

# Measurements of Heavy Cosmic-Ray Nuclei Spectra from CALET on the ISS

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# CALET Collaboration Team



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# Measurements of Cosmic-Ray Nuclei

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Nuclei measurements in GeV – TeV energy region

- Primary individual spectra
  - cosmic-ray acceleration and propagation
  - hardening of spectra
- Secondary-to-primary flux ratio
  - cosmic-ray propagation
  - energy dependence of diffusion coefficient

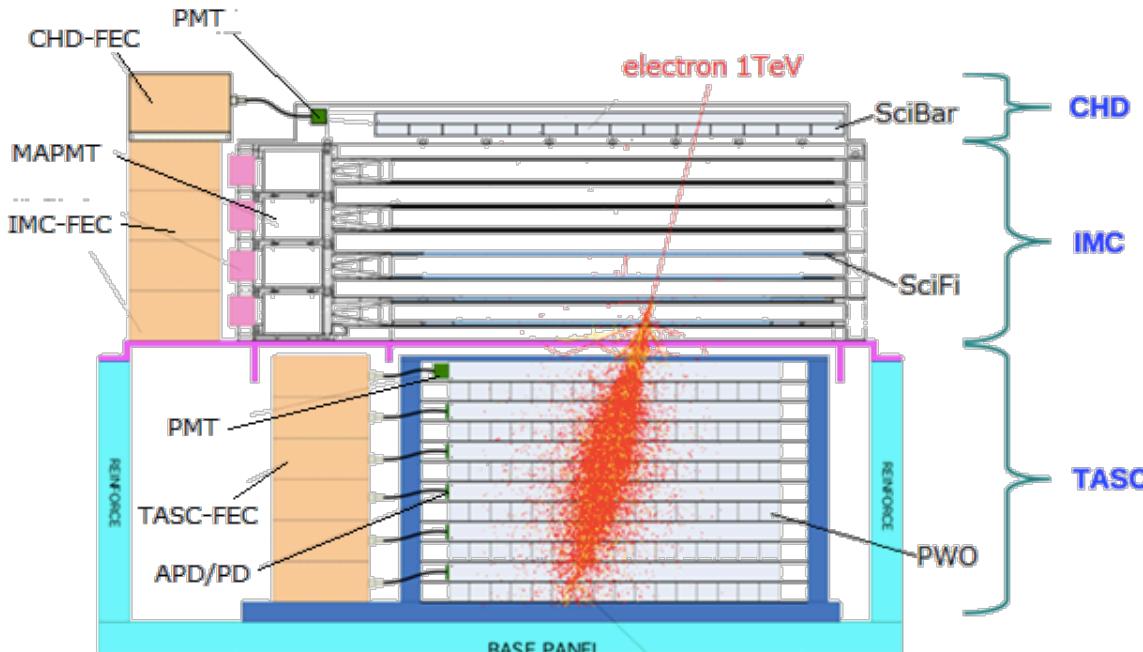
Direct measurements with CALET

## Energy spectra from Proton to Iron

- Charge measurement in  $Z = 1 - 40$ 
  - charge resolution:  $0.15e(C)-0.3e(Fe)$
- Energy measurement in  $10\text{GeV} - 1\text{PeV}$ 
  - dynamic range :  $1 - 10^6\text{MIP} (\sim 1\text{PeV})$

# CALET-CAL Detector

A 30 radiation length deep calorimeter designed to detect electrons and gamma-rays to 20 TeV and cosmic rays up to 1 PeV



## Performances

- Energy resolution
  - 2% ( $>100\text{GeV}$ ) for electron,  $\gamma$
  - 30~35% for protons, nuclei
- Charge resolution
  - 0.15 – 0.3e
- Angular resolution
  - 0.2 deg for electron,  $\gamma > \sim 50\text{GeV}$

	<b>CHD (Charge Detector)</b>	<b>IMC (Imaging Calorimeter)</b>	<b>TASC (Total Absorption Calorimeter)</b>
Function	Charge Measurement ( $Z=1-40$ )	Arrival Direction, Particle ID	Energy Measurement, Particle ID
Sensor (+ Absorber)	<b>Plastic Scintillator : 14 × 2 layers</b> Unit Size: 32mm x 10mm x 450mm	<b>Scintillating fibers: 448 x 16 layers</b> Unit size: 1mm <sup>2</sup> x 448 mm <b>Total thickness of Tungsten: 3 <math>X_0</math></b>	<b>PWO log: 16 x 12 layers</b> Unit size: 19mm x 20mm x 326mm <b>Total Thickness of PWO: 27 <math>X_0</math></b>
Readout	PMT+CSA	64 -anode MAPMT + ASIC	APD/PD+CSA PMT+CSA (for Trigger)@top layer

# Analysis procedure for nuclei

## 1. HE (High Energy) trigger

- Period: Oct. 13 2015 - Oct. 31 2017 (750days)

## 2. Offline shower trigger

## 3. Tracking with IMC

- select events satisfied Geom.A
- identify the impact point

## 4. Charge consistency with CHD and IMC

- remove backgrounds
- maintain charge resolution

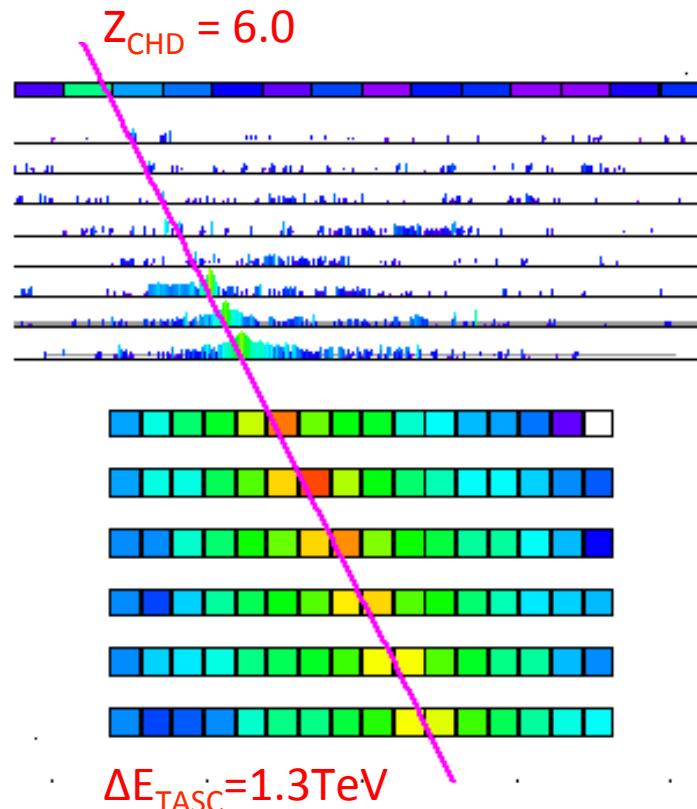
## 5. Charge selection with CHD

- estimate background

## 6. Energy measurements and unfolding

- measure energy with TASC
- unfold energy spectrum by Iterative Bayesian process

