

Latest Results from Observation of High-energy Gamma-rays by CALET

CALETによる高エネルギー ガンマ線観測の最新成果

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他CALETチーム

M. Mori, D. Zenita, Y. Asaoka, S. Torii, Y. Kawakubo, N. Cannady, M.L. Cherry for the CALET collaboration

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Japan Physical Society 2020 Fall Meeting (online), September 14-17, 2020

Gamma Ray Event Selection

= Electron Selection Cut + Gamma-ray ID Cut w/ Lower Energy Extension

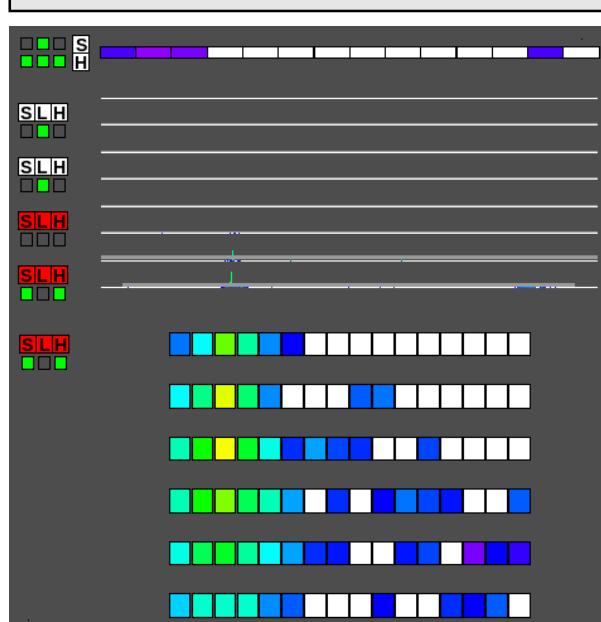
100 GeV Event Examples

gamma ray

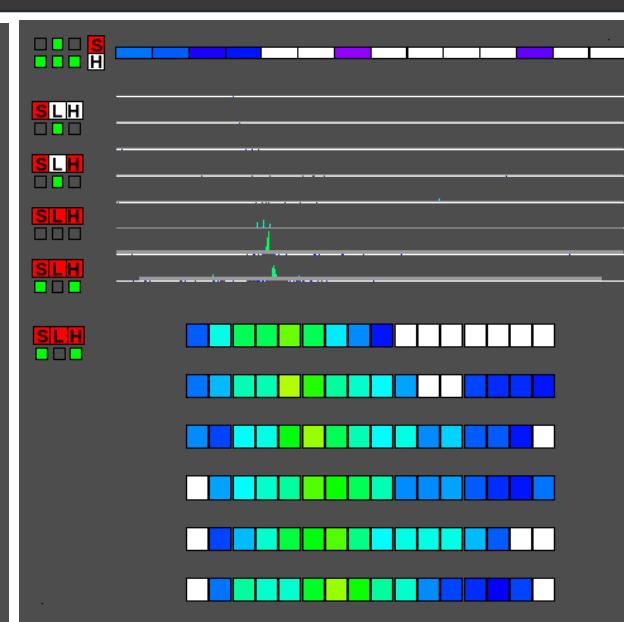
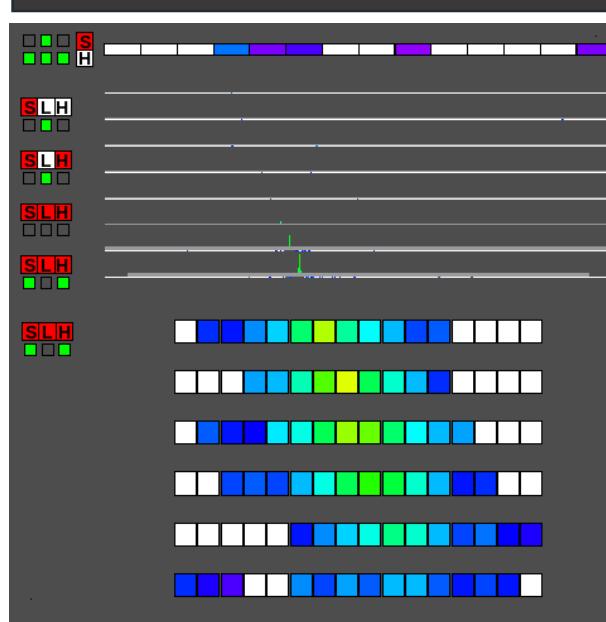
electron

proton

Charge Z=0



Charge Z=1



Electromagnetic Shower

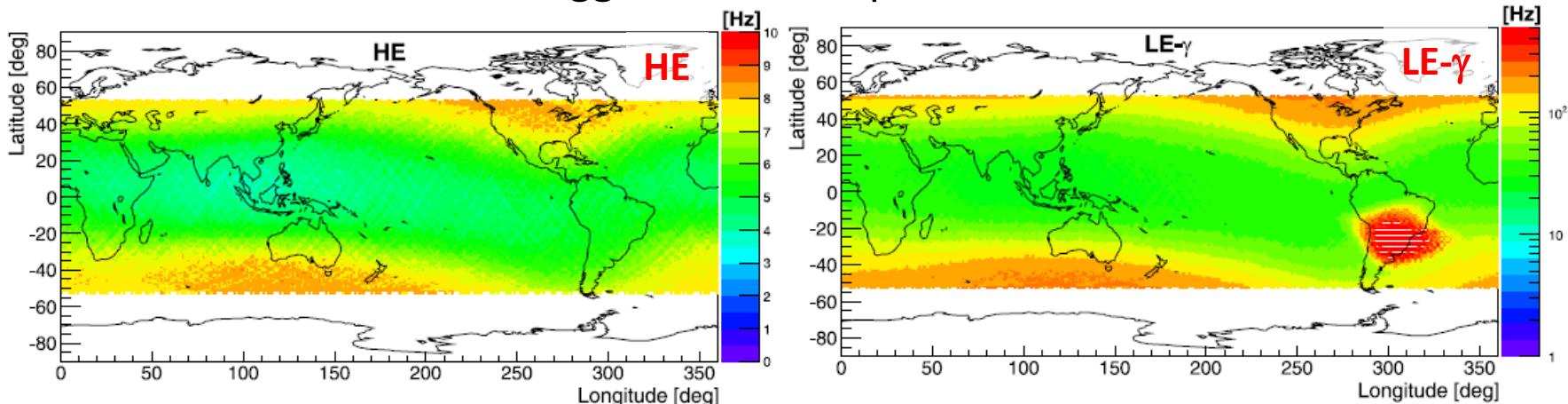
well contained, constant shower development

Hadron Shower

larger spread 2

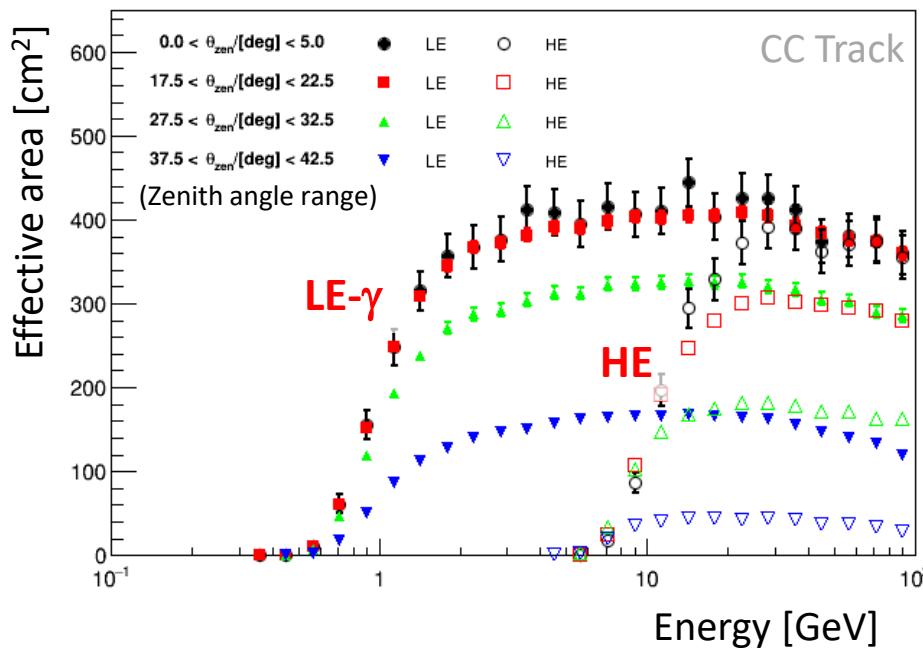
CALET triggers and gamma-ray observation

