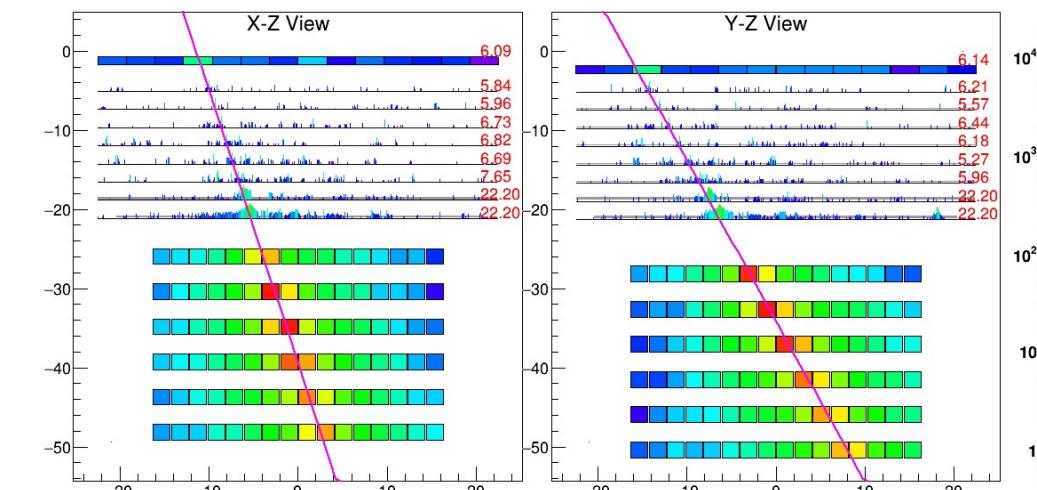
Charge measurement with CHD and IMC:
dE/dx measurement with CHD and IMC
IMC can identify the 1st interaction points

- $\Delta Z/Z$ (CHD): 0.15 for B, C, O, 0.30 for Fe
- $\Delta Z/Z$ (IMC): 0.20 for B, C, O



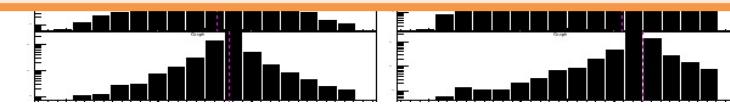
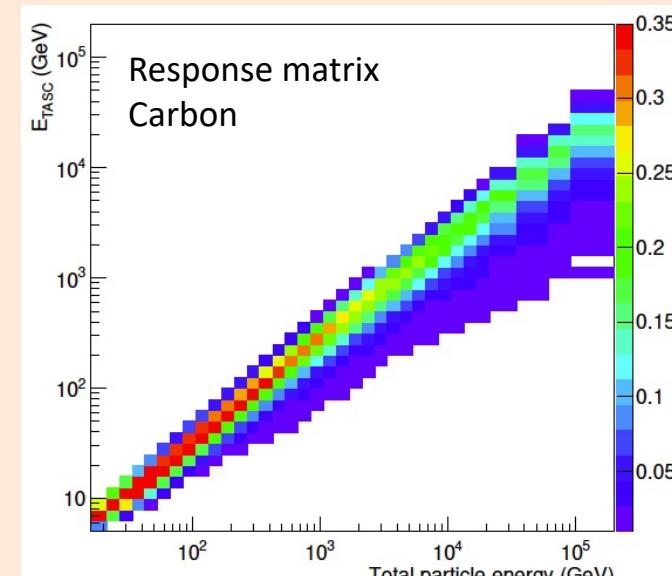


Event display: Carbon ($\Delta E_{\text{TASC}}=2.06\text{TeV}$)



Energy measurement with TASC:

Calorimeter thickness: $1.3\lambda_l$ ($30X_0$)
- Energy resolution: $\sim 30\%$
→ Energy unfolding using MC simulation
Accuracy of the MC is tested
by 150 GeV/n ion beam at CERN-SPS





Flux measurement

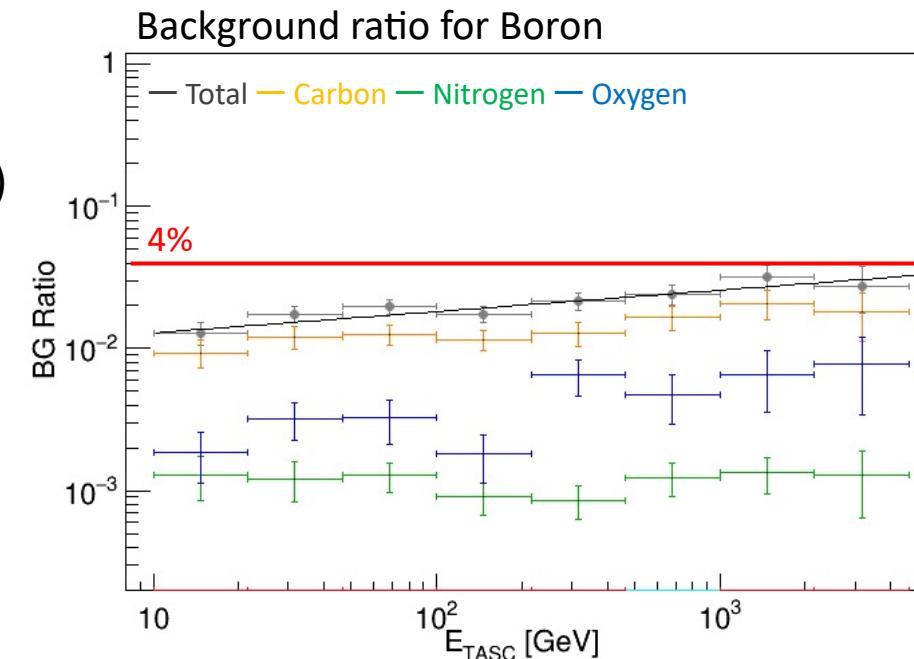
$$\Phi(E) = \frac{N(E)}{\Delta E \varepsilon(E) S \Omega T}$$

$$N(E) = \frac{U}{\text{The number of observed events}} [N_{\text{obs}}(E_{\text{TASC}}) - N_{\text{bg}}(E_{\text{TASC}})]$$

Energy unfolding

The number of background events

- $\varepsilon(E)$: Efficiency
 - 15-35% for B, C and O
 - 20-40% for Fe
- $S\Omega$: Geometrical factor (510 cm²sr)
- T: livetime
 - 1,480 days for C and O
 - 1,613 days for Fe
 - 1,815 days for B



Background ratios of C, O and Fe are ~1 %

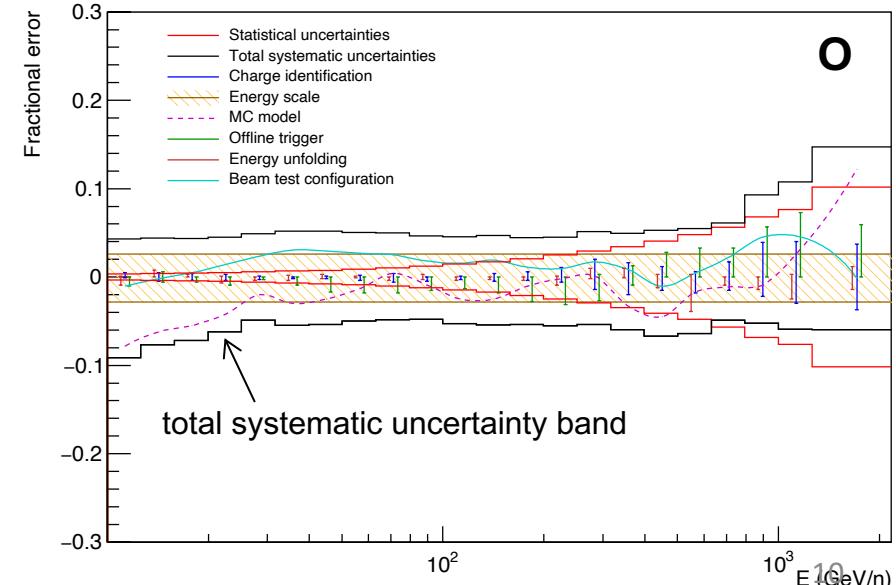
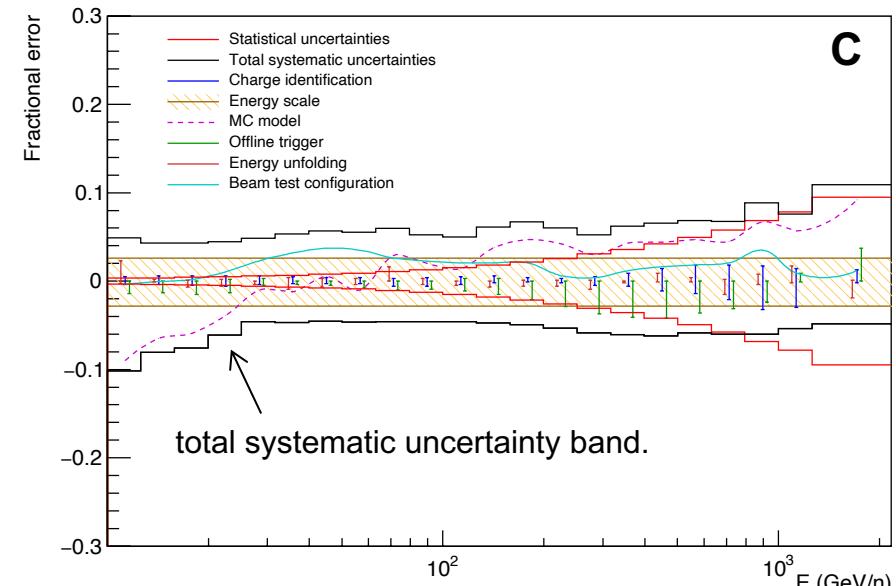