## 7. Flux Calculation

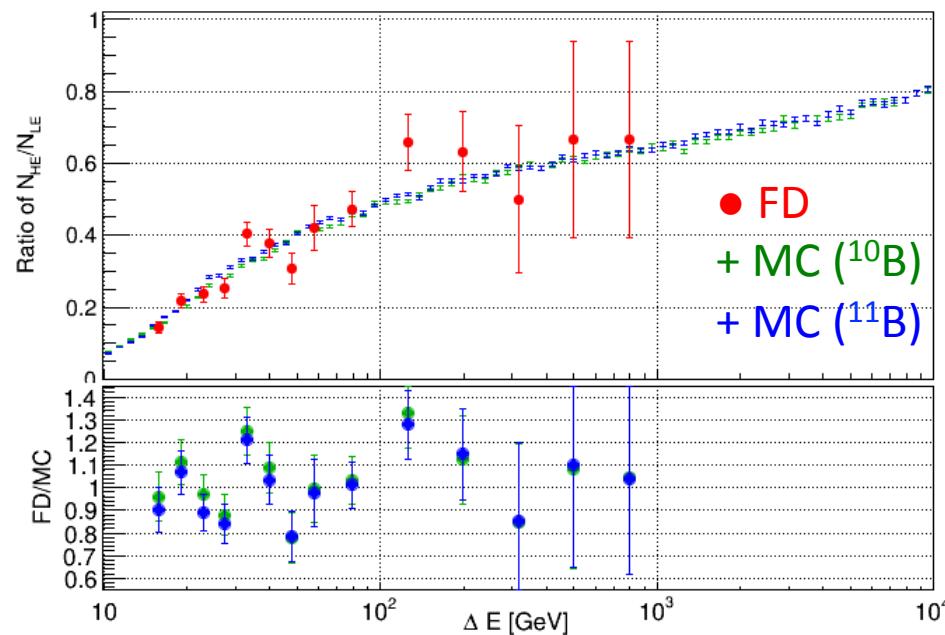


# Trigger efficiency of HE trigger

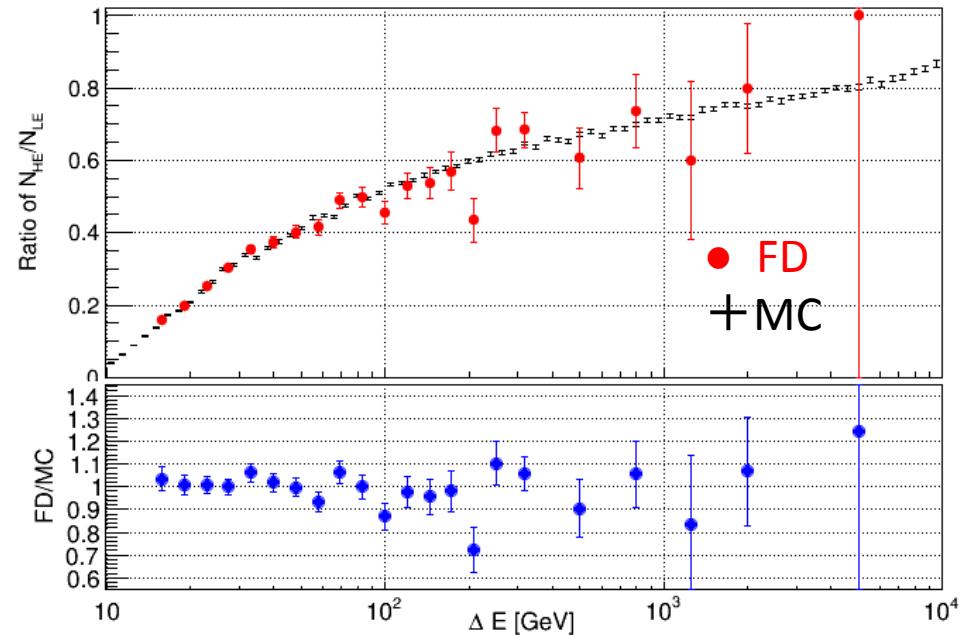
HE trigger is a shower trigger to detect 10 GeV shower events

The HE trigger efficiency was evaluated using data by **low threshold trigger** which can detect all events of  $Z \geq 5$  including penetrating particles because of the large  $dE/dx$