Trigger rate vs ISS position



Asaoka et al., Astropart.Phys. 100, 29 (2018)

Effective area

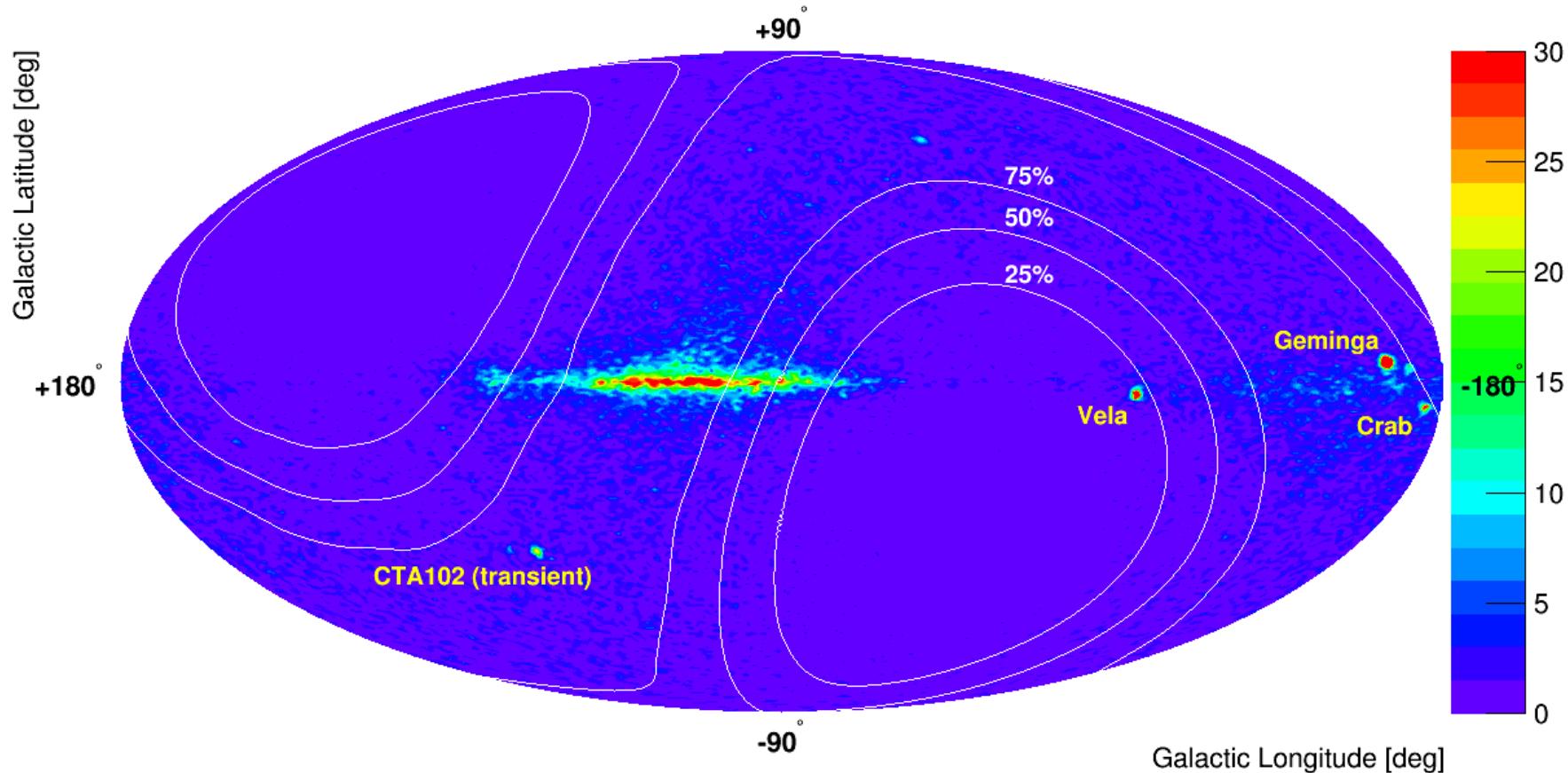


HE trigger: $E_{\gamma} > 10 \text{ GeV}$
LE- γ trigger: $E_{\gamma} > 1 \text{ GeV}$

- HE trigger mode: always ON
- LE- γ mode is activated when the geomagnetic latitude is below 20° and following a CALET Gamma-ray Burst Monitor (CGBM) burst trigger

Cannady et al., ApJS 238:5 (2018)

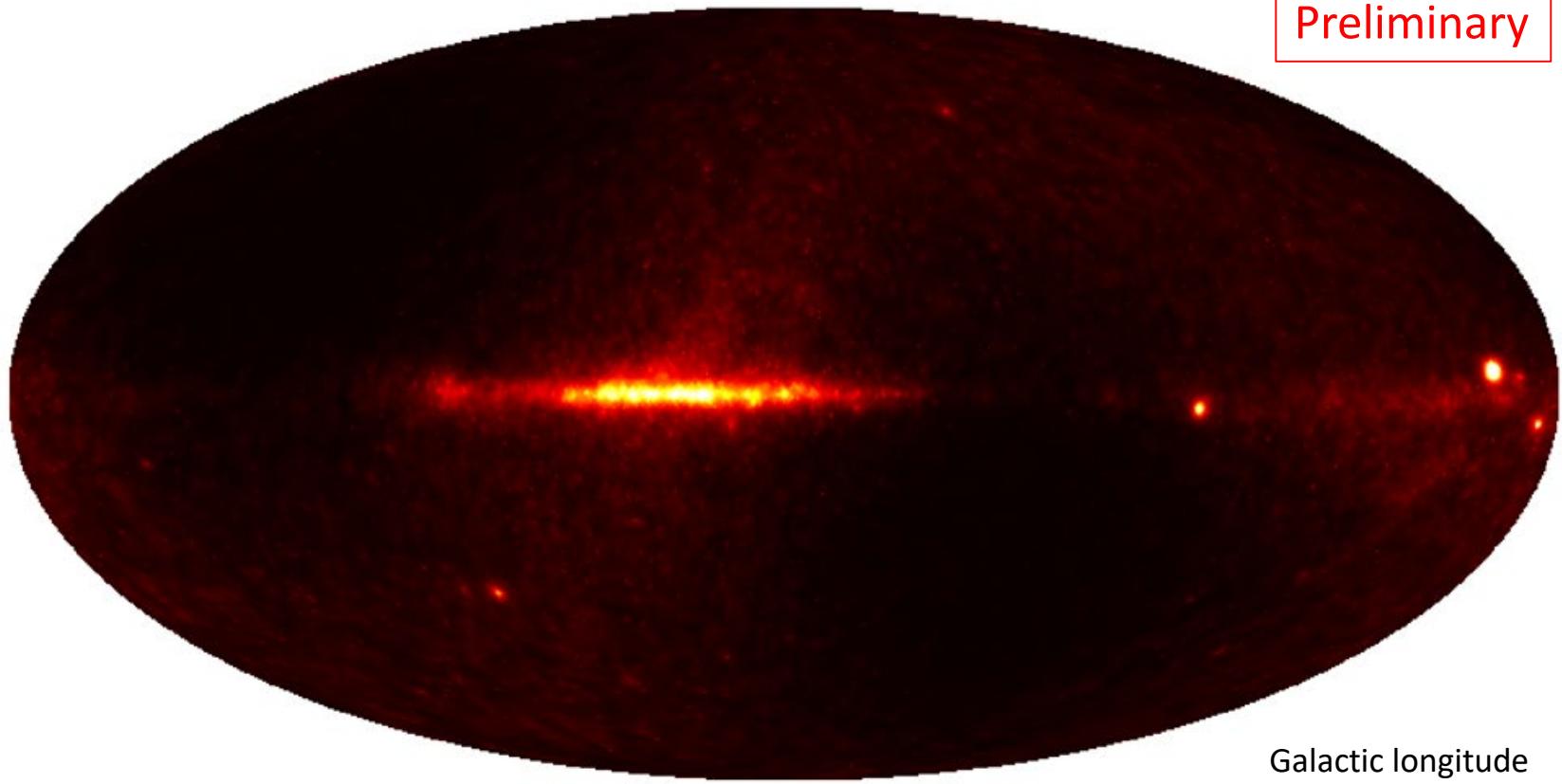
Gamma-ray skymap (2.5 yr)