Systematic uncertainties

We check the stability of the spectrum by varying the analysis cuts and w/ different MC simulations for efficiencies and unfolding.

Main sources of systematics uncertainties:

- **Normalization:**
 - Live time
 - Long-term stability
 - Energy scale
- **Energy dependent:**
 - Tracking
 - Charge ID
 - Trigger
 - Unfolding
 - Beam test configuration
 - MC model (EPICS, FLUKA , GEANT4)





Systematic uncertainties

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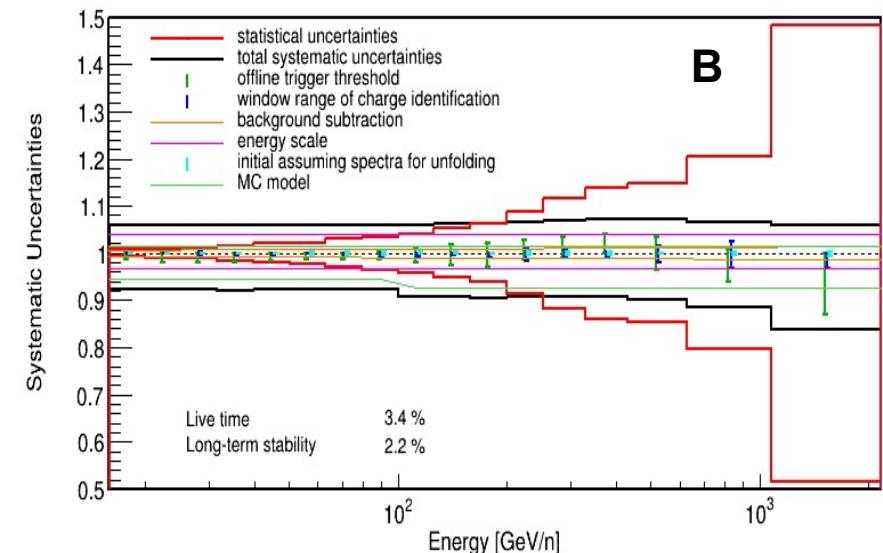
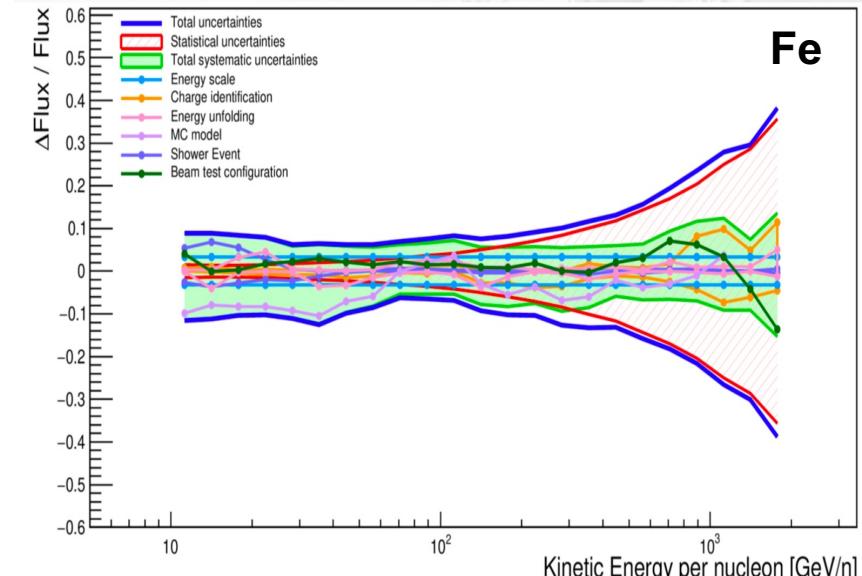
Main sources of systematics uncertainties:

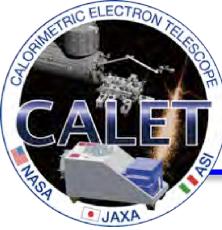
➤ **Normalization:**

- Live time
- Long-term stability
- Energy scale

➤ **Energy dependent:**

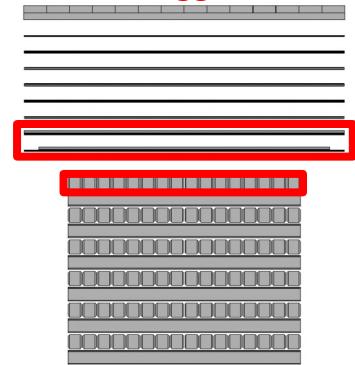
- Tracking
- Charge ID
- Trigger
- Unfolding
- Beam test configuration
- MC model (EPICS, FLUKA , GEANT4)





