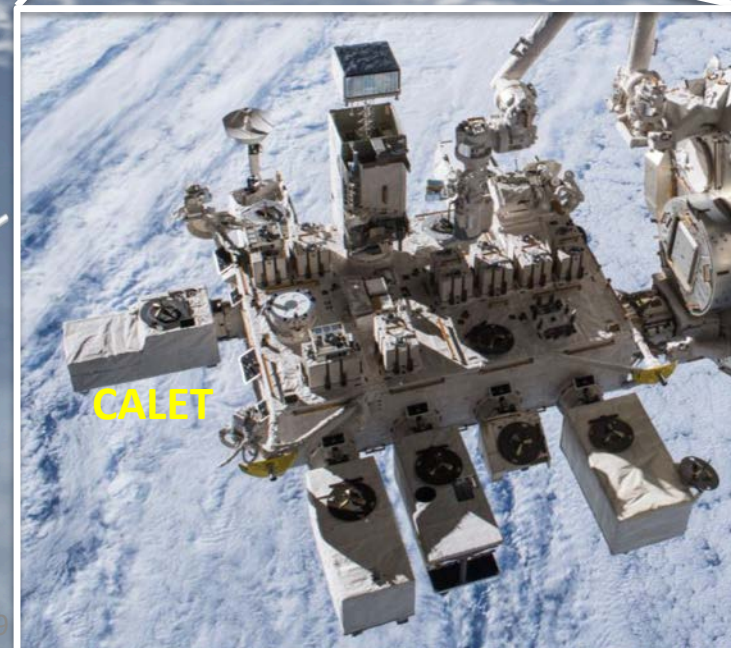




# Extended Measurement of Cosmic-Ray Electron and Positron Spectrum from CALET on the ISS

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Waseda University  
& Kanagawa University  
for the CALET collaboration