LE- γ mode, from 2015 November to 2018 May
(Contours show relative exposures)

Gamma-ray skymap (4 yr)

Galactic latitude

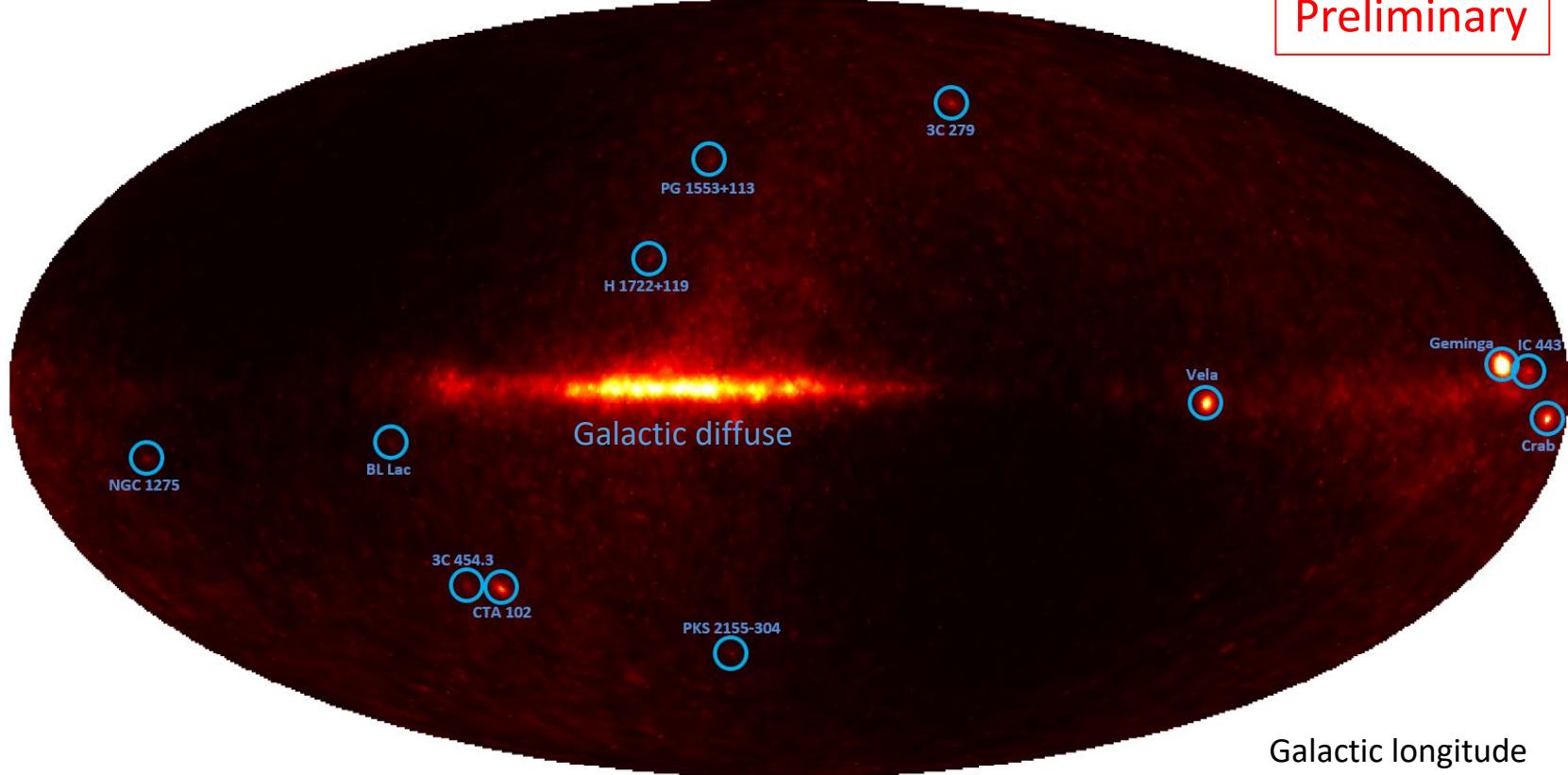


Galactic longitude

LE- γ mode, from 2015 November to 2019 Dec
(Note that the exposure is non-uniform)

Gamma-ray skymap (4 yr)

Galactic latitude

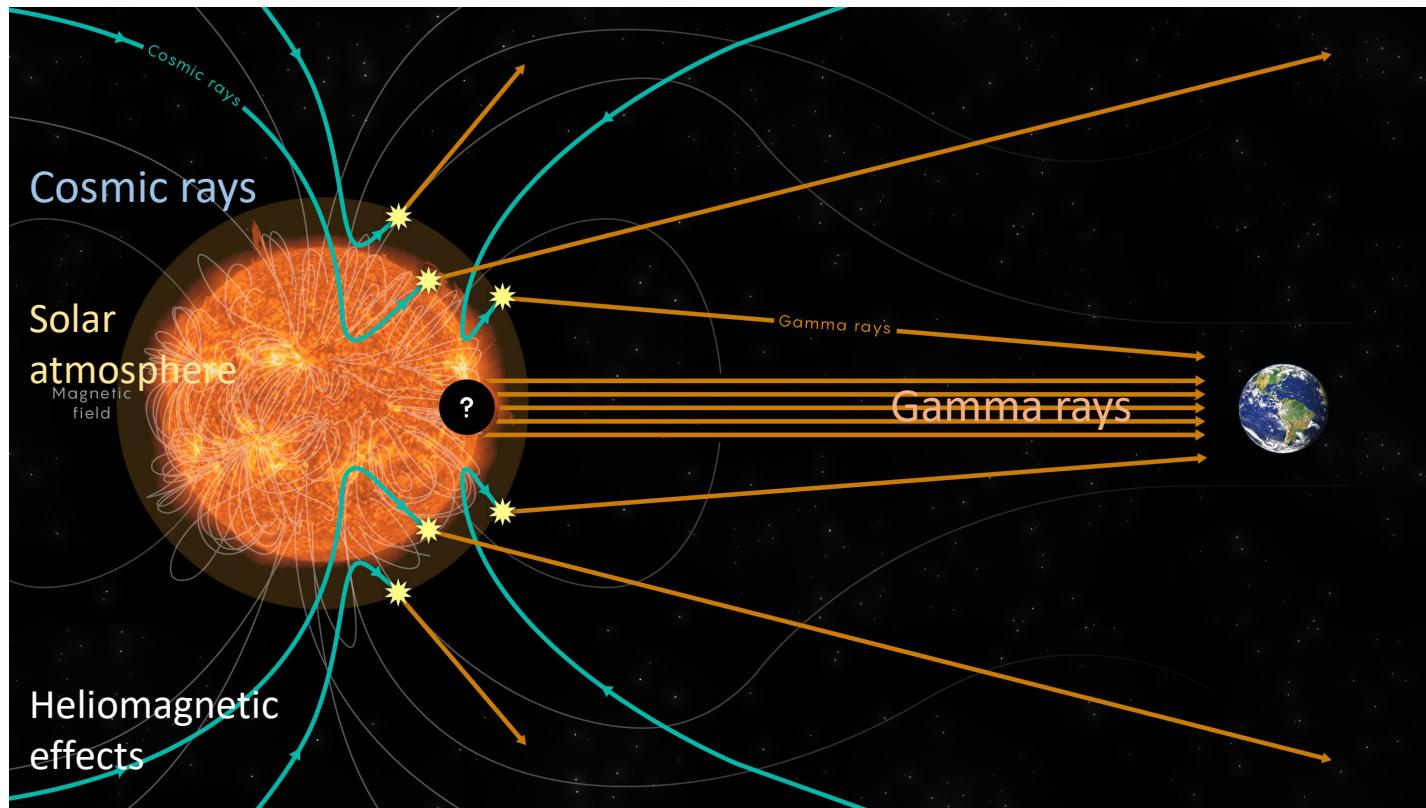


LE- γ mode, from 2015 November to 2019 Dec
(Note that the exposure is non-uniform)