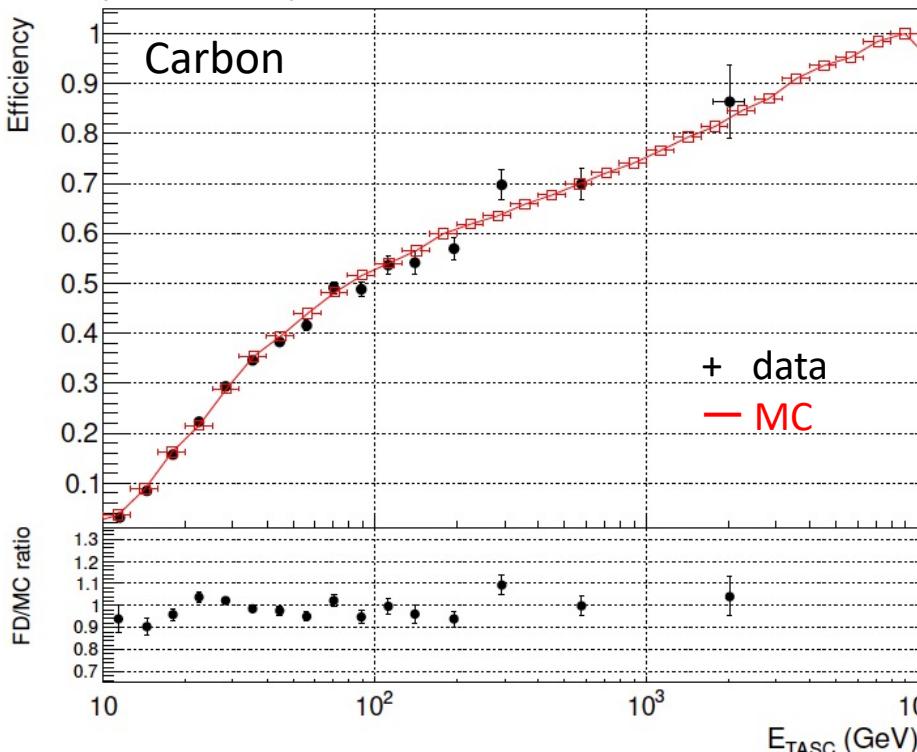
Trigger efficiency

Onboard trigger sources

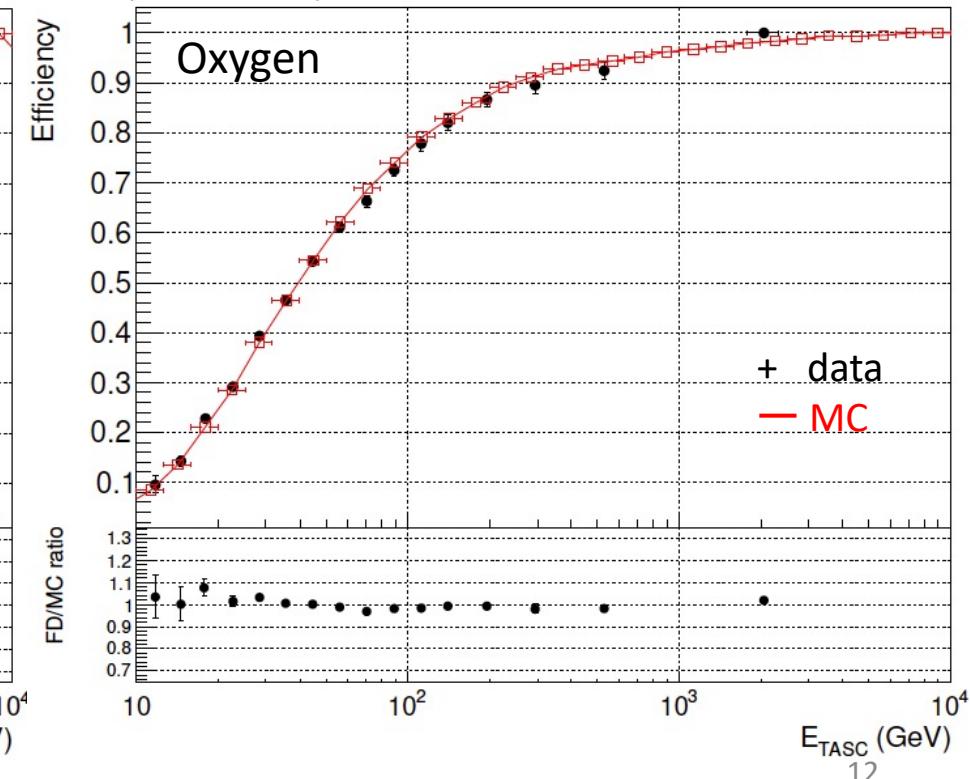


- High-Energy Trigger (HET), which is the primary CALET mission trigger, is the coincidence of signals in last two IMC layers and top TASC layer
- Low-Energy Trigger (LET) is the same trigger logic as HET, but lower threshold allowing to trigger also penetrating particles of C and O
- HET is modelled in simulation: good agreement between MC and flight data

(HET&LET) / LET

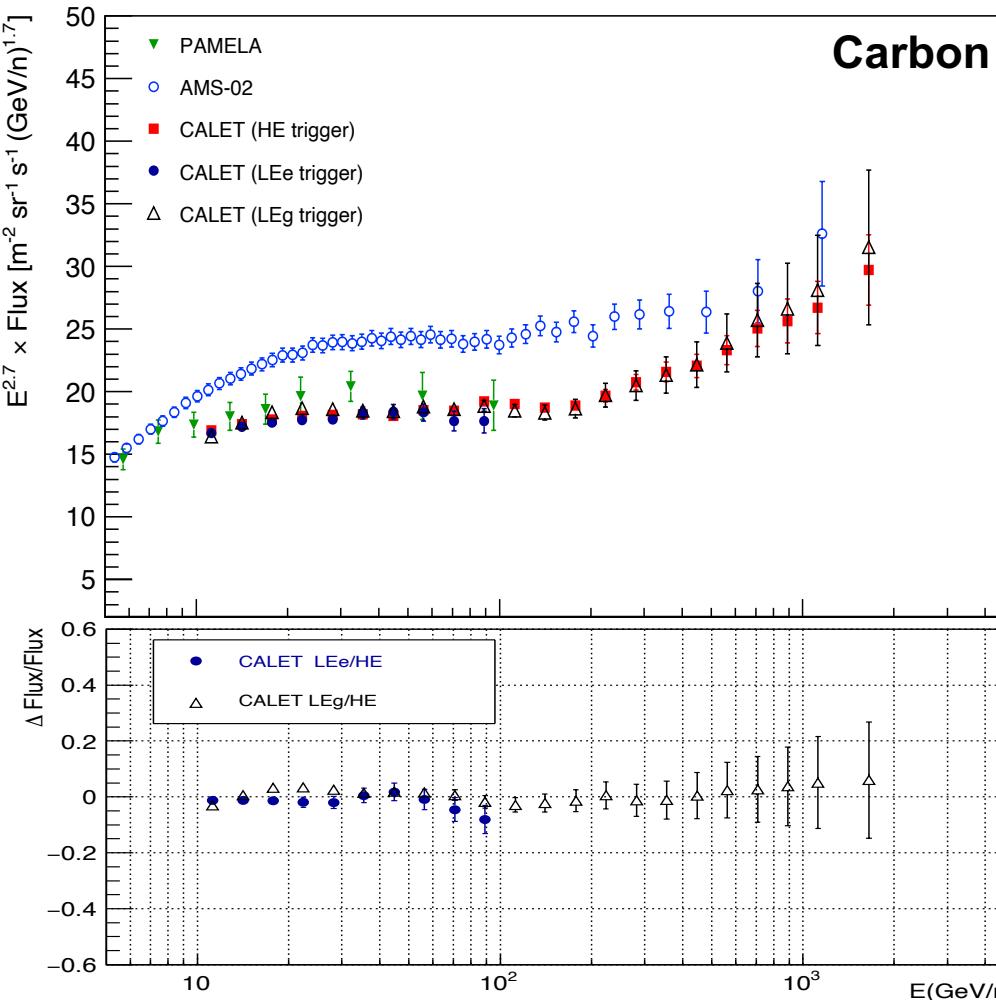


(HET&LET) / LET





High vs. low energy triggers



Carbon

High-energy (HE) trigger: $E > 10\text{GeV}$

- Coincidence of last two pairs of IMC layers (50 MIP thr.) & top TASC layer (100 MIP thr.)

Low-energy gamma (LEg) trigger: $E > 1\text{GeV}$

- Coincidence of last two pairs of IMC layers (5 MIP thr.) & top TASC layer (10 MIP thr.)
- livetime ~10% of HE livetime

Low-energy electron (LEe) trigger: $E > 1\text{GeV}$

- same as LEg with additional coincidence of CHD & upper IMC layers (0.3 MIP thr.)
- operated at a high geomagnetic latitude
- livetime ~2% of HE livetime

The resultant fluxes using data from the different trigger modes show consistent normalization and spectral shapes.