CRD2c 25 July, 2019



25, 2019



# CALET Payload

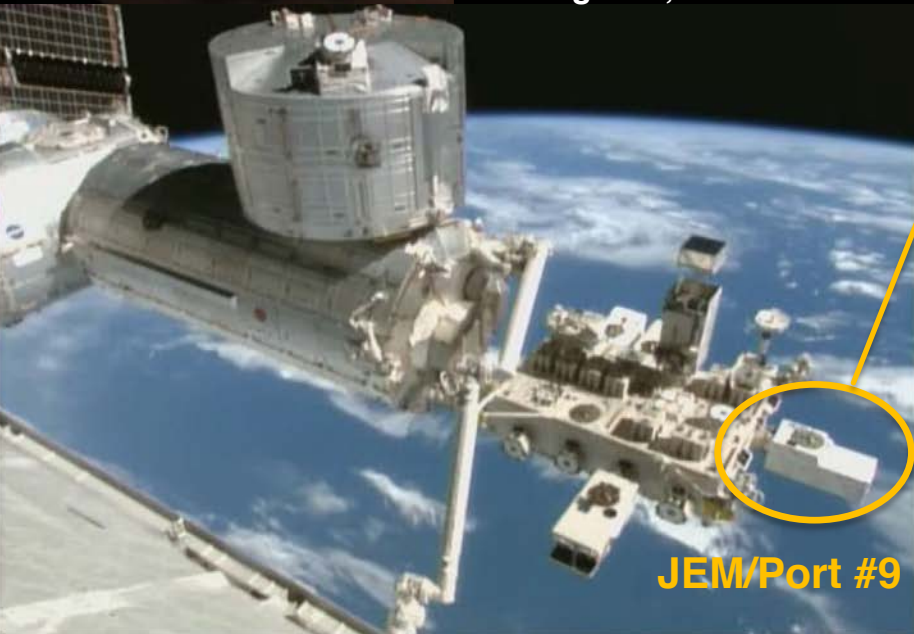


Kounotori (HTV) 5

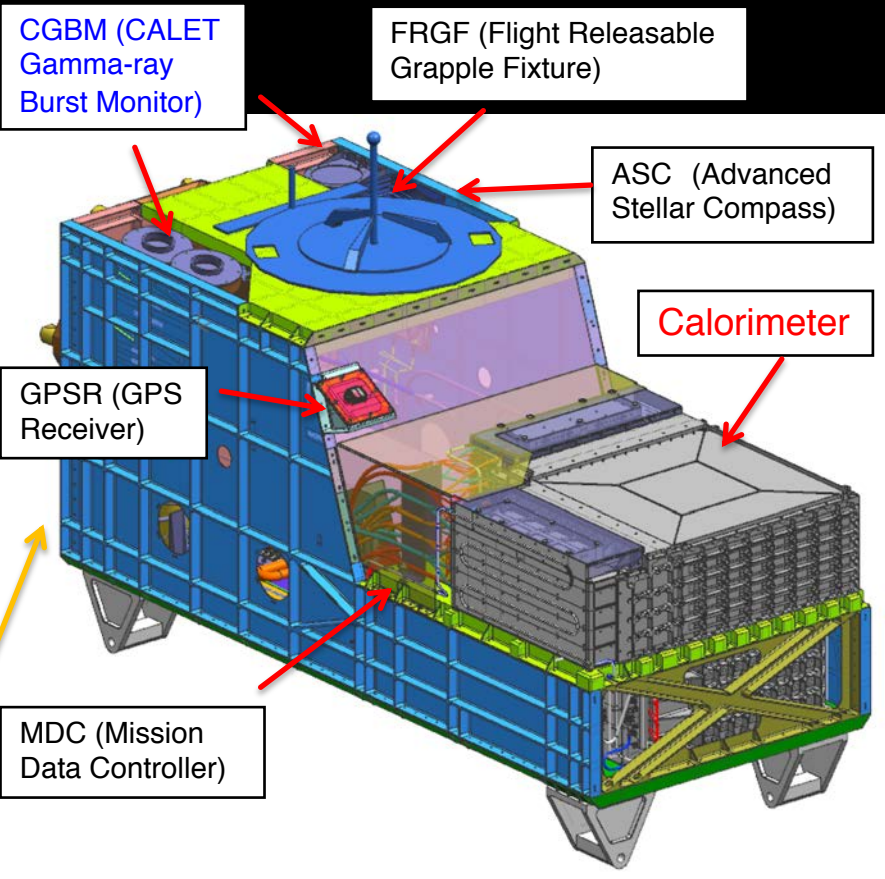


Launched on Aug. 19<sup>th</sup>, 2015  
by the Japanese H2-B rocket

Emplaced on JEM-EF port #9  
on Aug. 25<sup>th</sup>, 2015

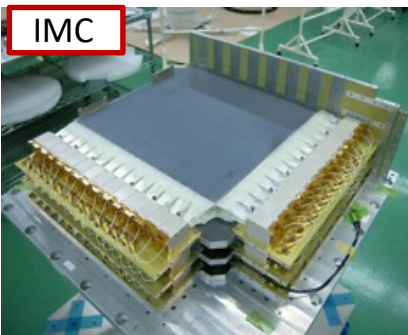
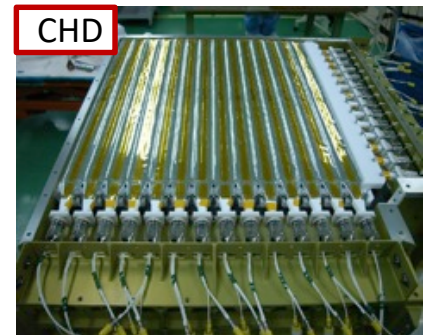
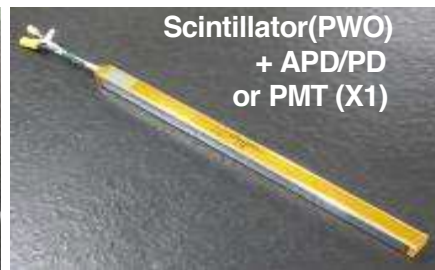


JEM/Port #9

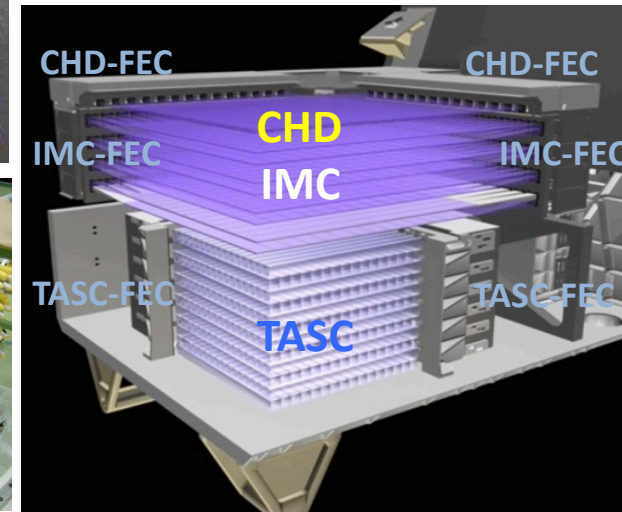


- Mass: 612.8 kg
- JEM Standard Payload Size:  
1850mm(L) × 800mm(W) × 1000mm(H)
- Power Consumption: 507 W (max)
- Telemetry:  
Medium 600 kbps (6.5GB/day) / Low 50 kbps

# CALET Instrument



## CALORIMETER



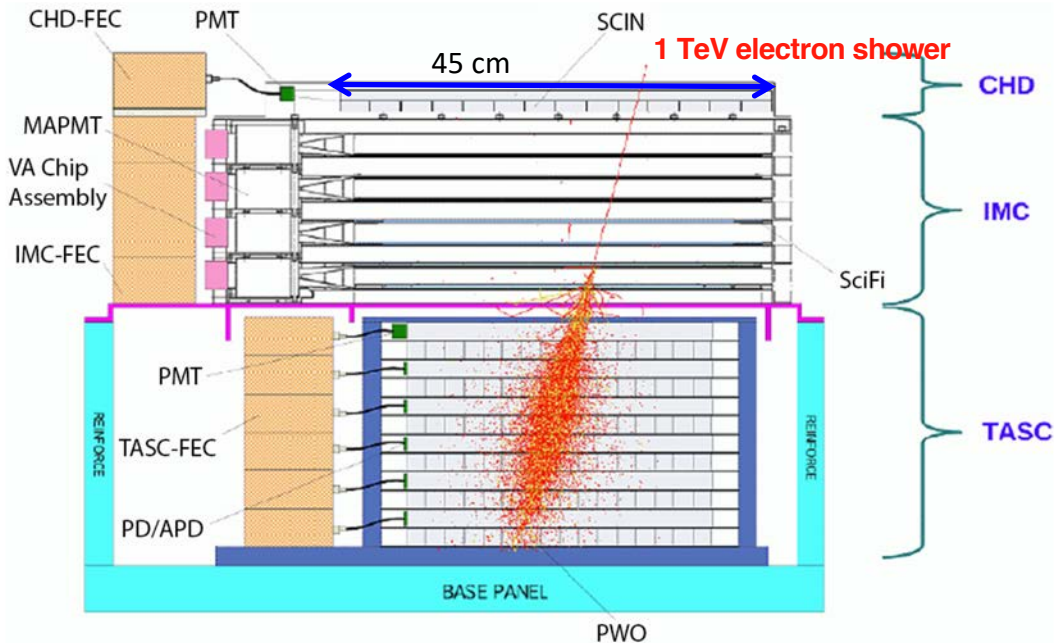
	CHD (Charge Detector)	IMC (Imaging Calorimeter)	TASC (Total Absorption Calorimeter)
Measure	Charge ( $Z=1-40$ )	Tracking , Particle ID	Energy, e/p Separation
Geometry (Material)	Plastic Scintillator 14 paddles x 2 layers (X,Y): 28 paddles Paddle Size: $32 \times 10 \times 450 \text{ mm}^3$	448 Scifi x 16 layers (X,Y) : 7168 Scifi 7 W layers ( $3X_0$ ): $0.2X_0 \times 5 + 1X_0 \times 2$ Scifi size : $1 \times 1 \times 448 \text{ mm}^3$	16 PWO logs x 12 layers (x,y): 192 logs log size: $19 \times 20 \times 326 \text{ mm}^3$ Total Thickness : $27 X_0$ , $\sim 1.2 \lambda_I$
Readout	PMT+CSA	64-anode PMT+ ASIC	APD/PD+CSA PMT+CSA (for Trigger)@top layer