Solar atmospheric gamma-rays

(Gamma-rays from solar disk and halo)

<https://www.quantamagazine.org/gamma-ray-data-reveal-surprises-about-the-sun-20190501/>



“Disk” (cascade) emission: Galactic CR + solar surface → gamma rays

“Extended” (IC) emission : CR electrons + solar photon halo → gamma rays

Fermi-LAT, ApJ 734:116, 2011

→ Reflect CR modulation (and solar activities) and HE CR electrons

Solar atmospheric gamma-rays

Tang et al., PR D98, 063019 (2018)
 Linden et al., PRL 121, 131103 (2018)

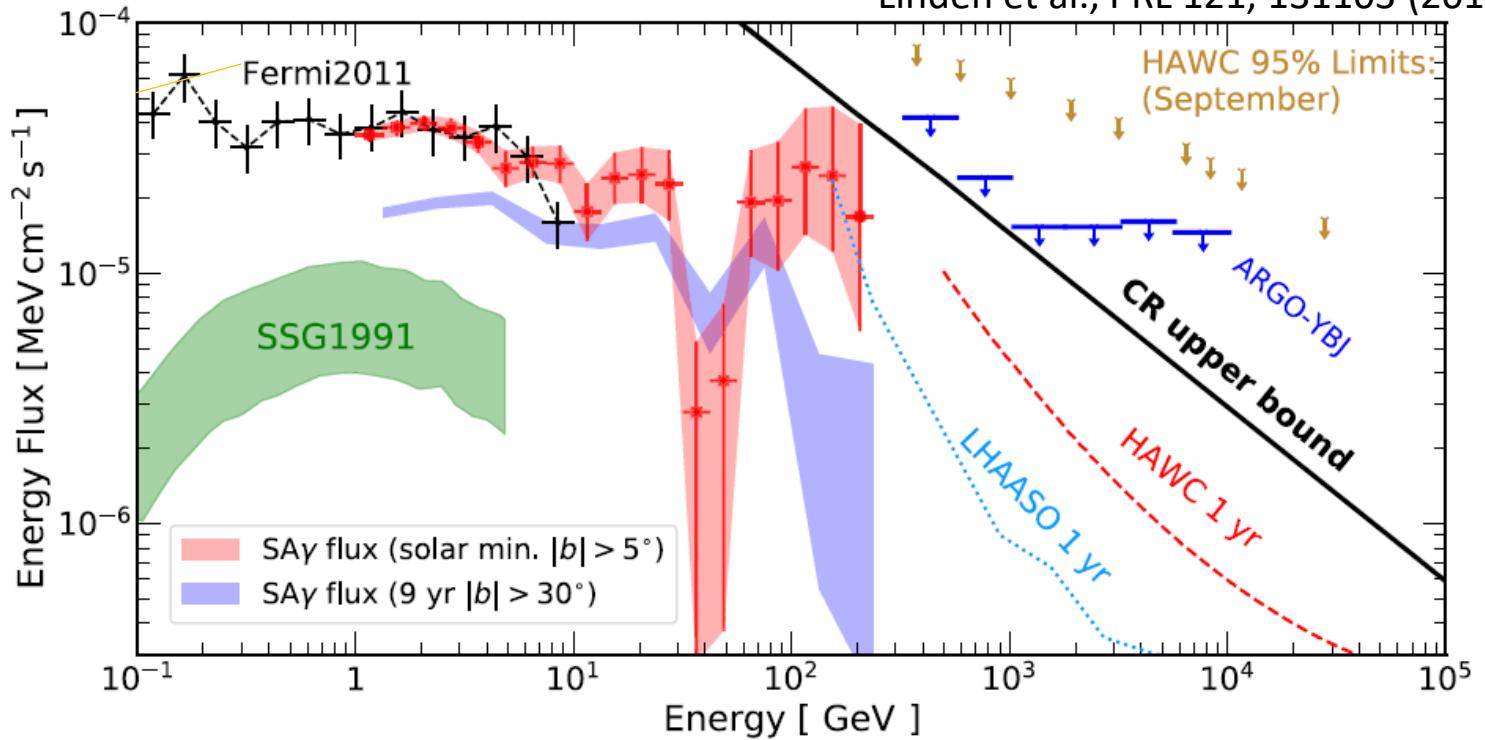


FIG. 5. The 9 year averaged SA γ flux and the solar minimum SA γ flux (76 weeks of data from 2008-8-7 to 2010-1-21), which is significantly harder above 100 GeV. We also show the point source sensitivities of HAWC [23] and LHAASO [24,25], as well as the preliminary limits from HAWC ([47], one month of data) and ARGO-YBJ [48]. The theoretical maximum gamma-ray flux that the Sun can produce with cosmic rays (CR upper bound; see text for detail) is shown by the black solid line.

SSG1991: Seckel, Stanev & Gaisser ApJ 382, 652 (1991)