TASC energy scale

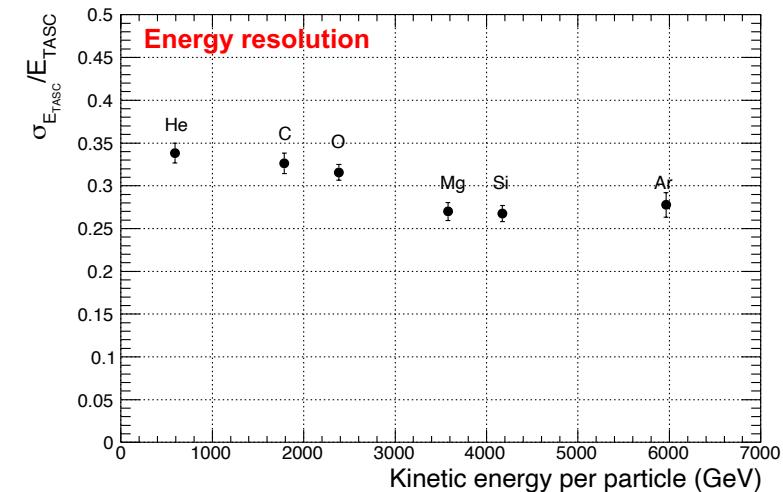
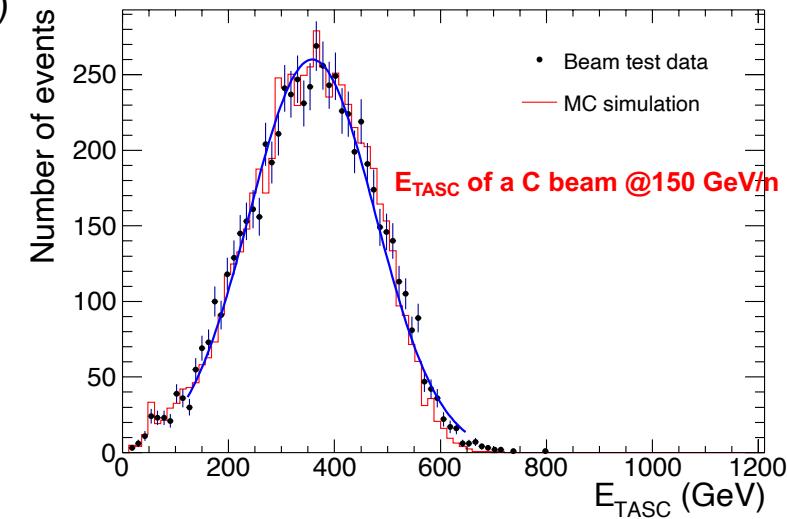
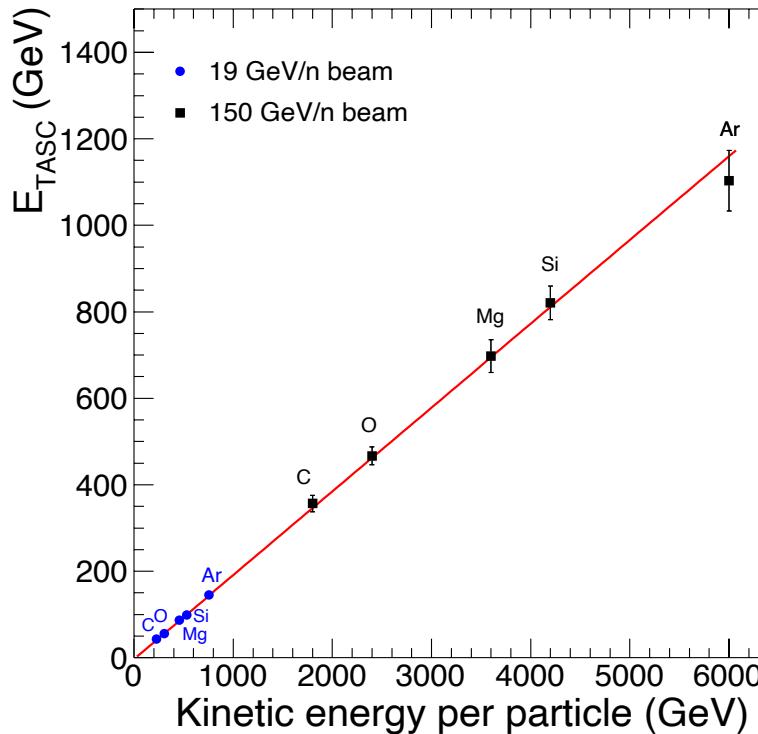
Beam test calibration at CERN-SPS with ion fragments beam ($Z/A=2$) at 13, 19 and 150 GeV/n.

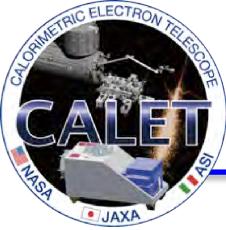
Good linearity up to max available beam energy (~ 6 TeV)

Fraction of particle energy released in TASC is $\sim 20\%$

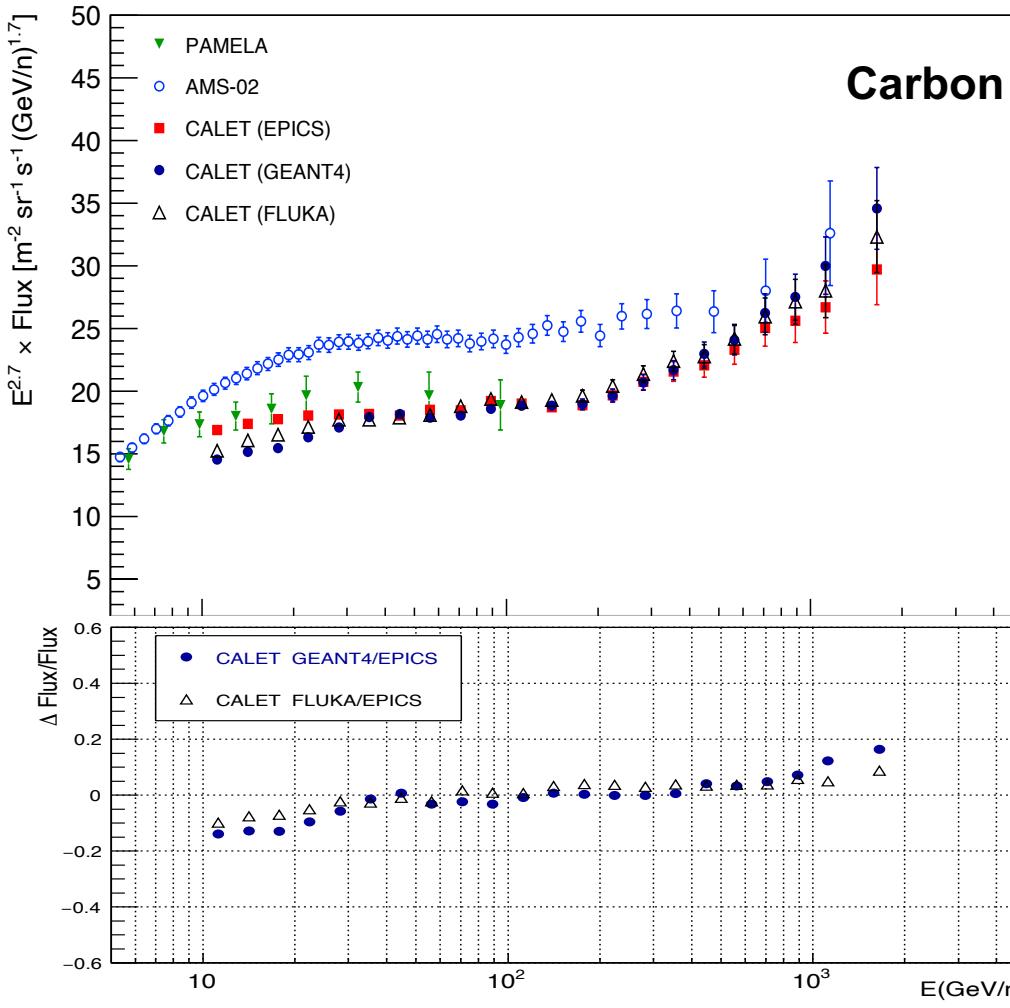
Energy resolution $\sim 30\%$

The energy response derived from MC simulations was tuned using the beam test results.





Mote Carlo models



MC simulations, reproducing detector configuration, physics processes, and detector signals, were developed based on three simulation packages

- EPICS 9.21 w/ DPMJET-III
- Fluka 2011 2c.6 w/ DPMJET-III
- GEANT4 10.5 w/ FTFP_BERT

MC simulations were tuned using beam test and flight data.