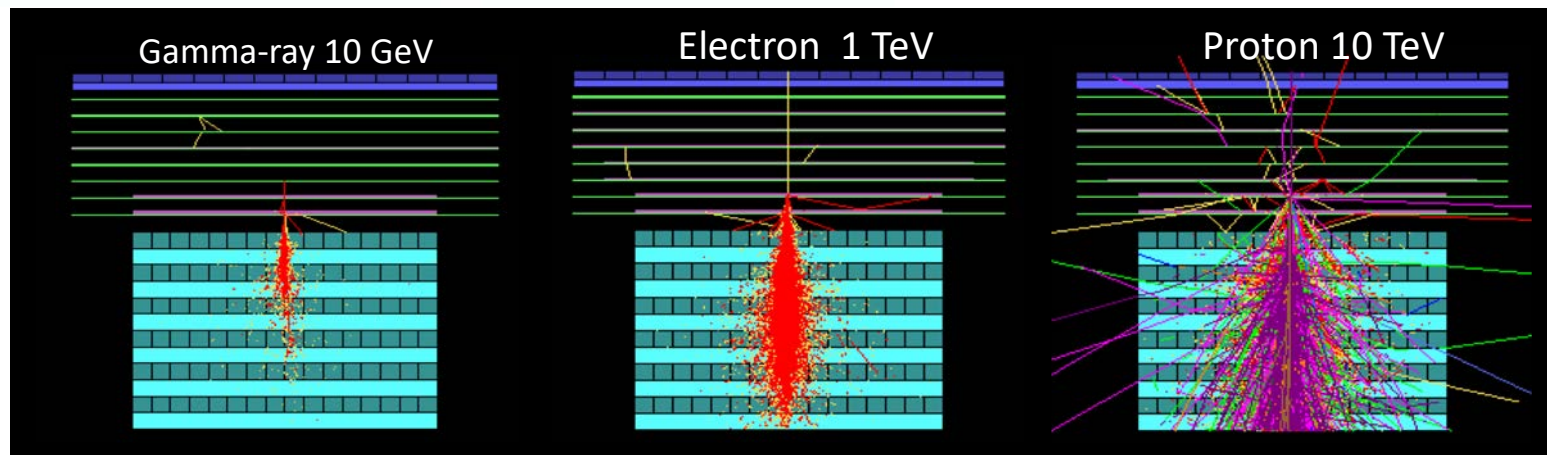
# CALET Capability

Field of view:  $\sim 45$  degrees (from the zenith)  
Geometrical Factor:  $\sim 1,040 \text{ cm}^2\text{sr}$  (for electrons)



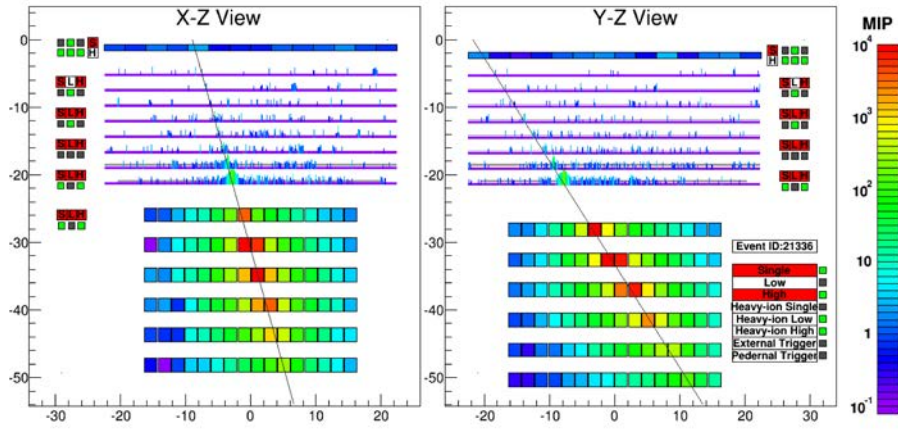
## Unique features of CALET

- ❑ A dedicated charge detector + multiple  $dE/dx$  track sampling in the IMC allow to identify individual nuclear species ( $\Delta z \sim 0.15-0.3 e$ ).
- ❑ Thick ( $\sim 30 X_0$ ), fully active calorimeter allows measurements well into the TeV energy region with excellent energy resolution ( $\sim 2-3\%$ )
- ❑ High granularity imaging pre-shower calorimeter accurately identify the arrival direction of incident particles ( $\sim 0.2^\circ$ ) and the starting point of electro-magnetic showers.
- Combined, they powerfully separate electrons from the abundant protons: contamination is much less than 10 % up to the TeV region.