CR electrons and protons interacting with Solar photosphere and corona

Solar atmospheric gamma-rays

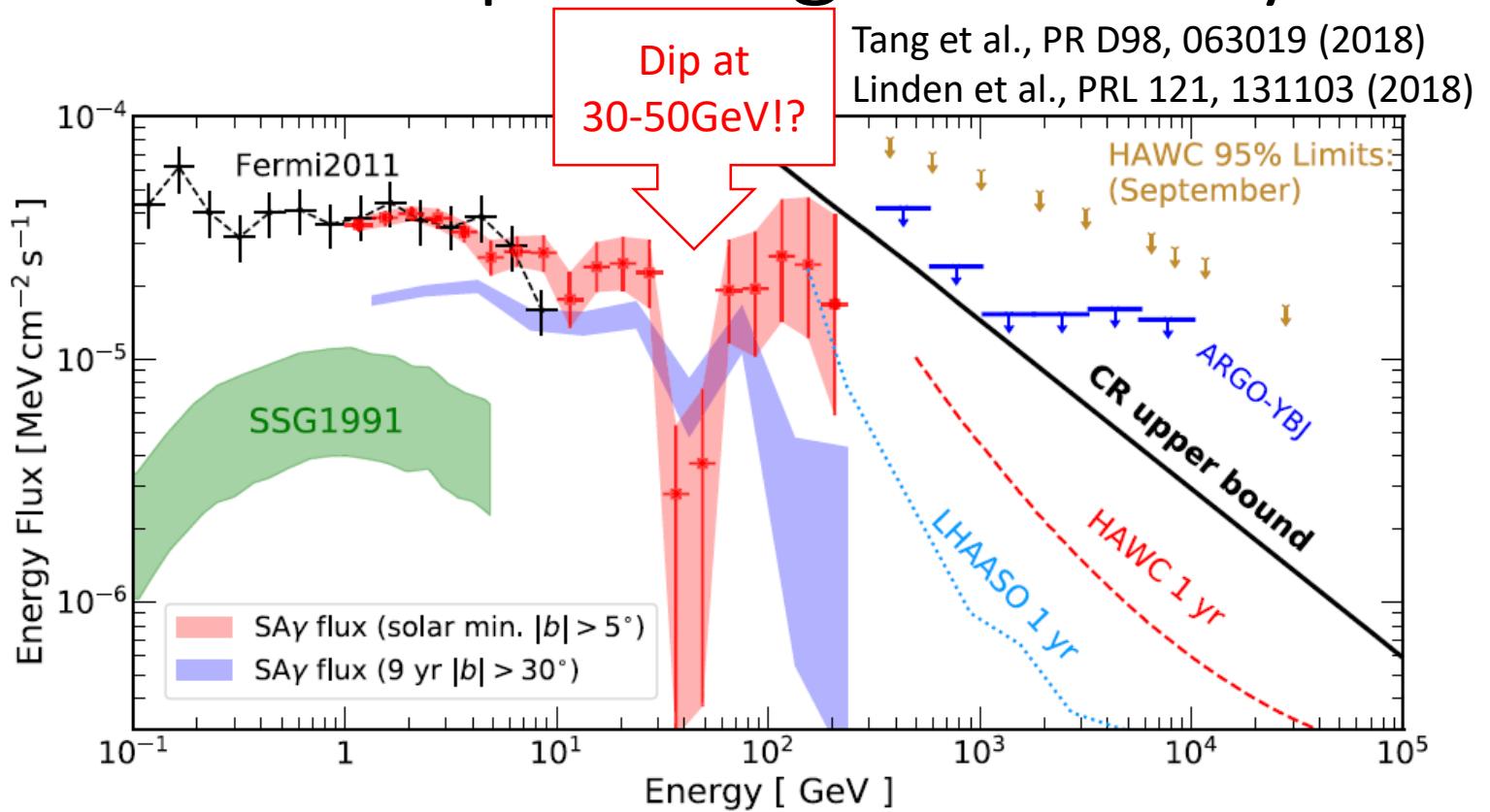


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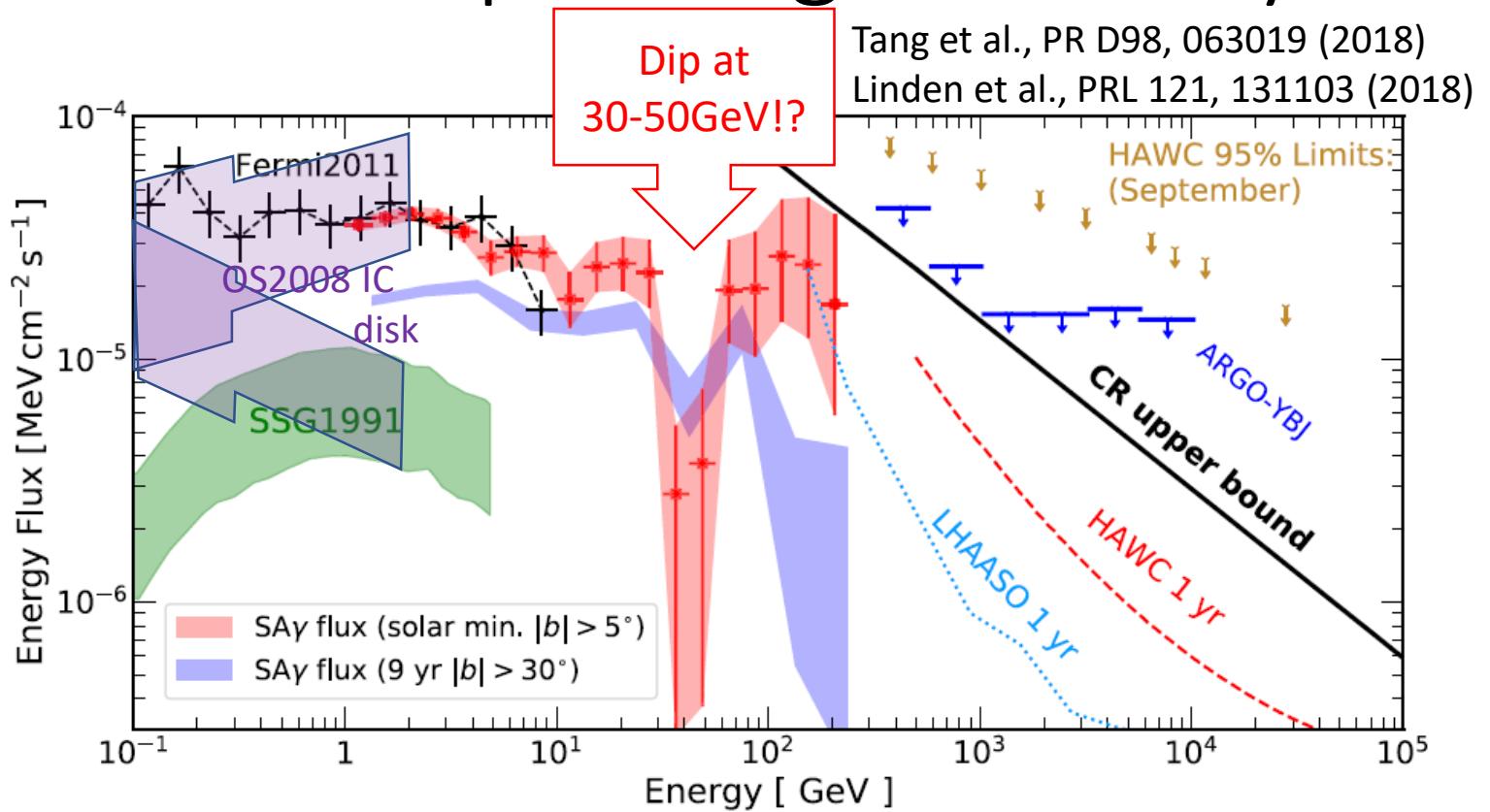


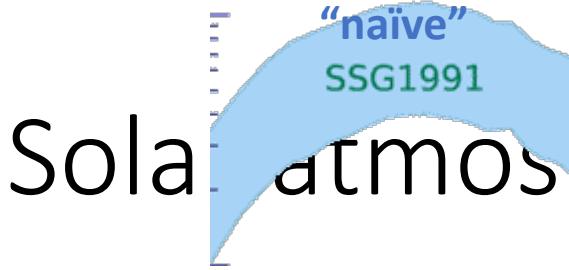
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CR electrons and protons interacting with Solar photosphere and corona

OS2008: Orlando & Strong AA 480, 847 (2008)

EGRET analysis for disk and extended (halo) emission



HAWC: PR D98, 123011 (2018)