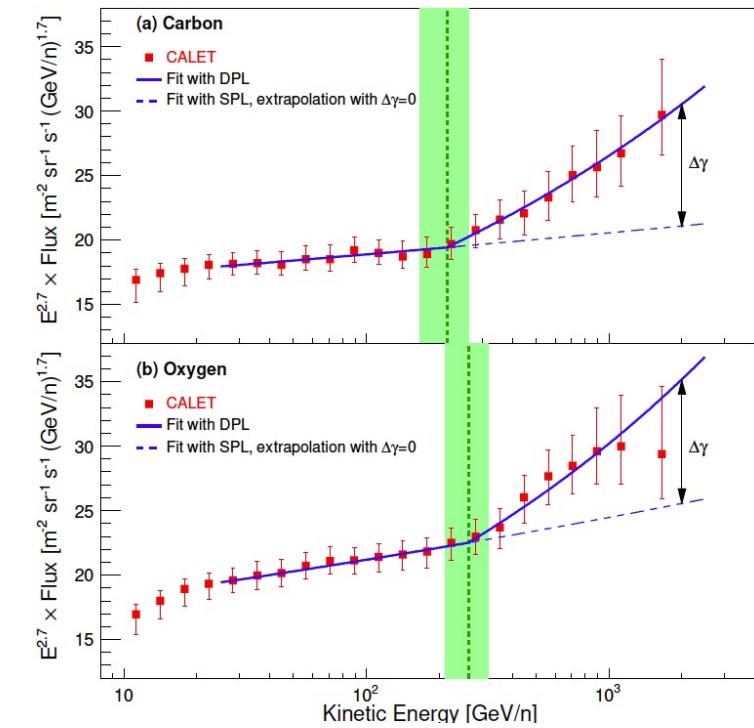
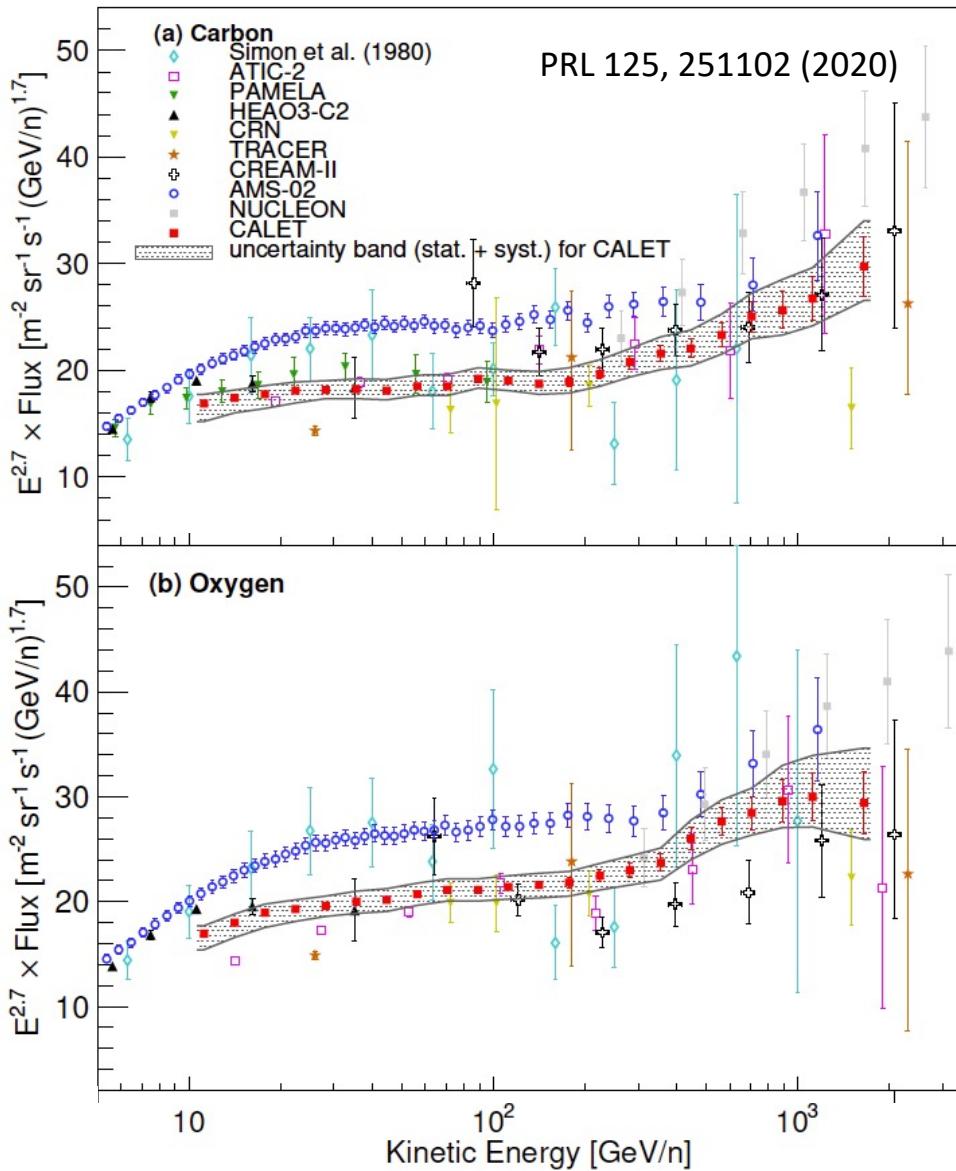
They are used to estimate selection efficiencies and response matrix.

Comparison of energy responses from different MC at high energy where no beam calibration is available.

The resultant fluxes from the analyses with different MC's show consistent normalization and spectral shapes.



Carbon and Oxygen Spectra



$$\Phi(E) = \begin{cases} C \left(\frac{E}{\text{GeV}}\right)^\gamma & E \leq E_0 \\ C \left(\frac{E}{\text{GeV}}\right)^\gamma \left(\frac{E}{E_0}\right)^{\Delta\gamma} & E > E_0 \end{cases}$$

Carbon

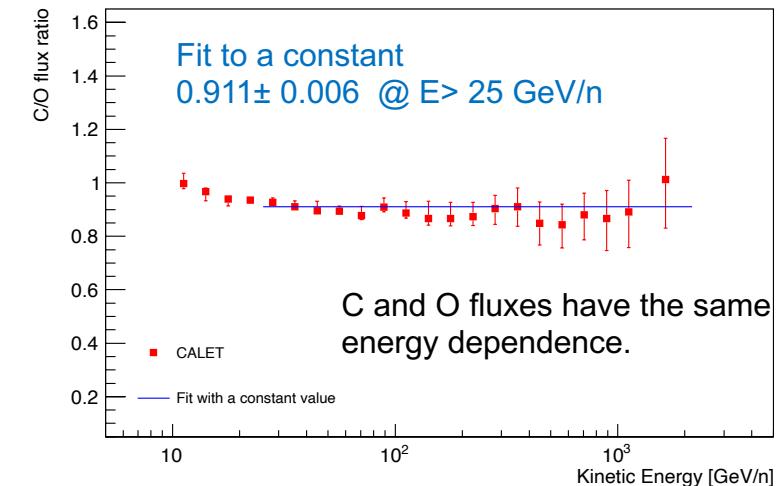
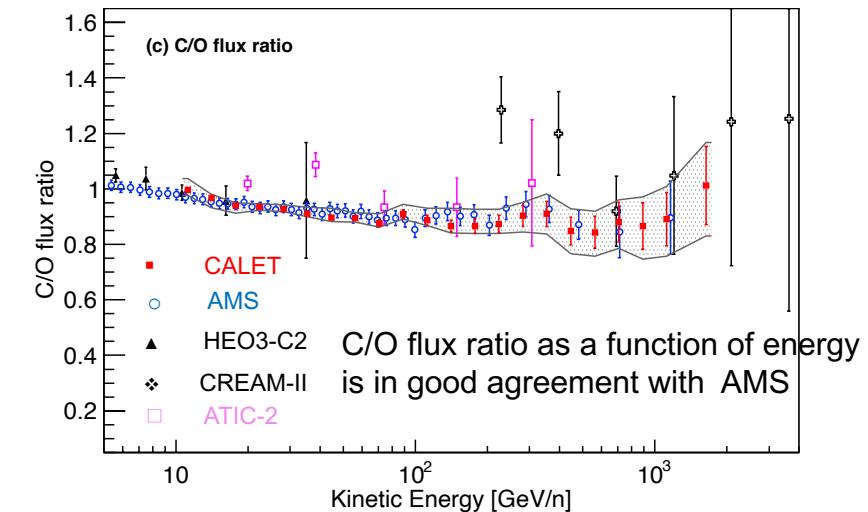
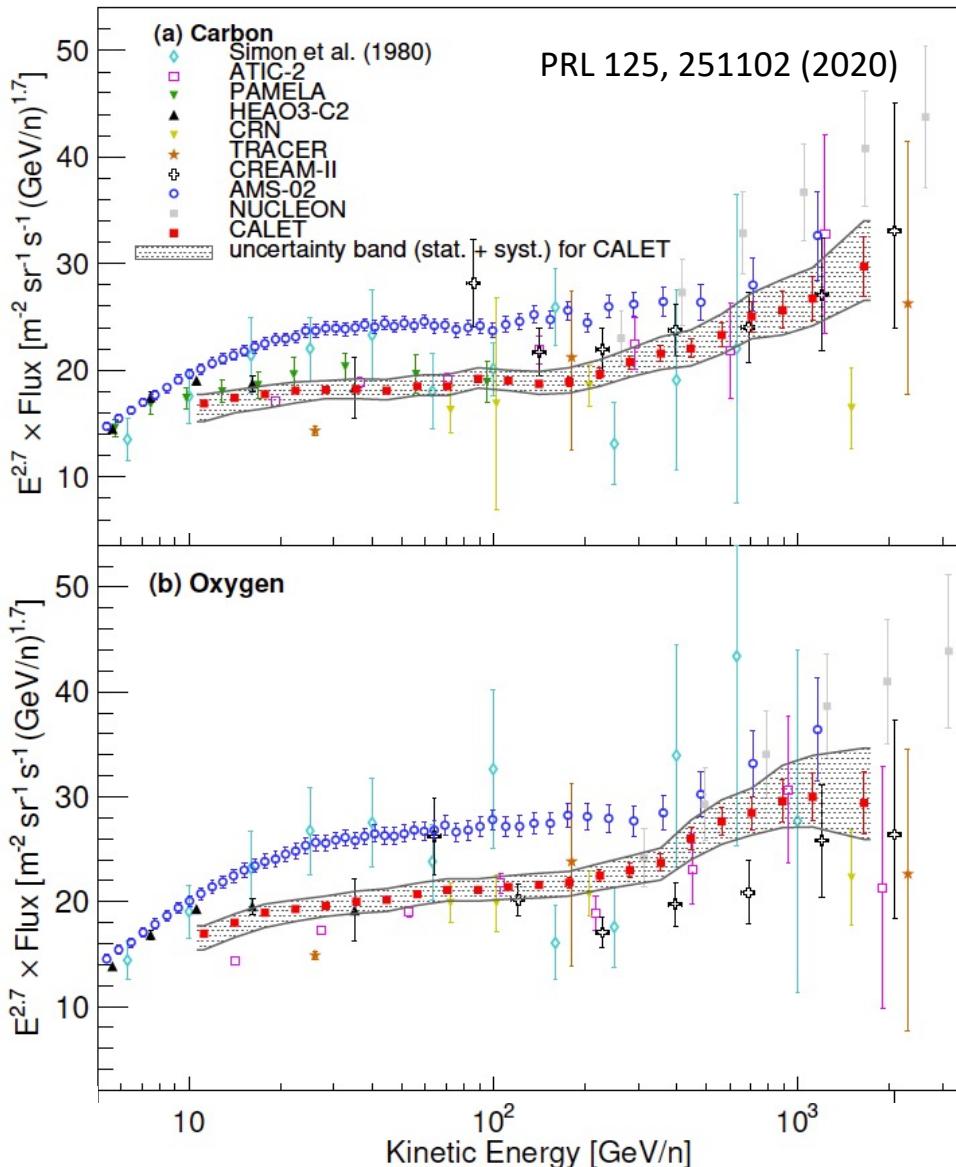
$$\left\{ \begin{array}{l} \gamma = -2.663 \pm 0.014 \\ E_0 = 215 \pm 54 \text{ GeV/n} \\ \Delta\gamma = 0.166 \pm 0.042 \text{ (4.0 }\sigma\text{)} \\ \text{with } \chi^2/\text{d.o.f.} = 9.0/8 \end{array} \right.$$