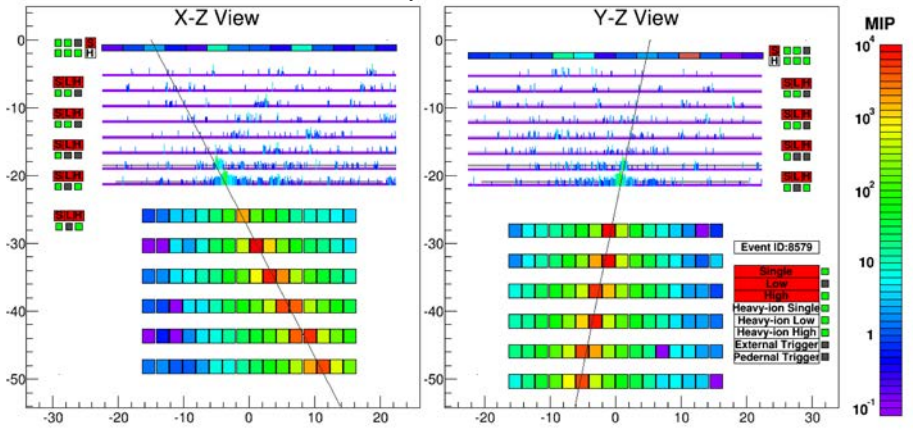


# Examples of Event Display

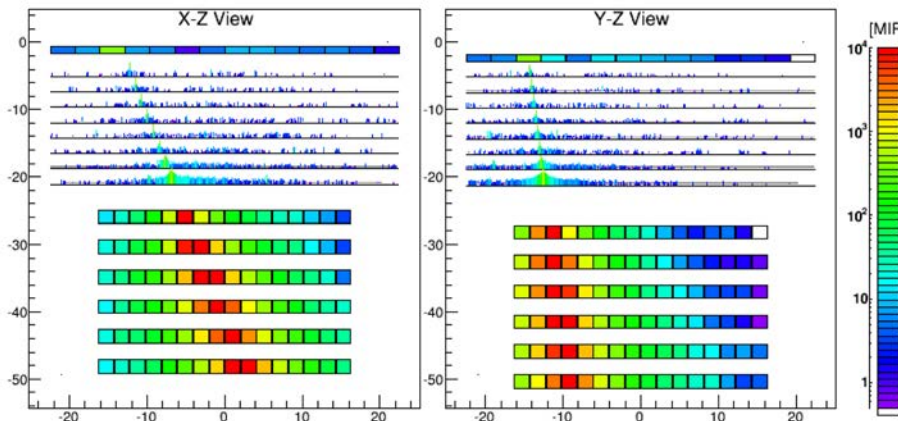
Electron,  $E=3.05$  TeV



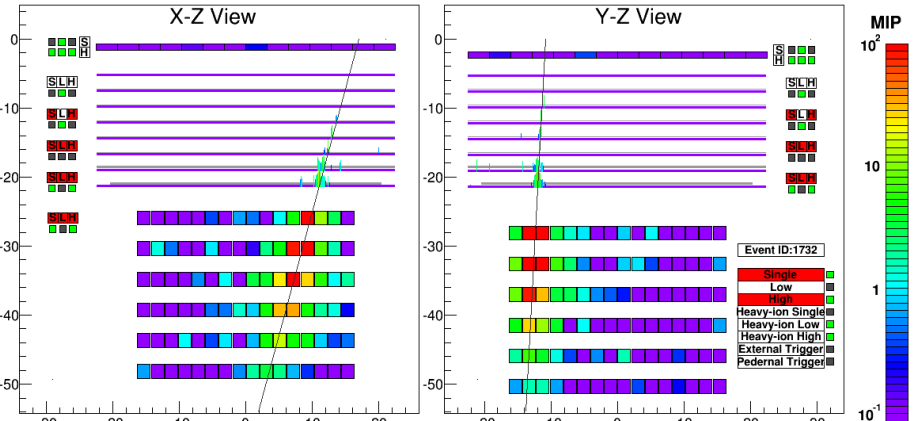
Proton,  $\Delta E=2.89$  TeV



Fe,  $\Delta E=9.3$  TeV

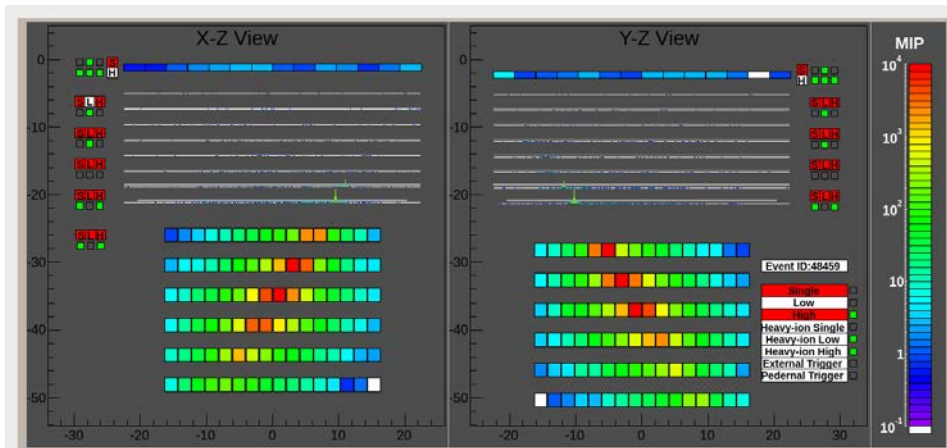
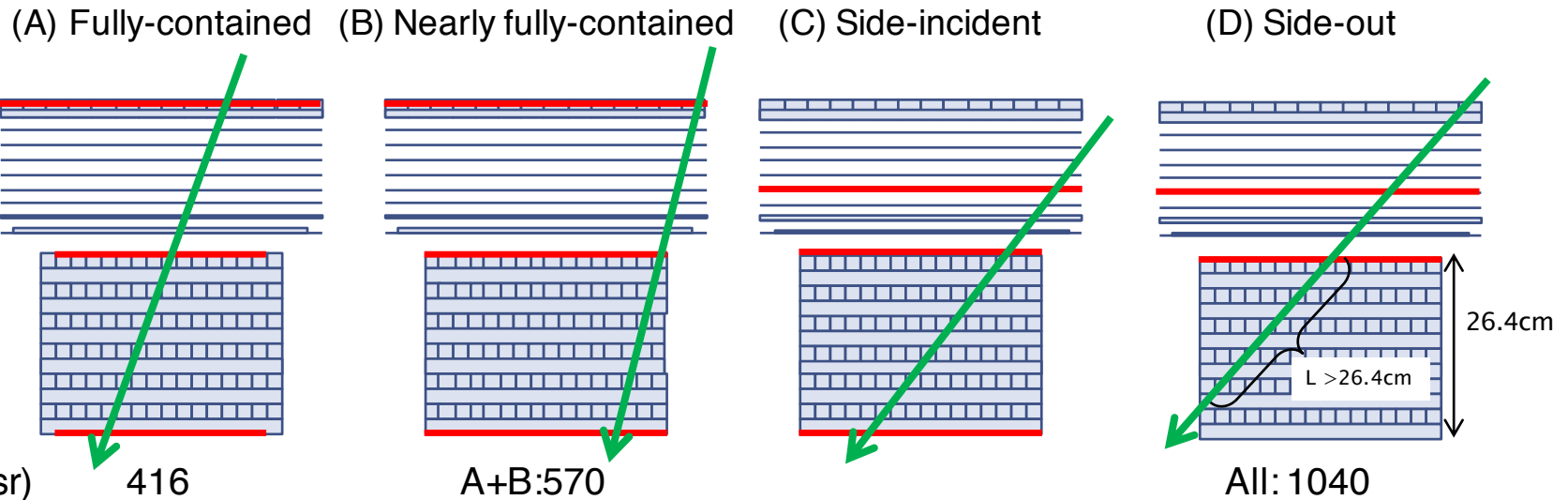