HE Trigger Efficiency for Boron



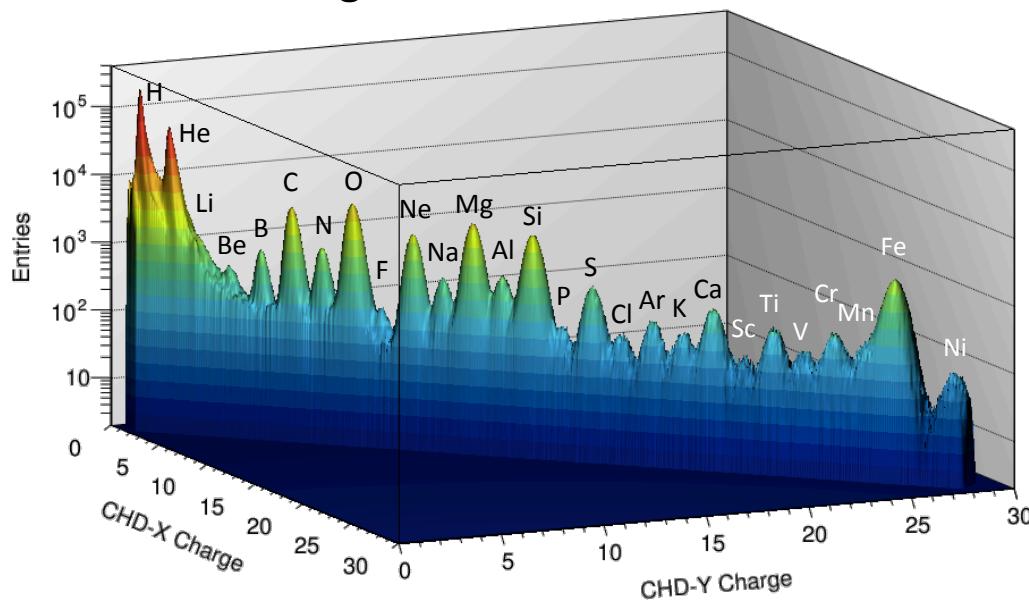
HE Trigger Efficiency for Carbon



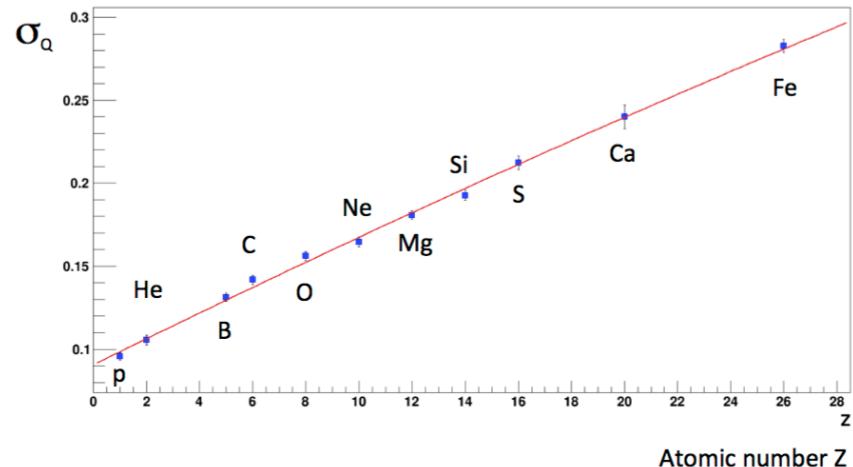
# Charge resolution

- Non-linear response to  $Z^2$  is corrected both in CHD and IMC using a model
- A clear separation between p, He,  $\sim Z=8$ , can be seen from CHD+IMC data analysis

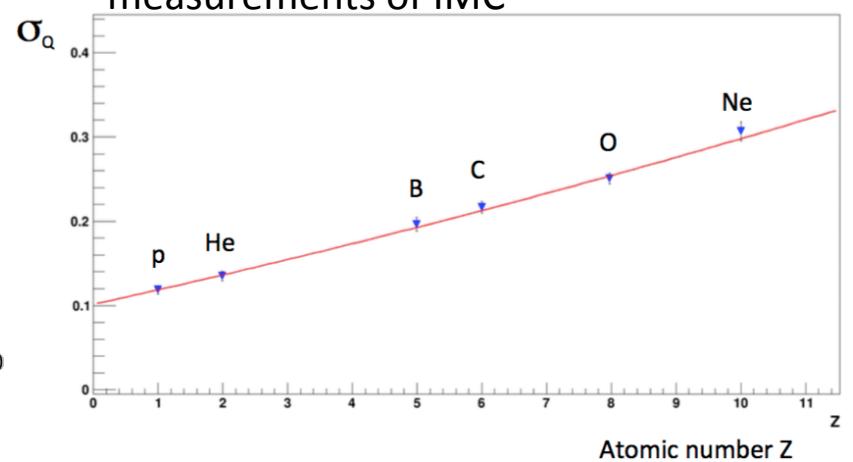
Charge with CHD-X vs CHD-Y



Charge resolution with 2 layers of CHD



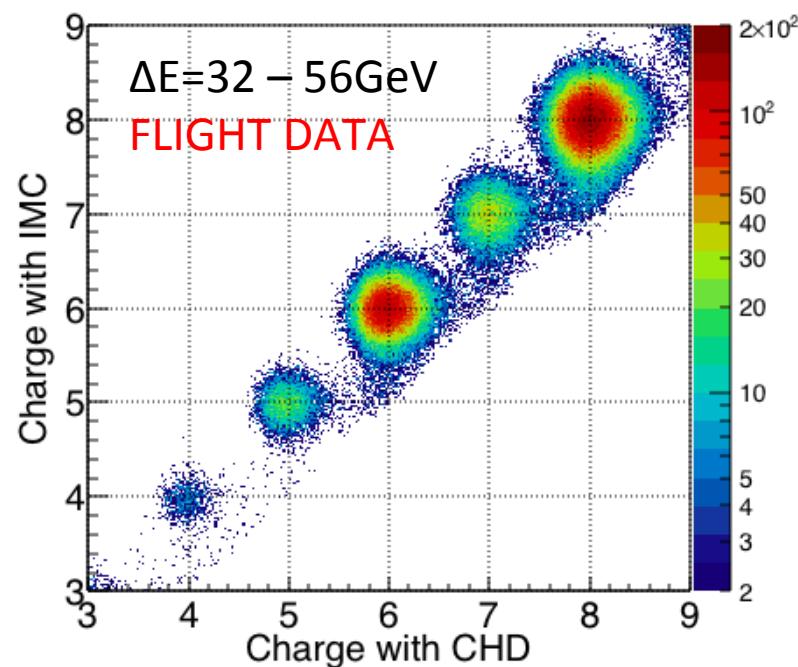
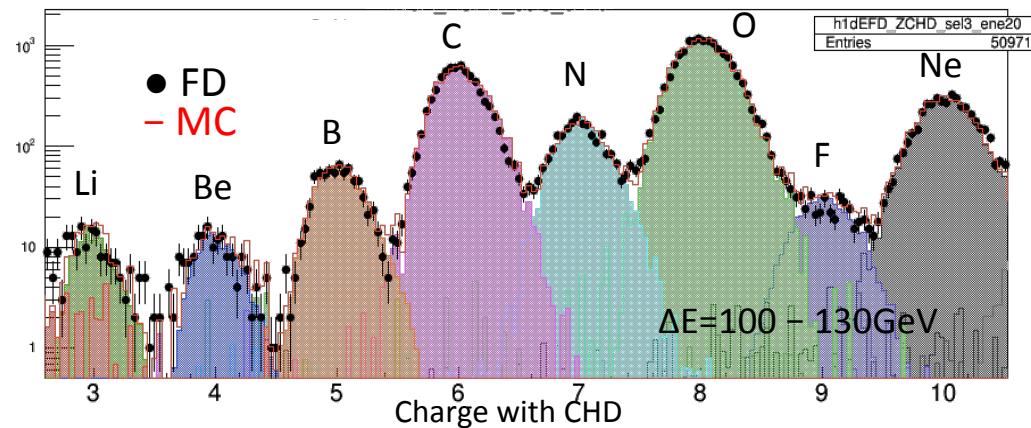
Charge resolution using multiple dE/dx measurements of IMC