Solar atmospheric gamma-rays

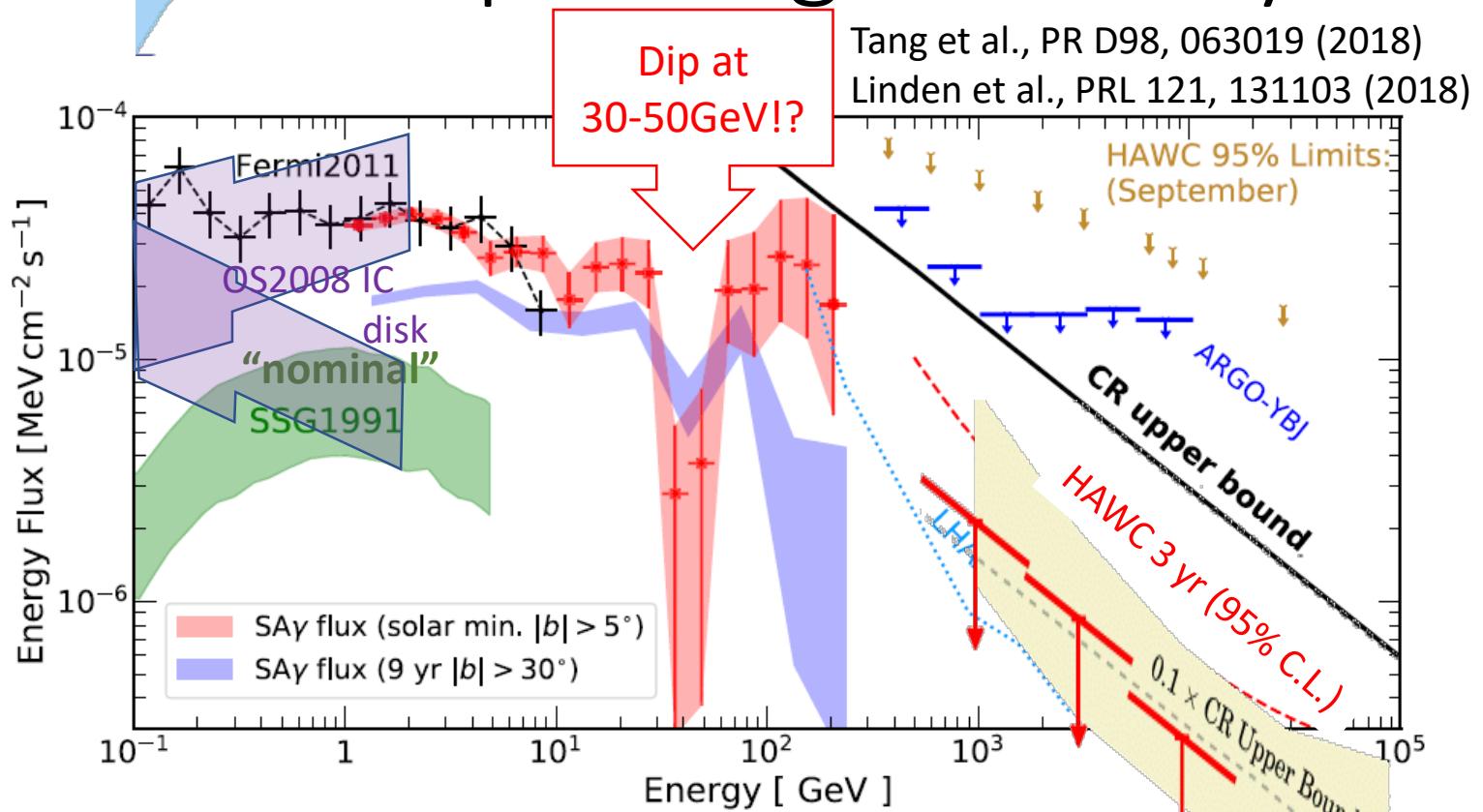


FIG. 5. The 9 year averaged S γ flux and the solar minimum S γ flux (76 weeks of data from 2008-8-7 to 2010-8-7) which is significantly harder above 100 GeV. We also show the point source sensitivities of HAWC [23] and LHAASO [24,25], as well as the preliminary limits from HAWC ([47], one month of data) and ARGO-YBJ [48]. The theoretical maximum gamma-ray flux that the Sun can produce with cosmic rays (CR upper bound; see text for detail) is shown by the black solid line.

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Observations with CALET

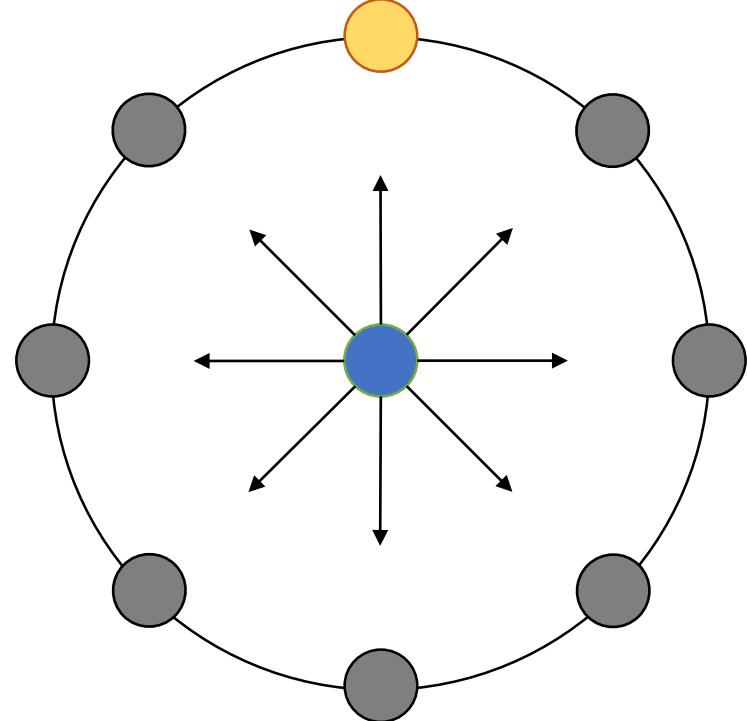
- Data sample covers the time period 2015/11 – 2019/10 (48 months)
 - LE- γ data ($E > 1$ GeV) available for low geomagnetic latitude
 - HE data ($E > 10$ GeV) available with >95% duty cycle
- Bright background sources are vetoed for event selection and exposure generation:
 - Galactic plane ($|b| < 20^\circ$)
 - Moon (5° window)
 - daily 12:00 moon position for exposures
 - moon position at precise time for event selection
 - Point sources (5° window)
 - Pulsars, AGN, etc.
 - Some redundancy with galactic plane cut
- Rotated into solar frame defined in terms of Sun position and normal to ecliptic plane

Analysis methodology

- Selection of events
 - Events chosen using energy- and tracking-dependent PSF relative to z-axis in Solar frame
 - Narrow window (< 2° overall, much smaller for most energies) contains primarily disk component. IC component contamination under investigation
- False Sun trials
 - Positions chosen along ecliptic leading or trailing actual Sun
 - In this analysis spaced by 45° (7 equally spaced false Suns)
 - Analysis repeated for these frames same as for real Sun frame
- Accumulation of exposure and background model
 - Full sky transformed for every day according to Sun position
 - Fermi-LAT background model used to validate observed event numbers

False Sun trials

- Better estimation of background than galactic diffuse model
- Sun position shifted further along the ecliptic plane
 - Calculate i^{th} fake Sun out of N using observation time $T_0 + (i * 365.25 \text{ days} / N)$
- Repeat analysis of Sun region in each false Sun window



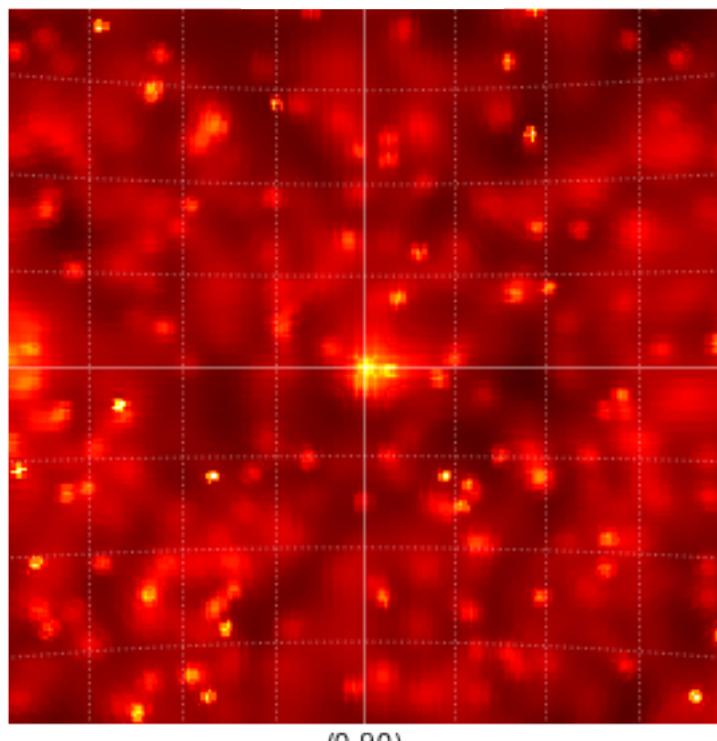
Currently using 7 false Sun trials, each shifted by 45° along the ecliptic, well outside of a 20° region of interest.