Gamma-ray,  $E=44.3$  GeV

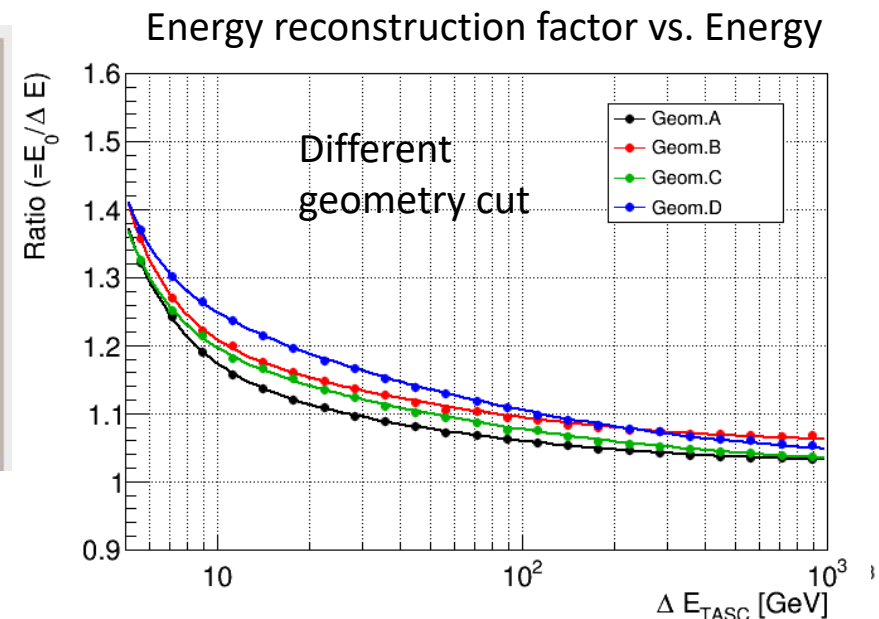


Unit in MIP

# Energy Reconstruction for Electromagnetic Showers



4 TeV electron candidate (well contained)  
 $\Rightarrow$  very small leakage ( $\sim$  a few %)





# Electron Identification

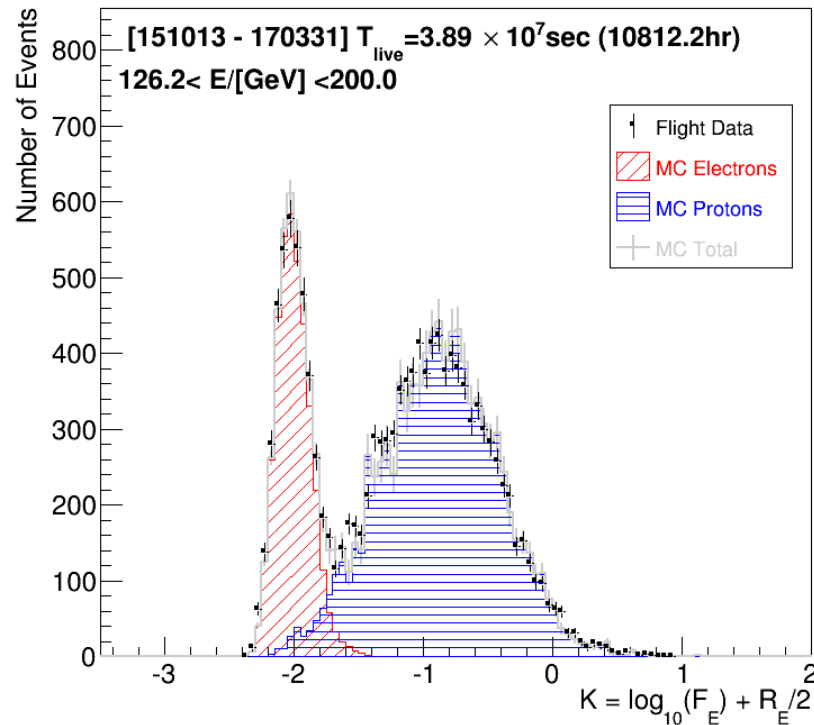
## Simple Two Parameter Cut

$F_E$ : Energy fraction of the bottom layer sum to the whole energy deposit sum in TASC

$R_E$ : Lateral spread of energy deposit in TASC-X1

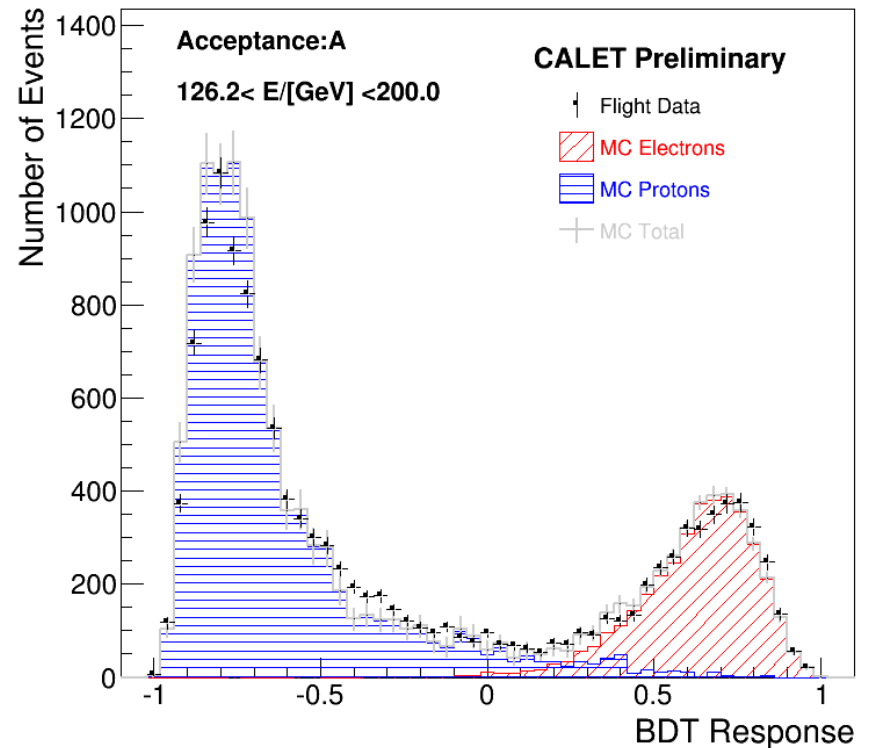
Cut Parameter K is defined as follows:

$$K = \log_{10}(F_E) + 0.5 R_E \text{ (/cm)}$$



## Boosted Decision Trees (BDT)

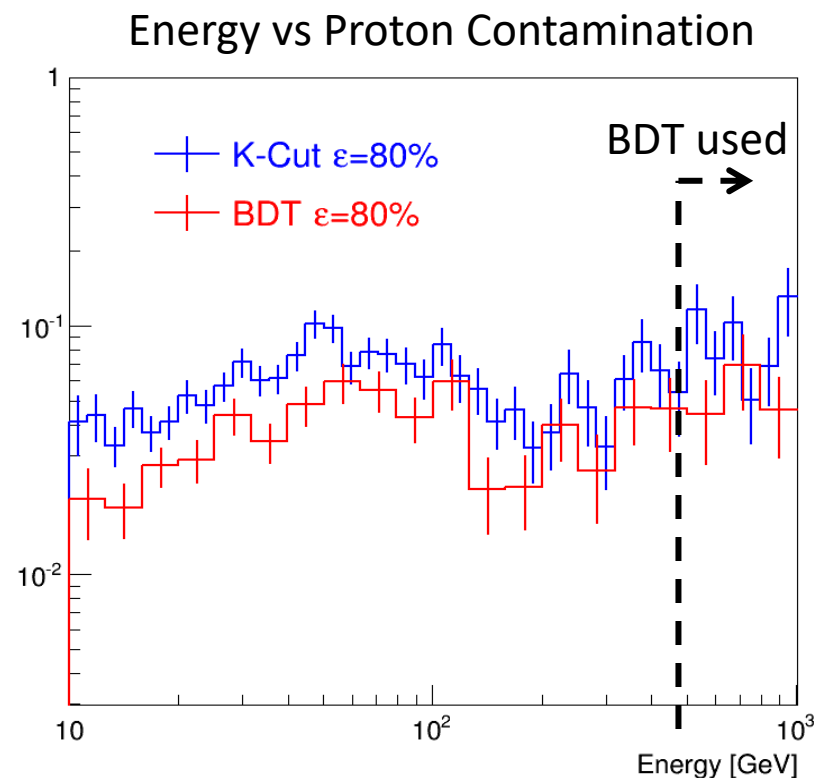
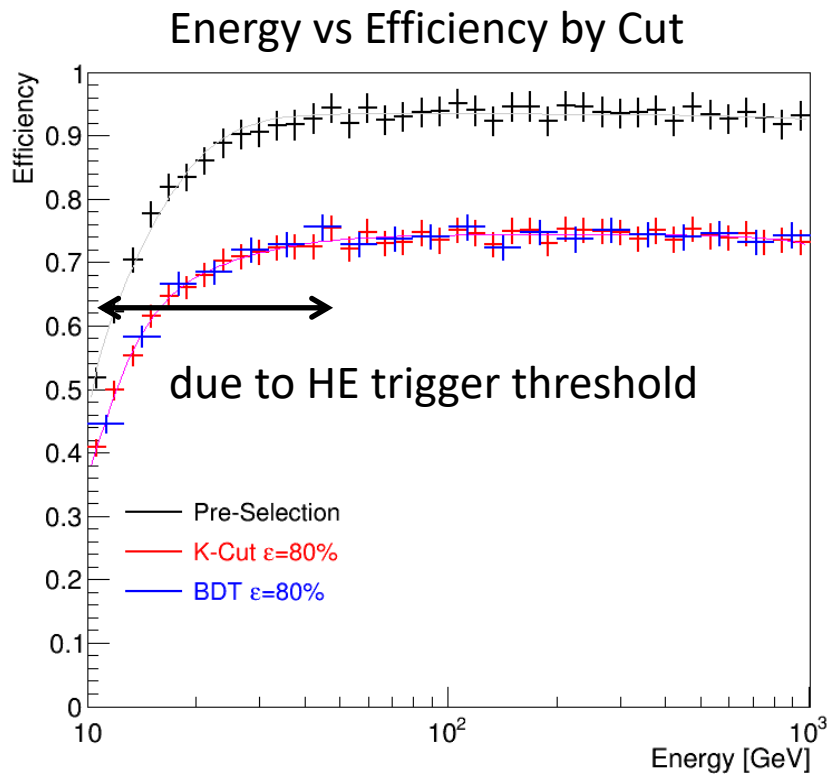
In addition to the two parameters in the left, TASC and IMC shower profile fits are used as discriminating variables.





# Electron Efficiency and Subtraction of Proton Contamination

- Constant and high efficiency is the key point in our analysis.
- Simple two parameter cut is used in the low energy region while the difference in resultant spectrum are taken into account in the systematic uncertainty.



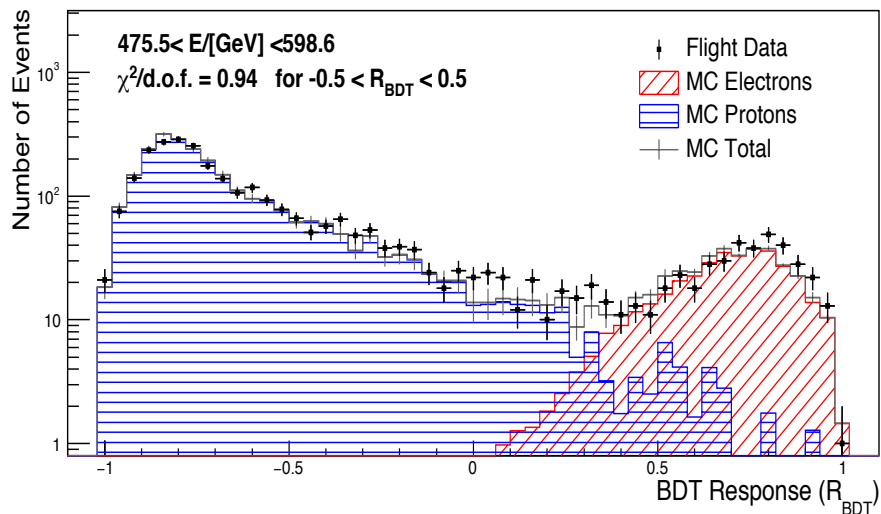


# BDT Response Distribution at Higher Energies

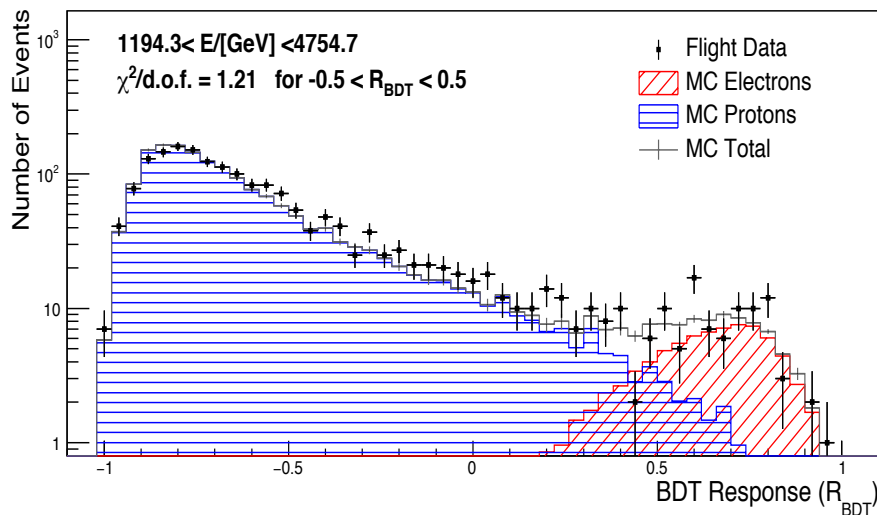
In the final electron sample, the resultant contamination ratios of protons are:

**5 % up to 1 TeV ; 10% - 20% in the 1 - 4.8 TeV region**  
 , while keeping a constant high efficiency of 80 % for electrons.