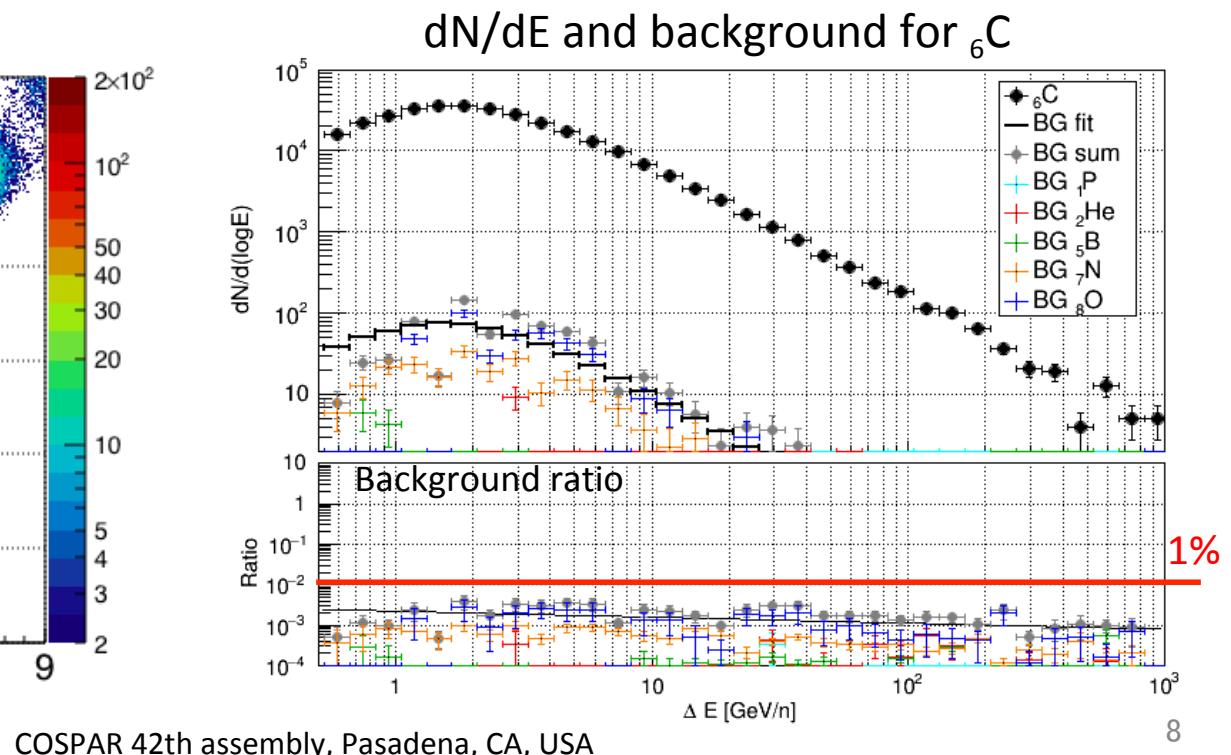
# Charge identification of Carbon

## Pre-selection

- HE trigger
- Tracking + geometrical condition
- Z-consistency with CHD-X and CHD-Y
- Z-consistency with CHD and IMC



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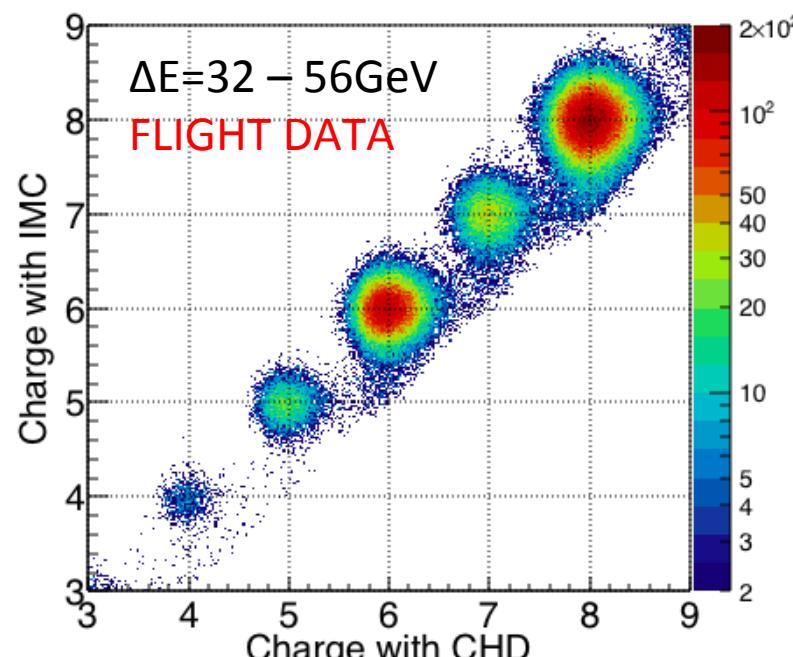
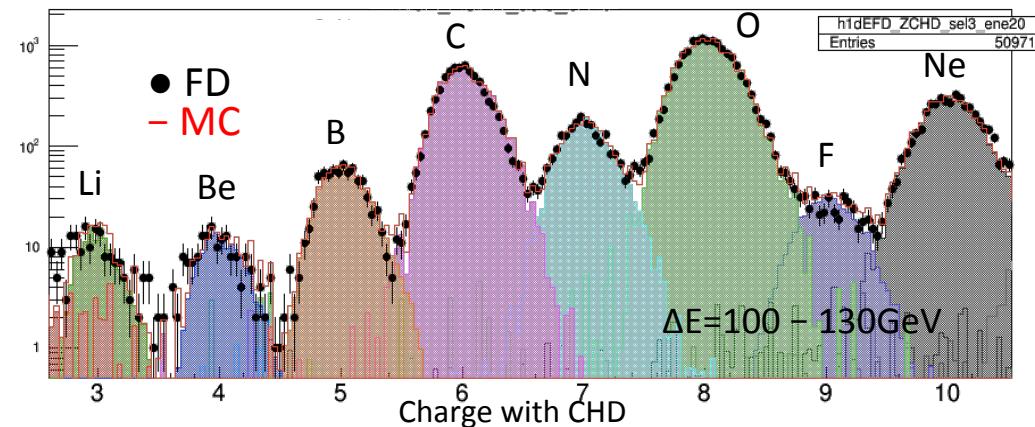