Oxygen

$$\left\{ \begin{array}{l} \gamma = -2.637 \pm 0.009 \\ E_0 = 264 \pm 53 \text{ GeV/n} \\ \Delta\gamma = 0.158 \pm 0.053 \text{ (3.0 }\sigma\text{)} \\ \text{with } \chi^2/\text{d.o.f.} = 3.0/8 \end{array} \right.$$

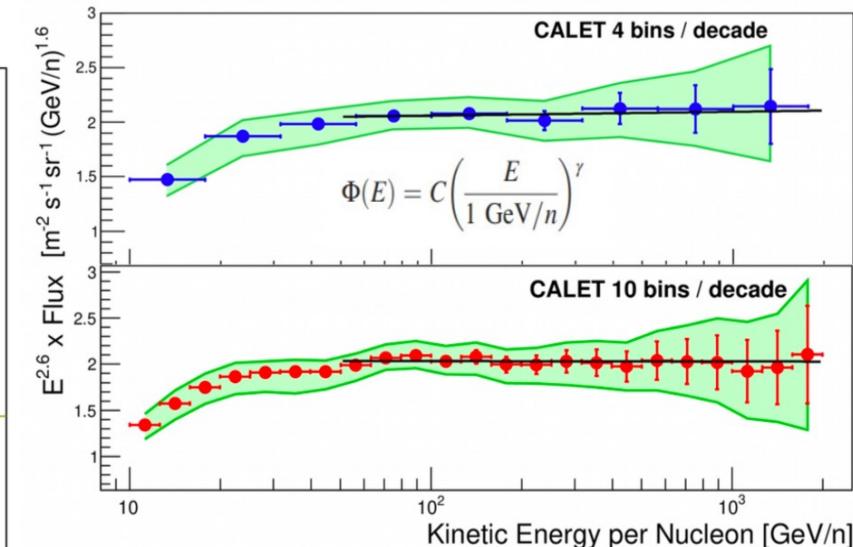
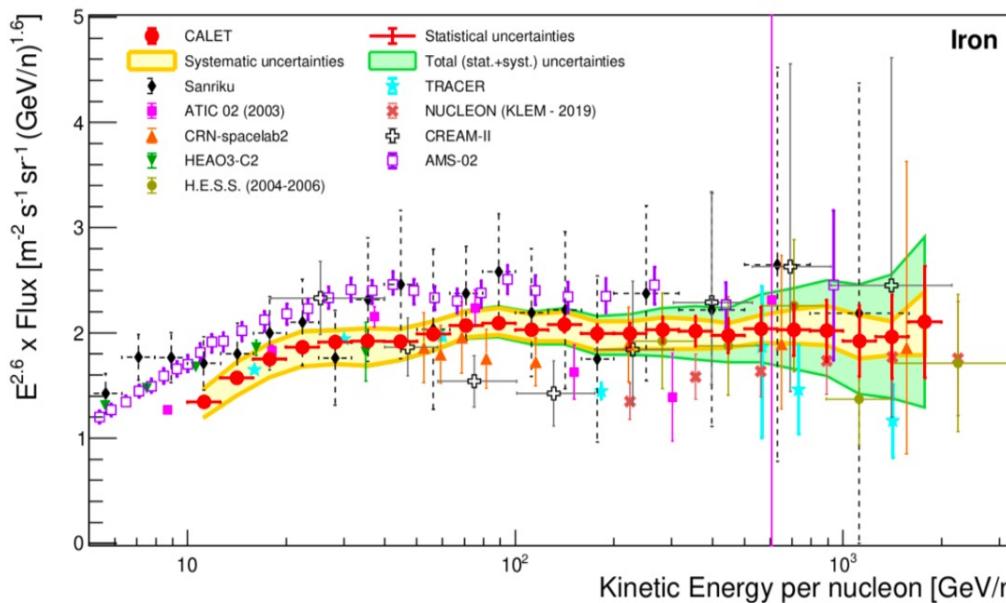