$476 < E < 599 \text{ GeV}$



$1196 < E < 4755 \text{ GeV}$   
 (highest energy bin)

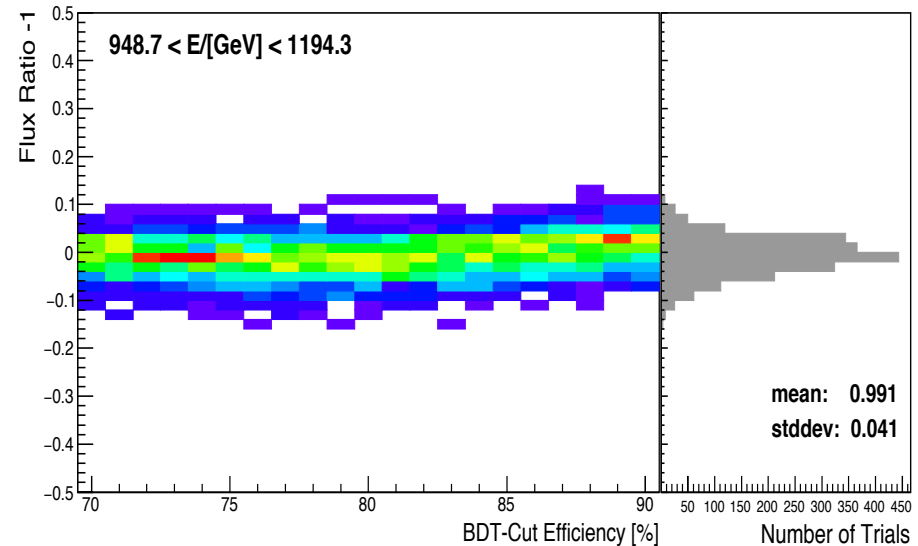




# Stability of BDT Analysis and Energy Dependence of Systematic Uncertainties

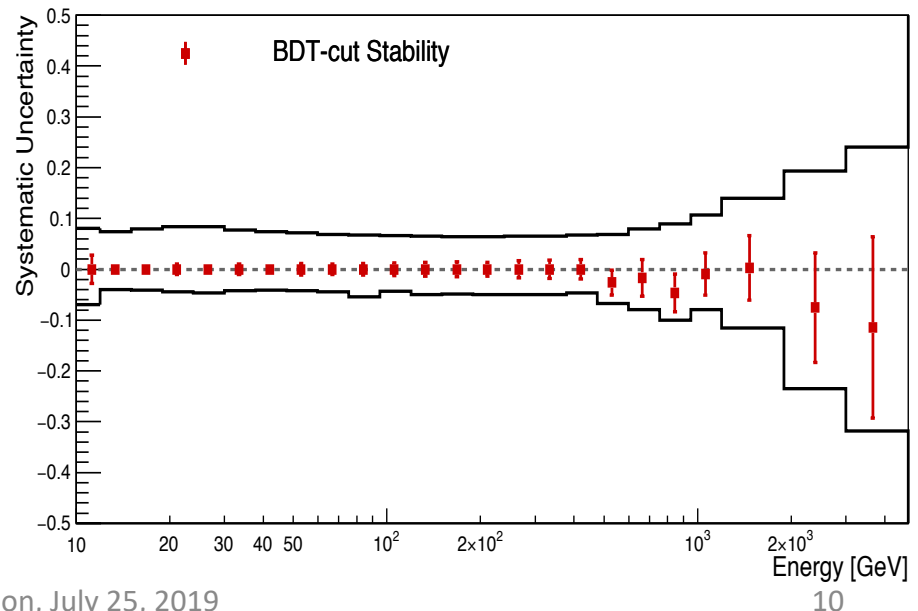
Stability of BDT analysis with respect to independent training samples and BDT-cut efficiency in the  $949 < E < 1194$  GeV

- Color maps show the flux ratio dependence on efficiency, where the bin value (number of trials) increases as color changes from violet, blue, green, yellow to red.
- A projection onto the Y-axis is shown as a rotated histogram (in gray color).



## Energy dependence of systematic uncertainties

- The red squares represent the systematic uncertainties stemming from the electron identification based on BDT.
- The bands defined by black lines show the sum in quadrature of all the sources of systematics, except the energy scale uncertainties.