COSPAR 42th assembly, Pasadena, CA, USA

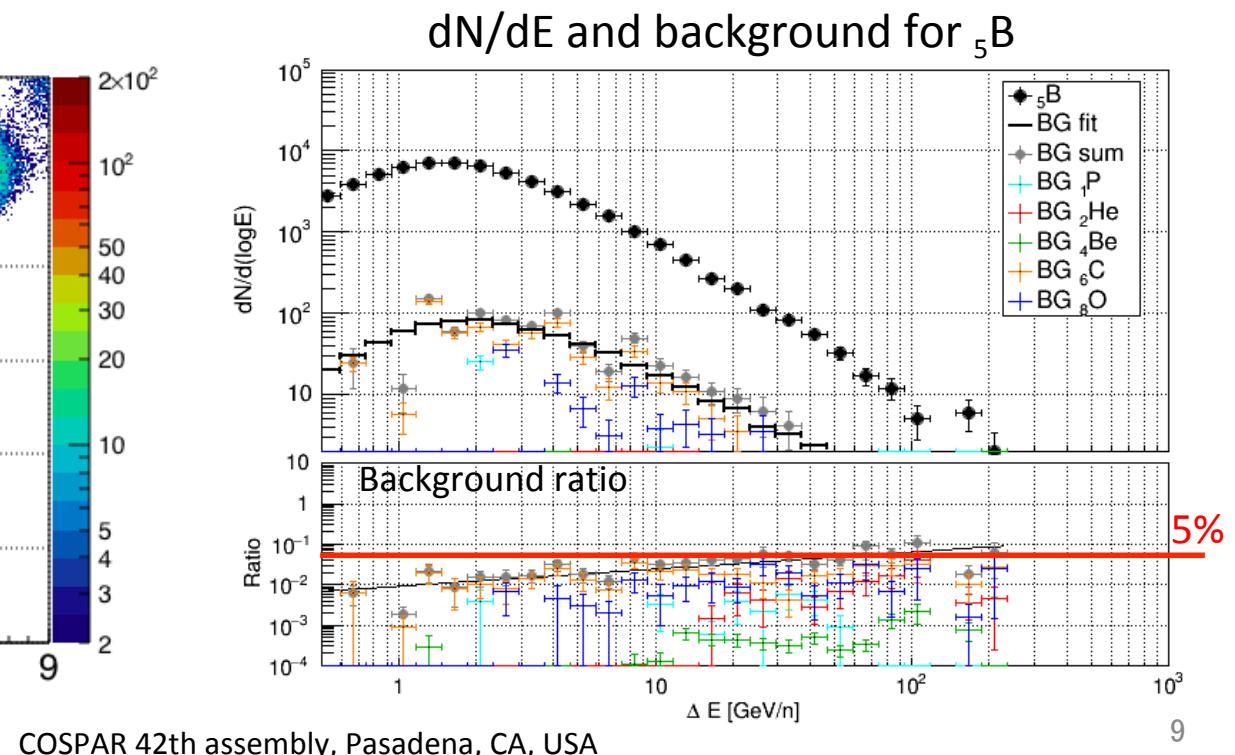
# Charge identification of Boron

## Pre-selection

- HE trigger
- Tracking + geometrical condition
- Z-consistency with CHD-X and CHD-Y
- Z-consistency with CHD and IMC