Carbon and Oxygen Spectra





Iron spectrum



10bins/dec: $\gamma = -2.60 \pm 0.02 \text{ (stat)} + 0.02 \text{ (sys)}$

4bins/dec: $\gamma = -2.59 \pm 0.02 \text{ (stat)} + 0.04 \text{ (sys)}$

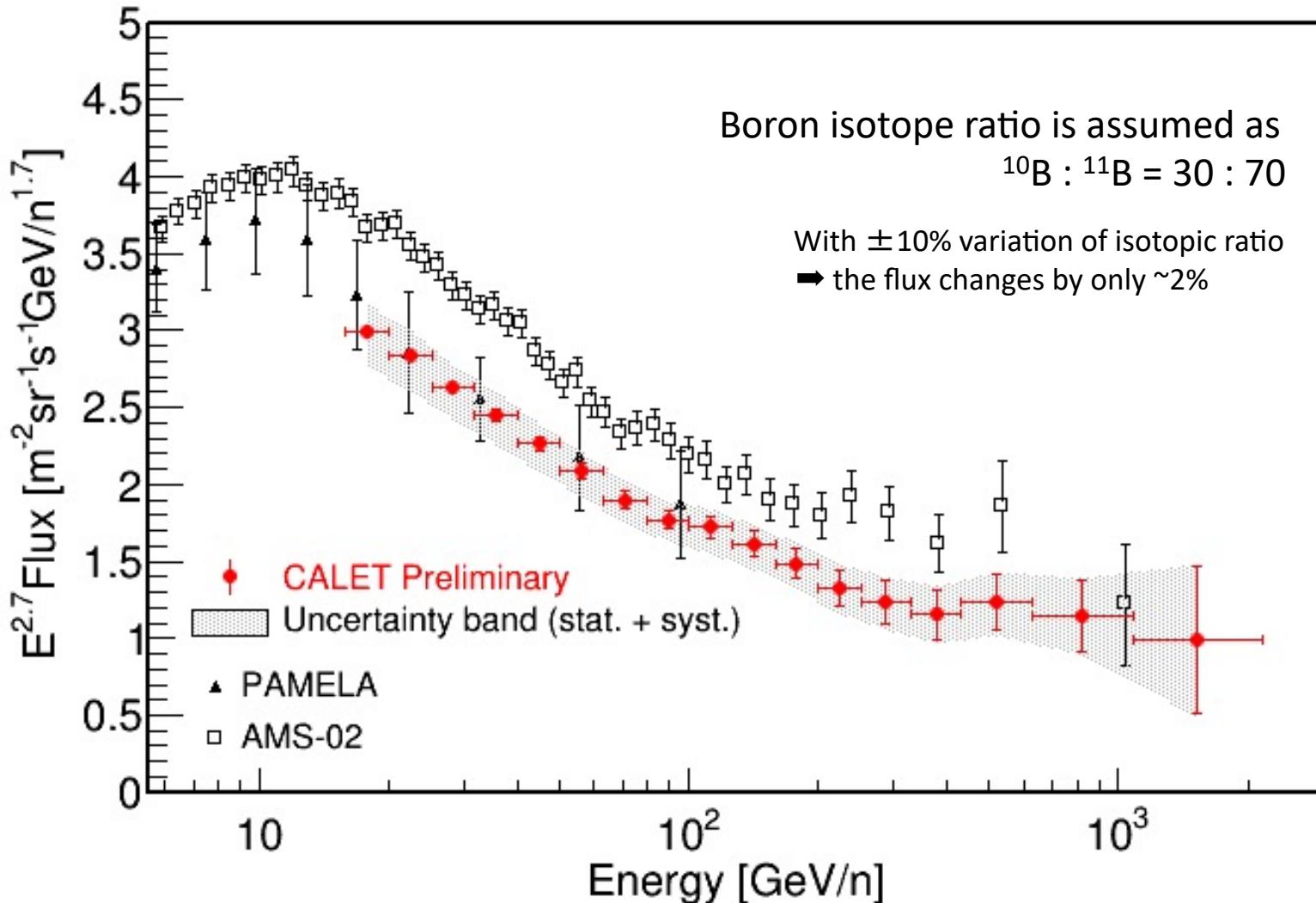
→ stable when larger energy bins are used

The iron flux, above 50 GeV/n, is compatible within the errors
With a single power law



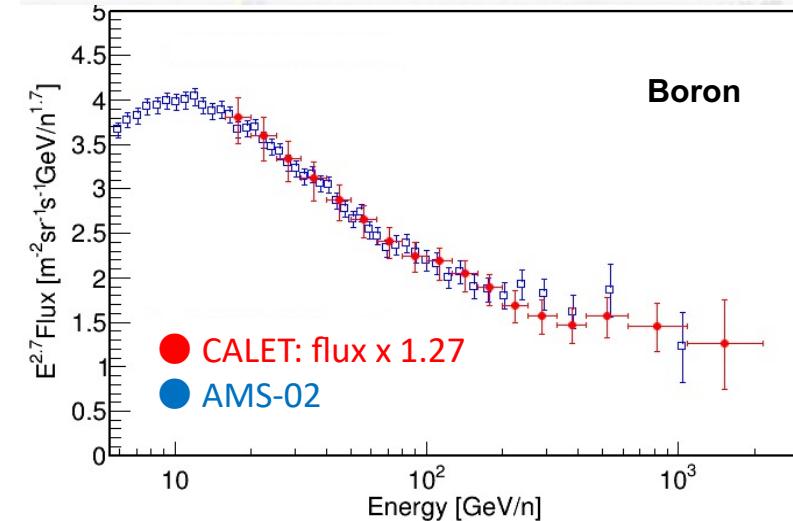
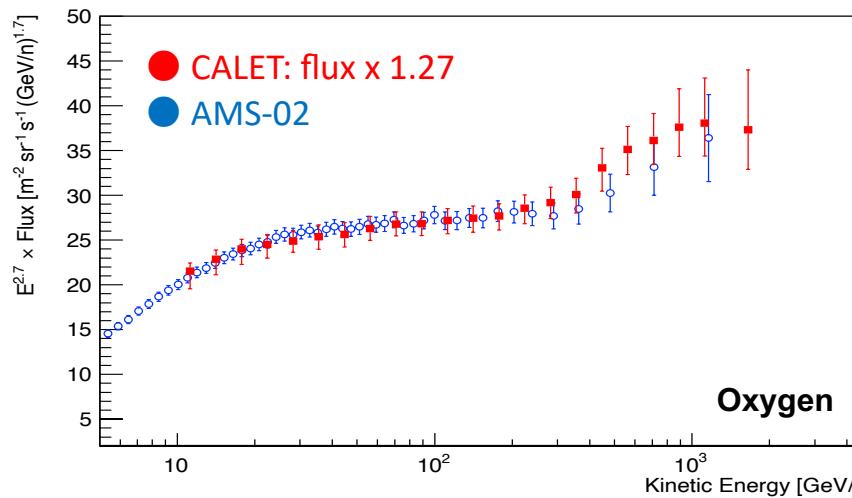
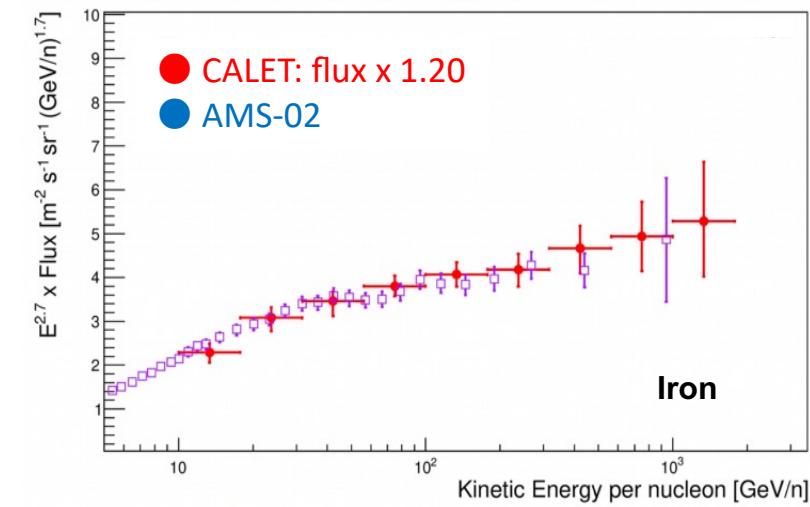
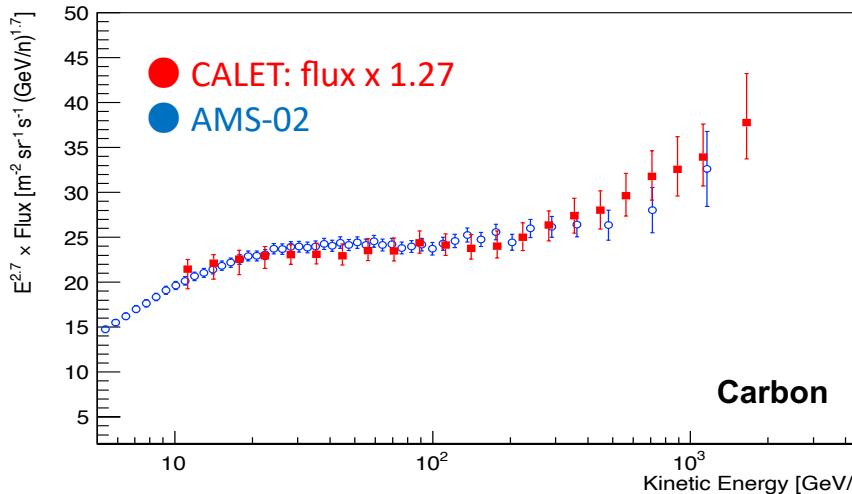
Boron spectrum