Count map: Sun frame vs. background frame

Preliminary

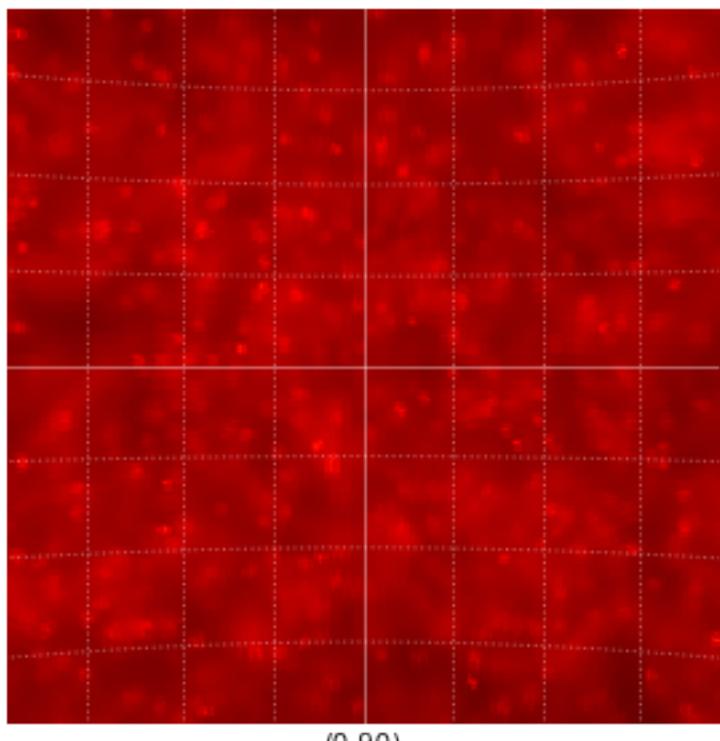
Sun frame

12 'pix, 200x200 pix



False Sun average

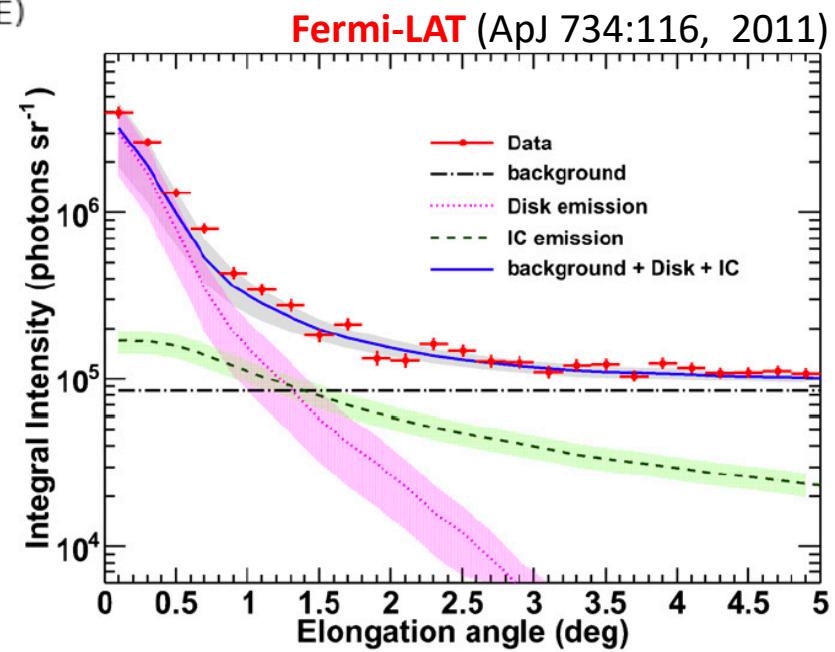
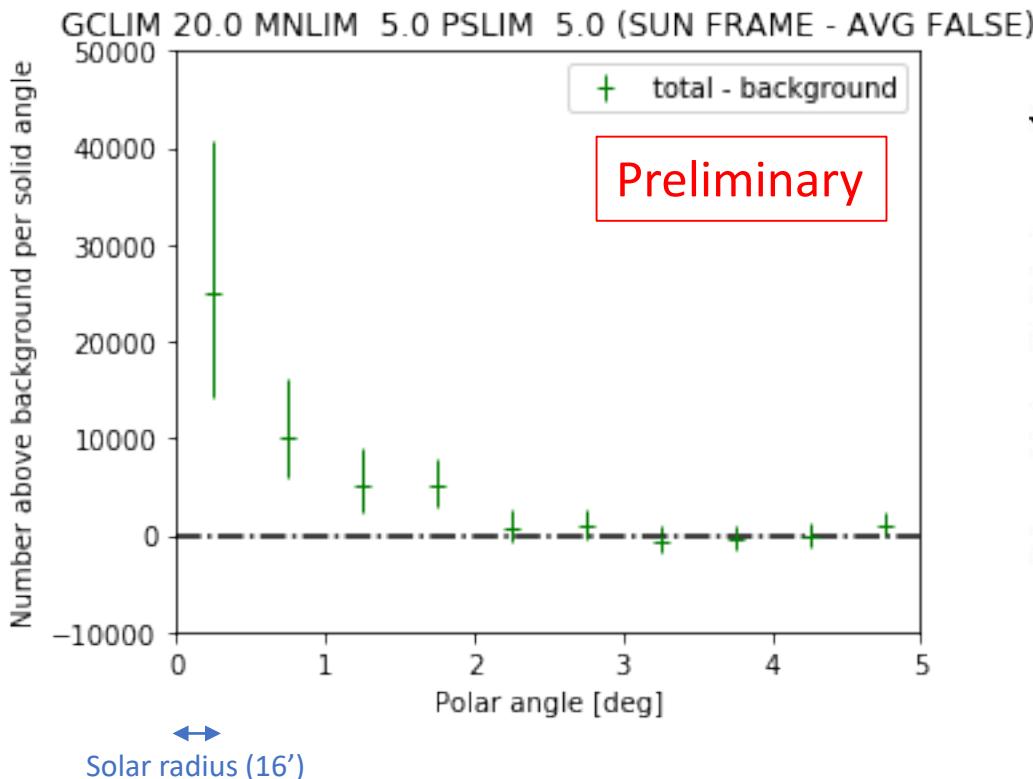
12 'pix, 200x200 pix