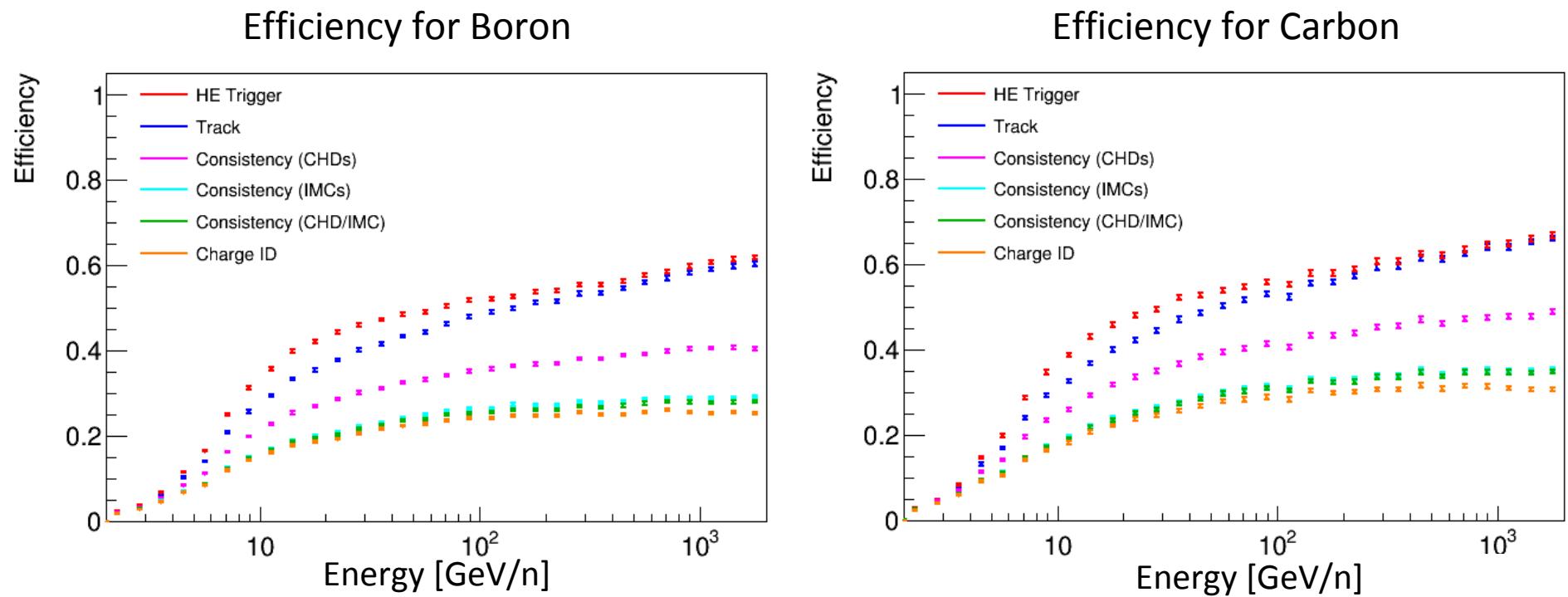


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# Efficiency



# Energy unfolding

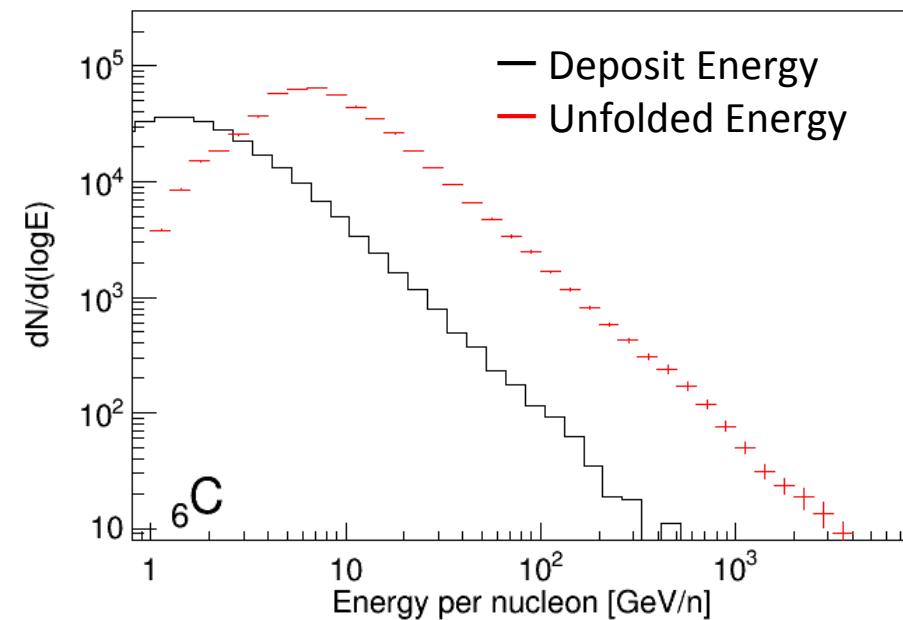
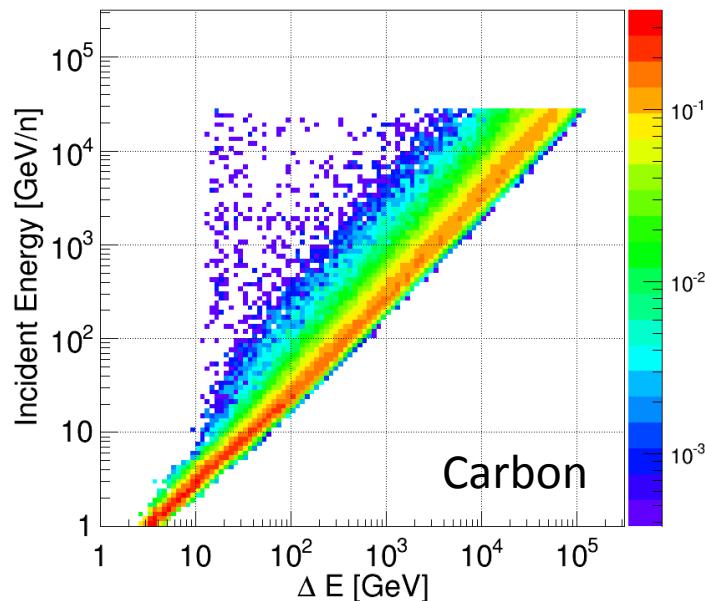
Characteristics of nuclei measurements with CALET calorimeter:

- thickness:  $30 X_0$  for electron,  $1.3\lambda$  for proton
- $\sigma(E)/E$  : 2% for electron, 30% for nuclei
- Need energy unfolding for nuclei to obtain primary energy spectrum