CALET preliminary result is consistent with PAMELA, but lower than AMS-02





Fluxes normalized to AMS-02

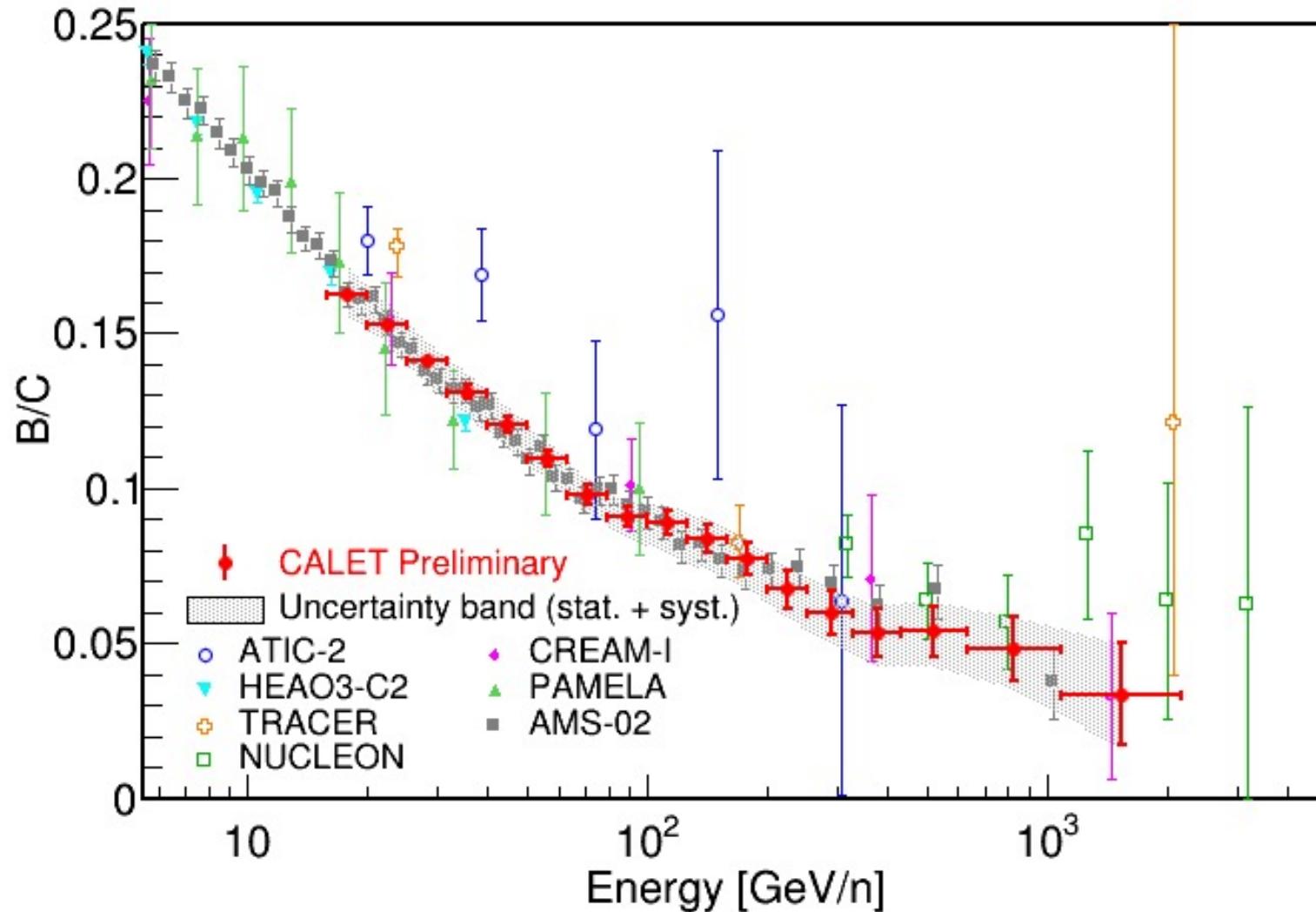


The spectral shapes and their spectral indices are consistent with AMS-02, but the absolute normalizations are lower than AMS-02.



Preliminary result of B/C ratio

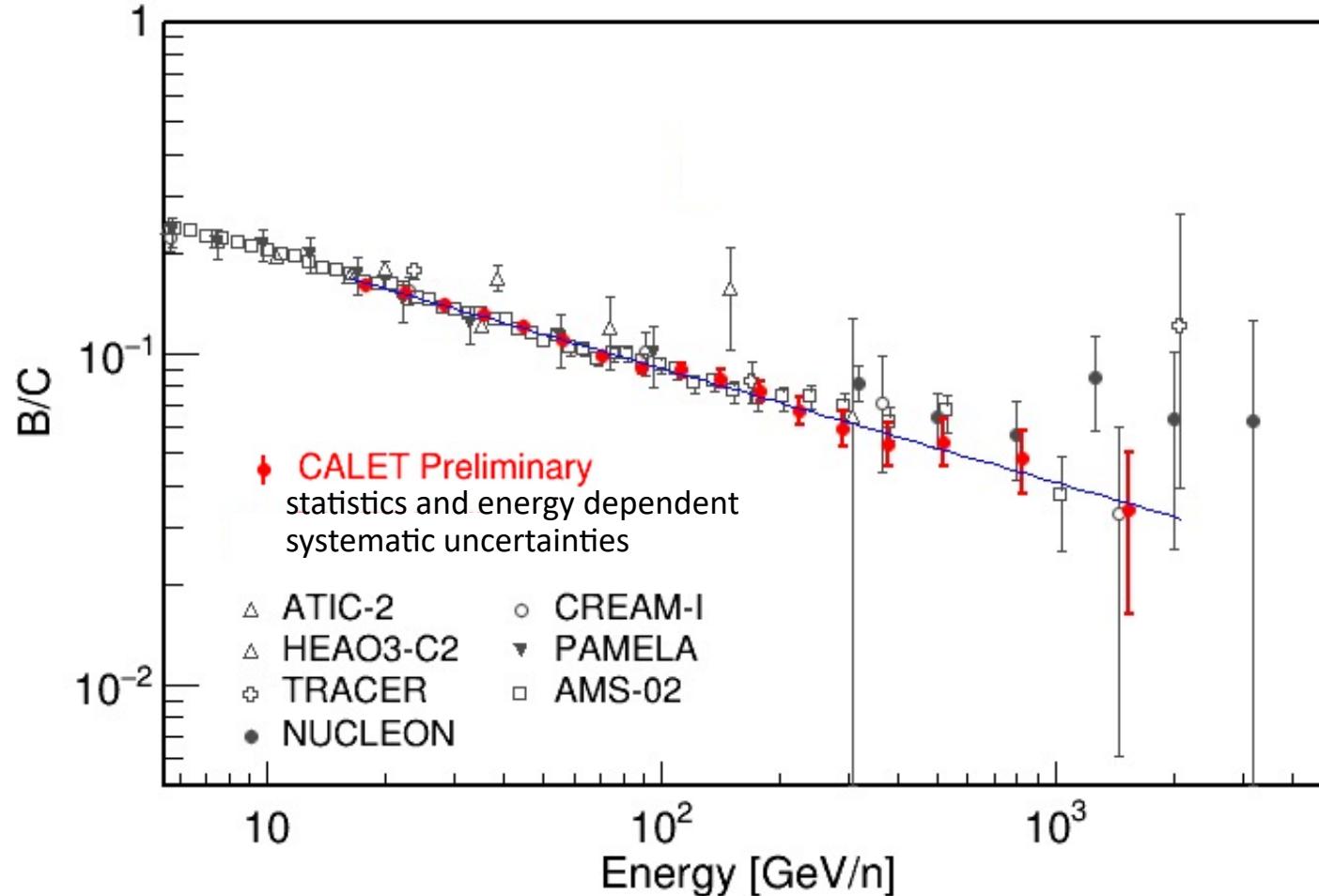
CALET preliminary result is well consistent with other experiments





Fit with a single power law function

The B/C ratio is fitted by a single power law function: $(B/C) = AE^\delta$
 $\delta = -0.344 \pm 0.012 (E > 20 \text{ GeV/n}), \chi^2/\text{ndf} = 5.6/15$





Summary

- CALET measures carbon, oxygen, iron and boron spectra and B/C ratio from 10 GeV/n to 2.2 TeV/n with 5 years of operations
- A single power law spectrum for C and O is excluded by more than 3σ
- A spectral index increase $\Delta\gamma=0.166 \pm 0.042$ for C and $\Delta=0.158 \pm 0.053$ for O is measured above 200 GeV/n
- The Fe spectral spectrum is consistent with the hypothesis of a SPL spectrum up to 2 TeV/n with a spectral index $\gamma = -0.260 \pm 0.03$
- Preliminary result of B/C ratio can be fit with a SPL function of $E^{-\delta}$; $\delta=0.344 \pm 0.012$ ($E>20$ GeV/n)
- Our results are consistent with the ones reported by AMS-02, as regards the spectrum shape and hardening. However, the absolute normalization of our data is significantly lower than AMS-02, but in agreement with other experiments
- We performed detailed systematics checks to search for possible causes of this normalization issue. We can exclude that it can stem from trigger inefficiencies, differences between MC simulation packages or hadronic models, or lacking modeling of the instrument.