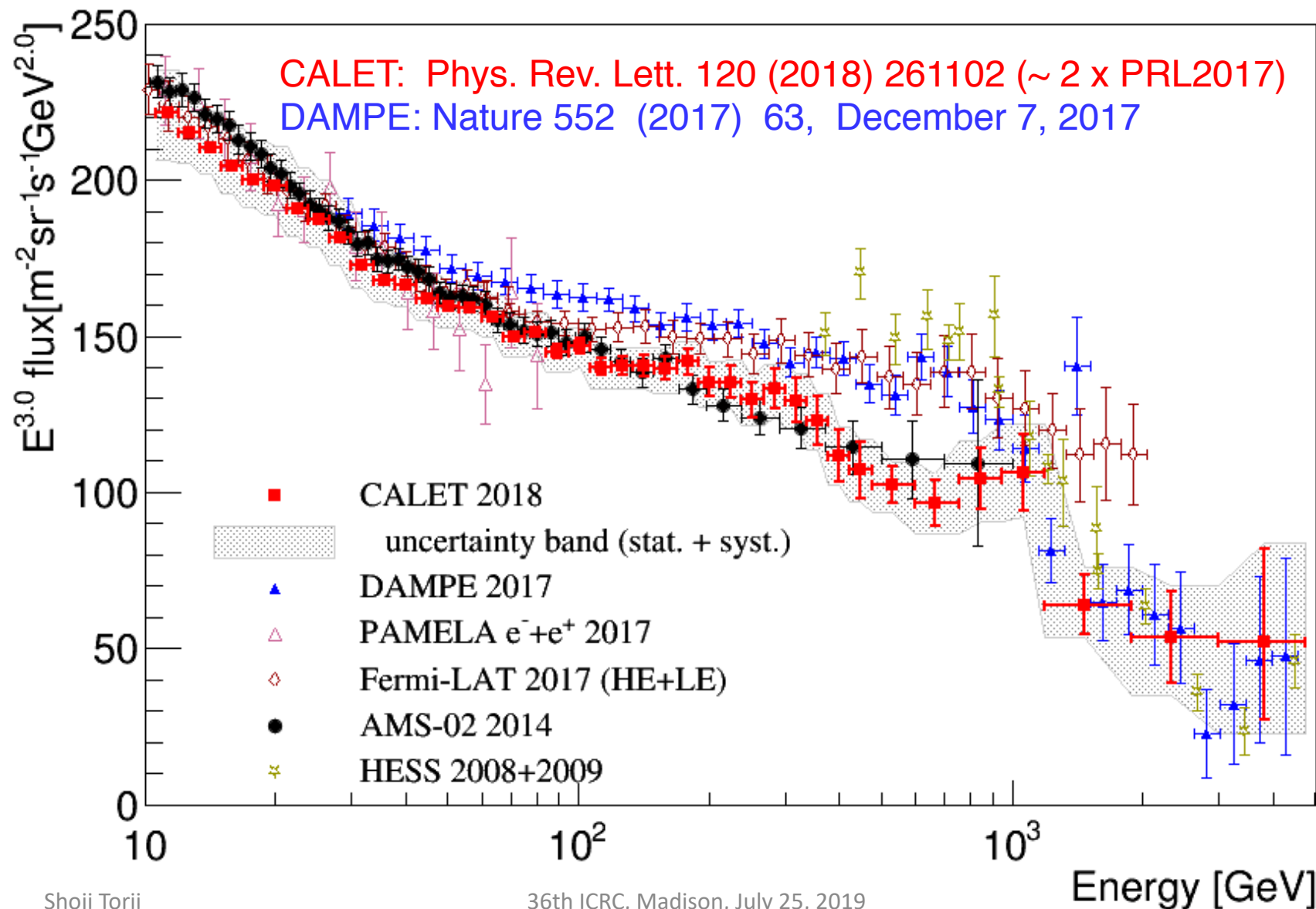






# Extended Measurement by Observation over 780 days

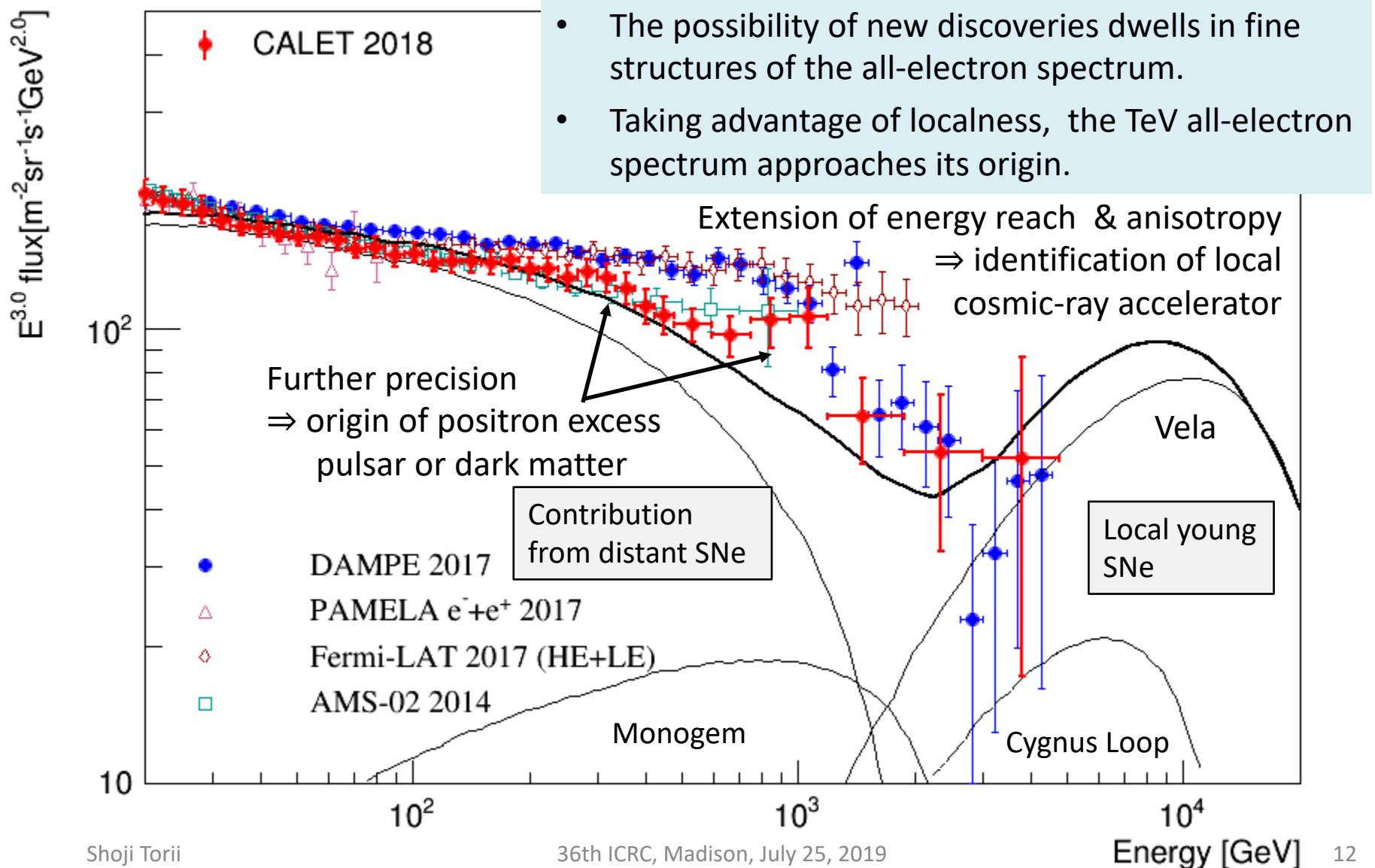
Approximately doubled statistics above 500GeV by using full acceptance of CALET





# Prospects for the CALET All-Electron Spectrum

Five years or more observations  $\Rightarrow$  3 times more statistics, reduction of systematic errors







# Summary and Future Prospects

- ❑ CALET was successfully launched on August 19th, 2015, and is successfully carrying out observations since October 2015 with stable instrument performance.
- ❑ As of the end of May, 2019, the exposure amount, SQT, reached more than  $110 \text{ m}^2 \text{ sr day}$  for electron observations over 10 GeV.
- ❑ We have reported results of the all-electron ( $e^+ + e^-$ ) spectrum in the energy range from 10 GeV to 4.8 TeV.
- ❑ Further observations will improve the measurement of electron spectrum by better statistics and a further reduction of the systematic errors, especially in the TeV region.
- ❑ We gratefully acknowledge JAXA's contribution to the development of CALET and to the operation on the ISS.