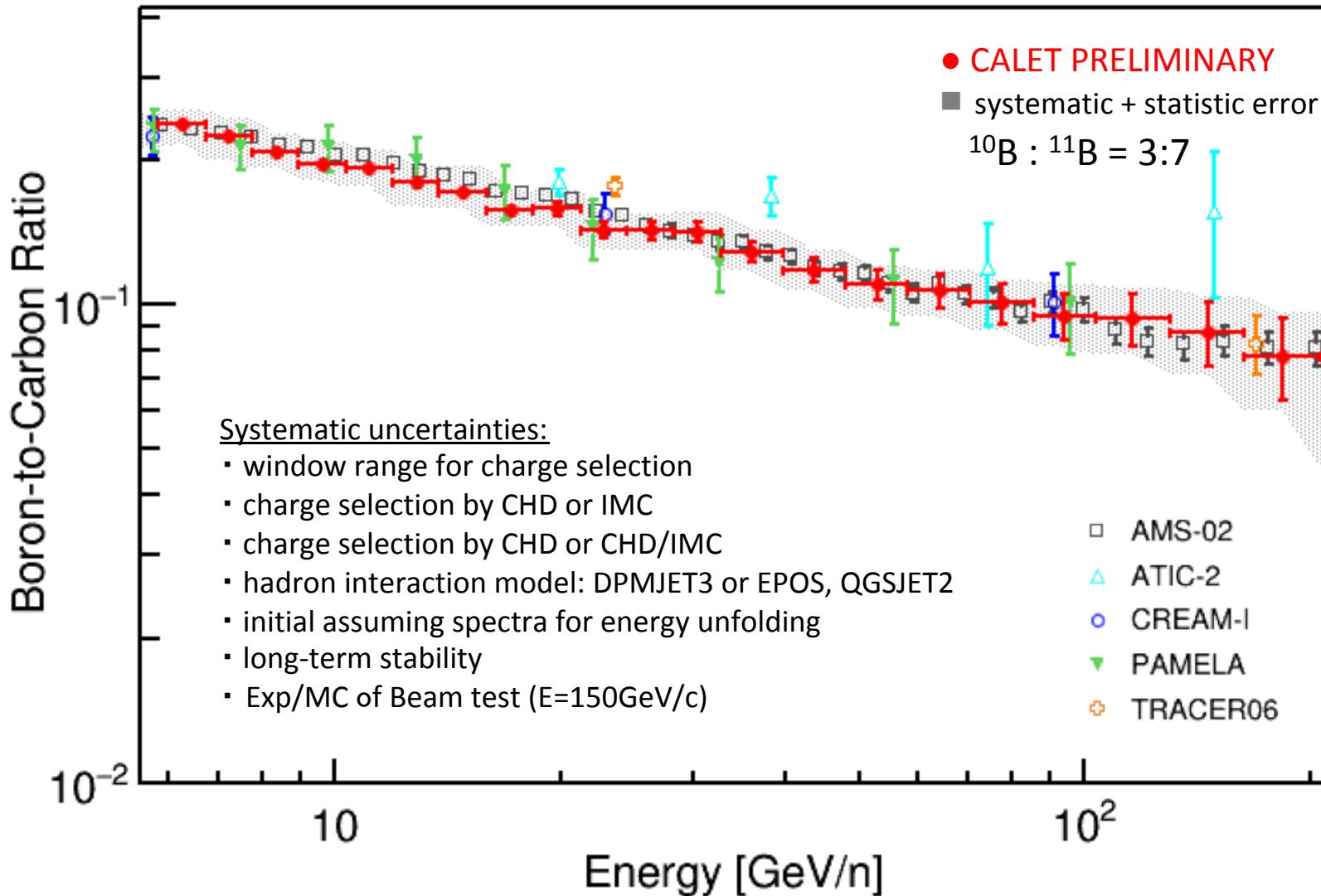
- Iterative Bayesian unfolding

- Initial assuming spectra:  $f(E) = A \times E^{-2.60}$   
A is normalized by charge distribution in CHD
- Response function:

$\Delta E$  [GeV] (deposit energy in calorimeter) vs  $E_0$  [GeV] (primary energy)

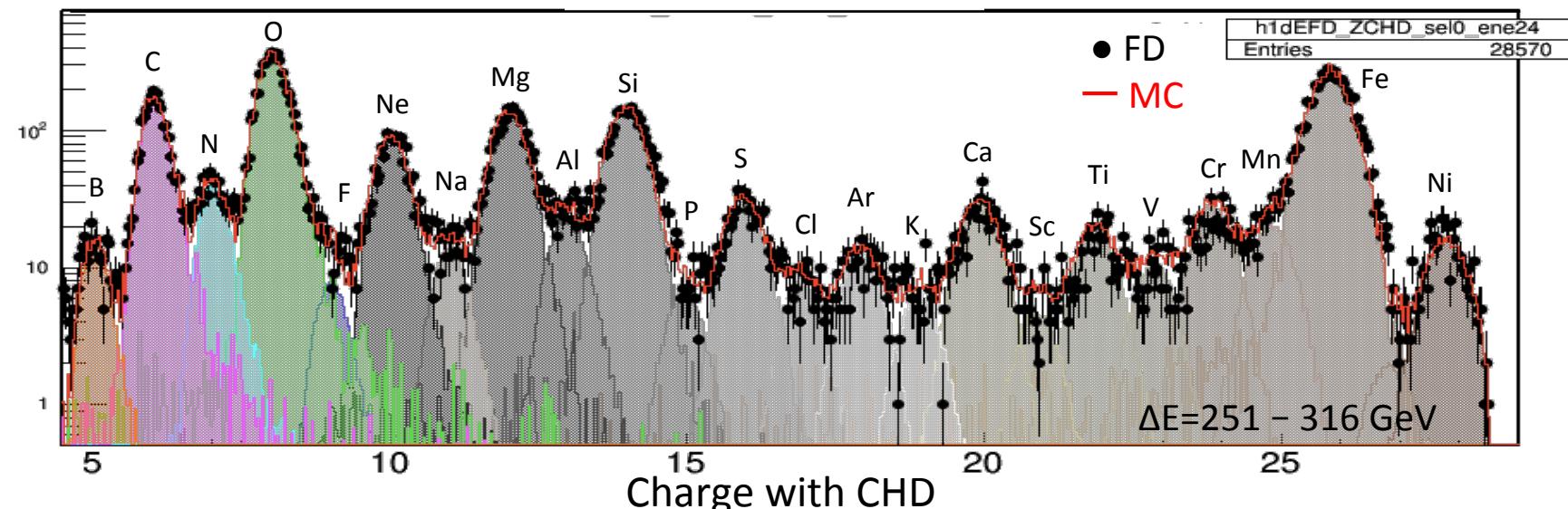
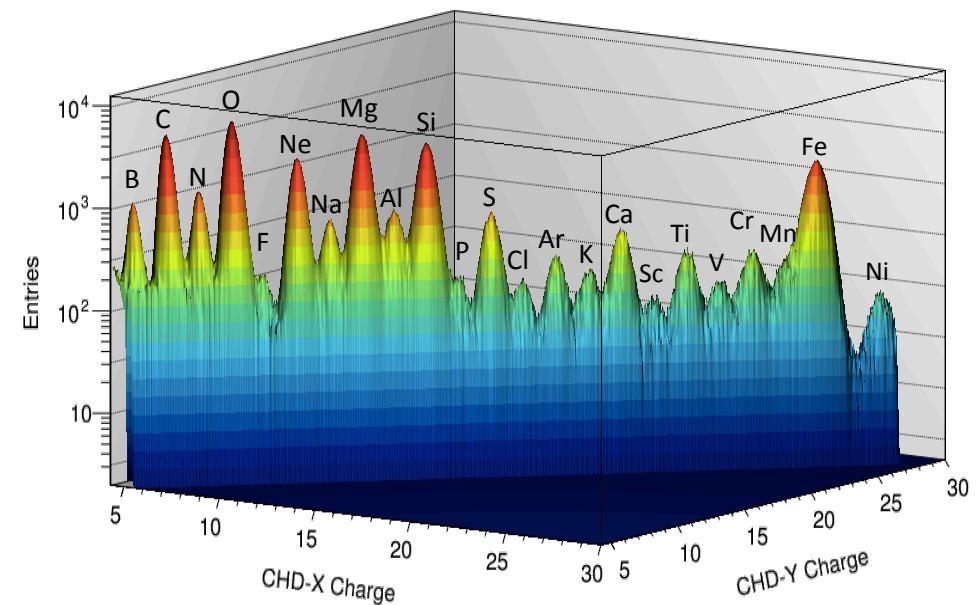