Events vs. opening angle

Background number of events calculated from average of false Sun trials

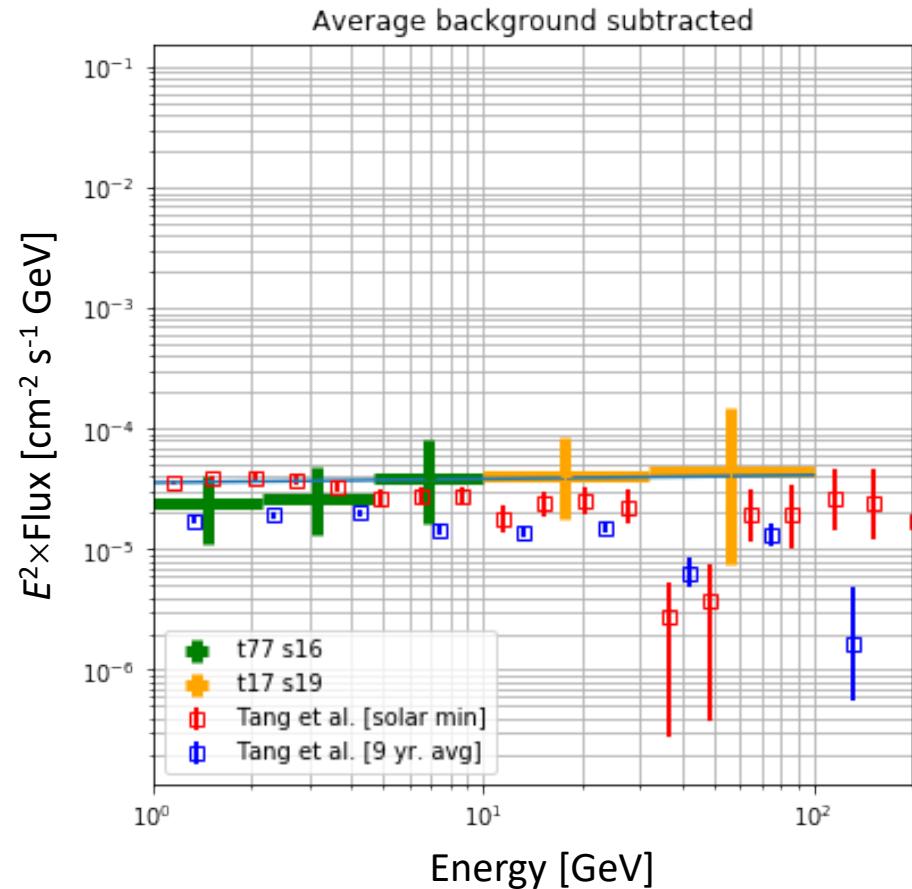
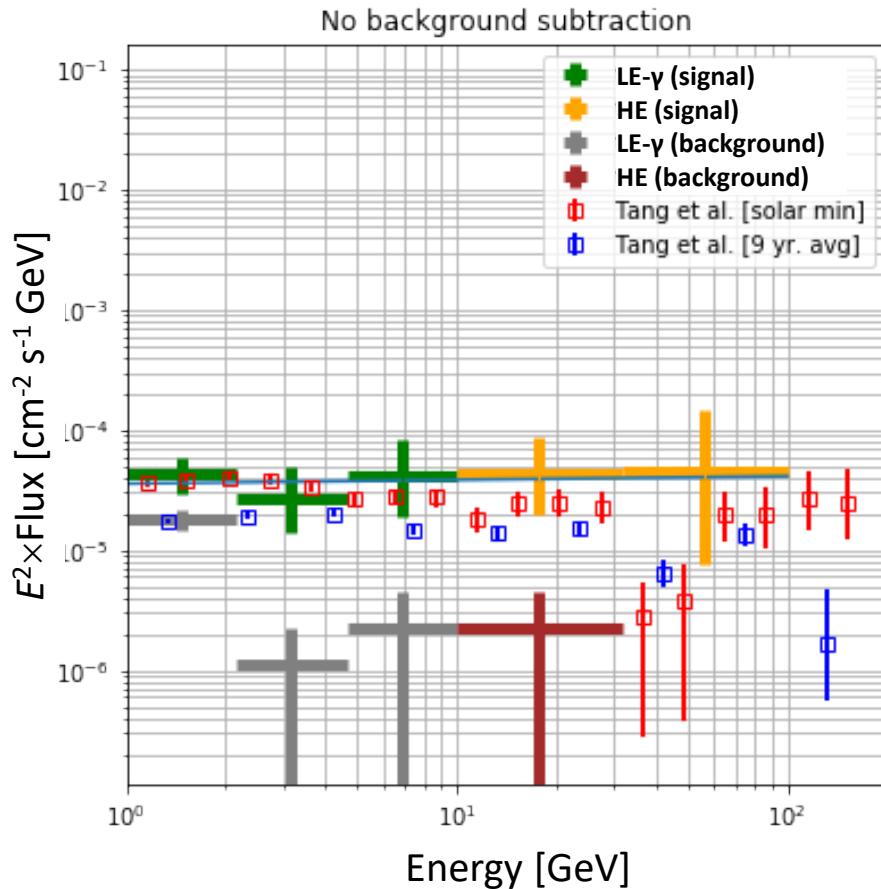
NB absolute value on y-axis is dependent on effective area and observation time



CALET gamma-ray spectrum from the Sun

Preliminary

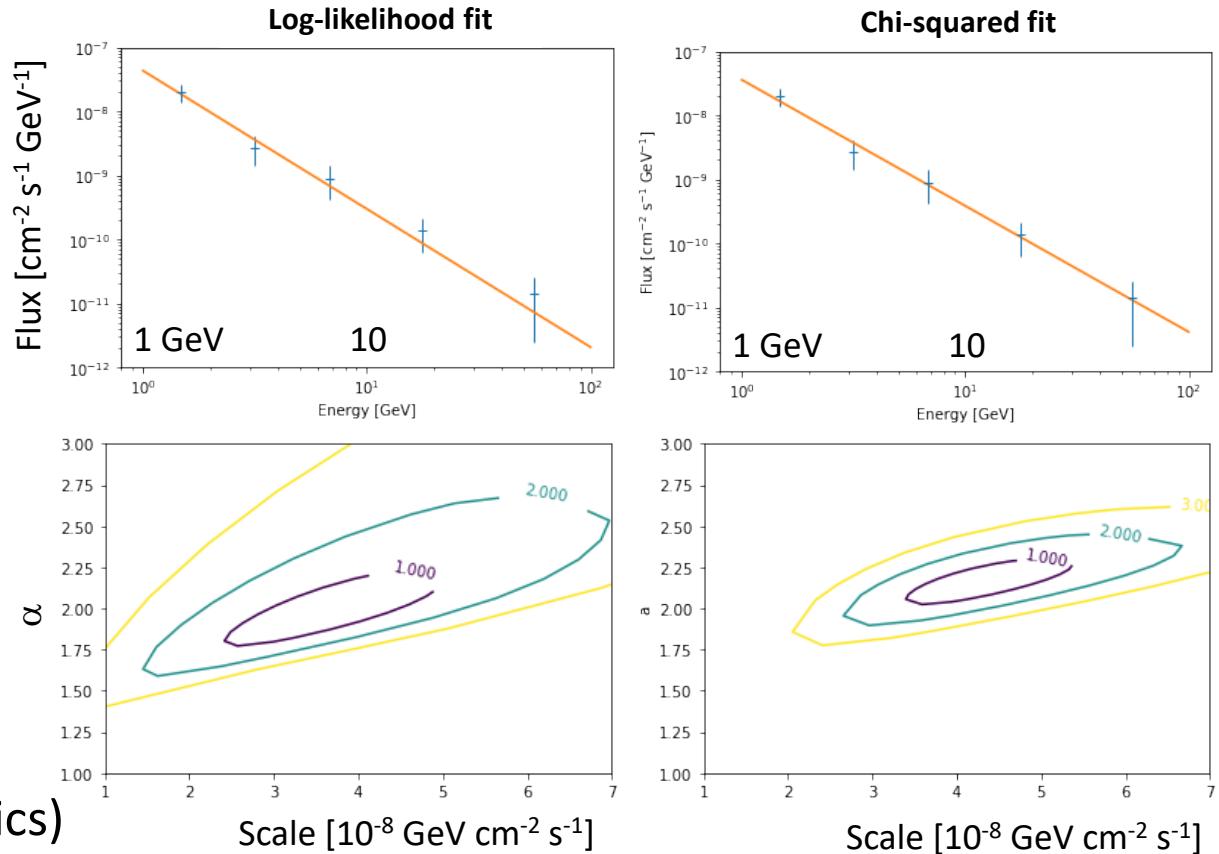
Flux considering only 2017 – 2018 data (low sunspot number)

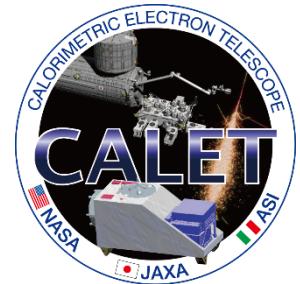


CALET spectral fit

Preliminary

- Preliminary flux calculated and a power law ($dN/dE \propto E^{-\alpha}$) was fit
 - CAVEAT: No treatment of the IC flux is taken into account (for now)
- Chi-squared fitting
 - Scale: $(3.6 \pm 1.2) \times 10^{-8}$
 - Index: (1.97 ± 0.22)
(error statistical only)
- Likelihood fitting
 - Scale: $(4.3 \pm 1.1) \times 10^{-8}$
 - Index: (2.16 ± 0.14)
(error statistical only)
- Fermi2011 results
 - Index: (2.11 ± 0.73)
(includes full systematics)





Summary

- CALET is monitoring the gamma-ray sky above 1 GeV for more than 4 years since 2015.
- CALET detects the Sun in GeV gamma rays significantly over the background signal
 - Events from the 48-month dataset clearly shows excess toward the Sun above 1 GeV
 - Power-law spectral fit is consistent with Fermi-LAT (Tang et al. 2018) continuum
 - Dip at 30-50 GeV not detected in CAL analysis, but not confidently ruled out
 - Continued accumulation of statistics in Solar minimum