# Preliminary Boron-to-Carbon Flux Ratio

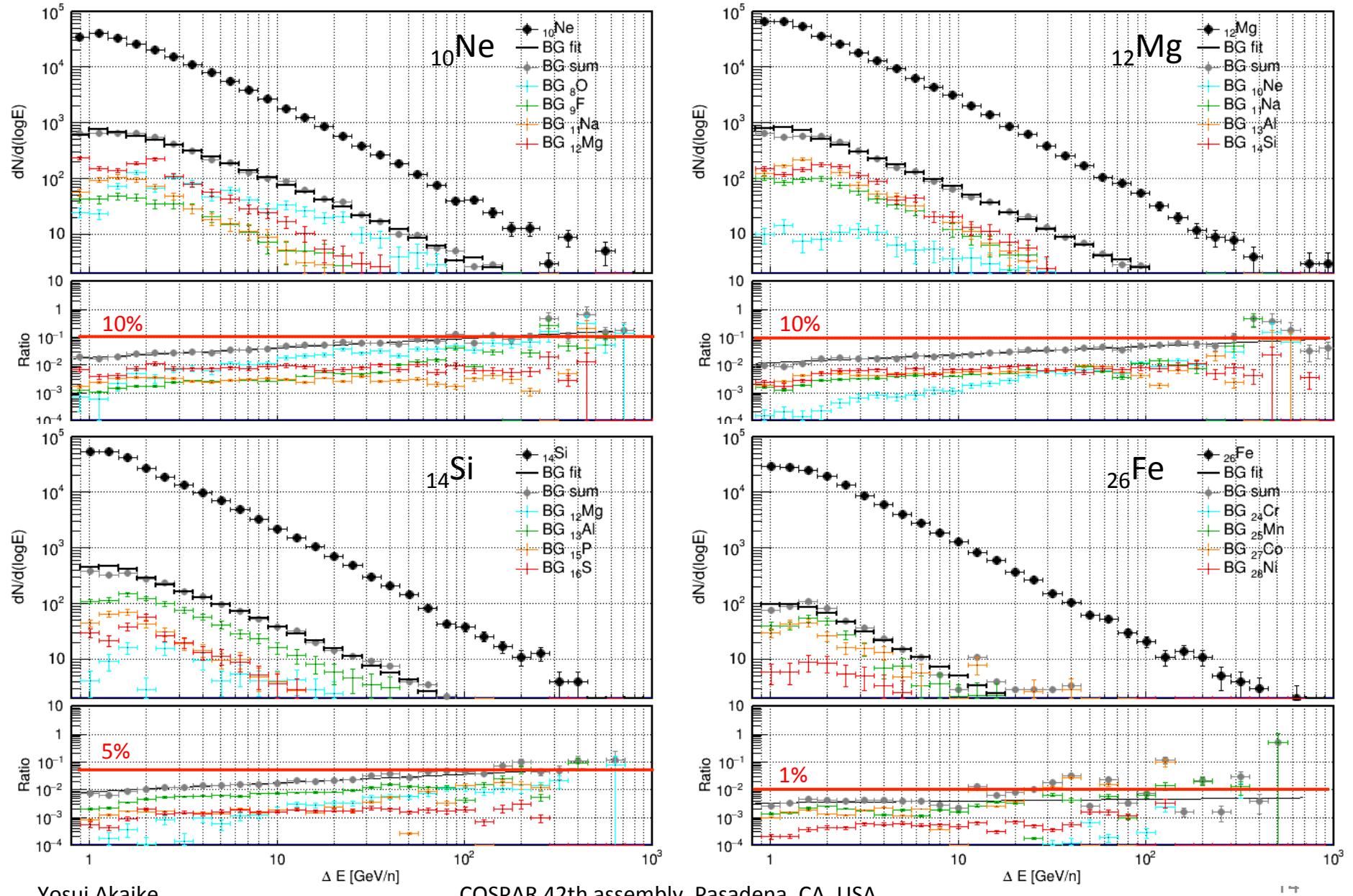


# Analysis of heavy nuclei ( $Z>8$ )

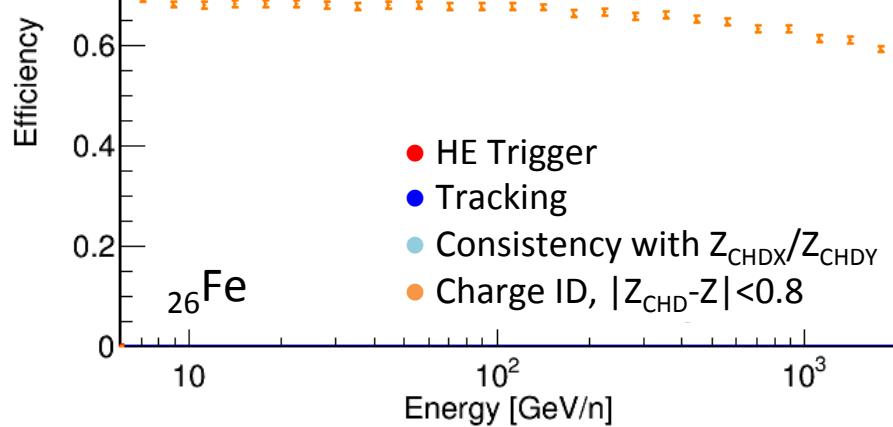
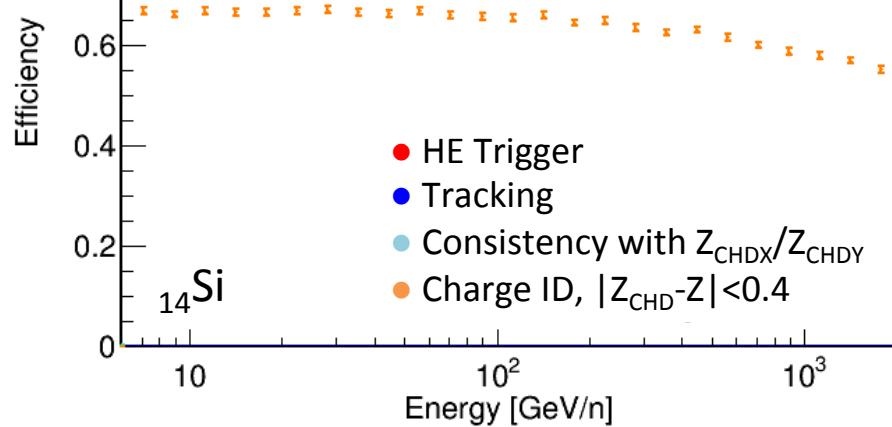
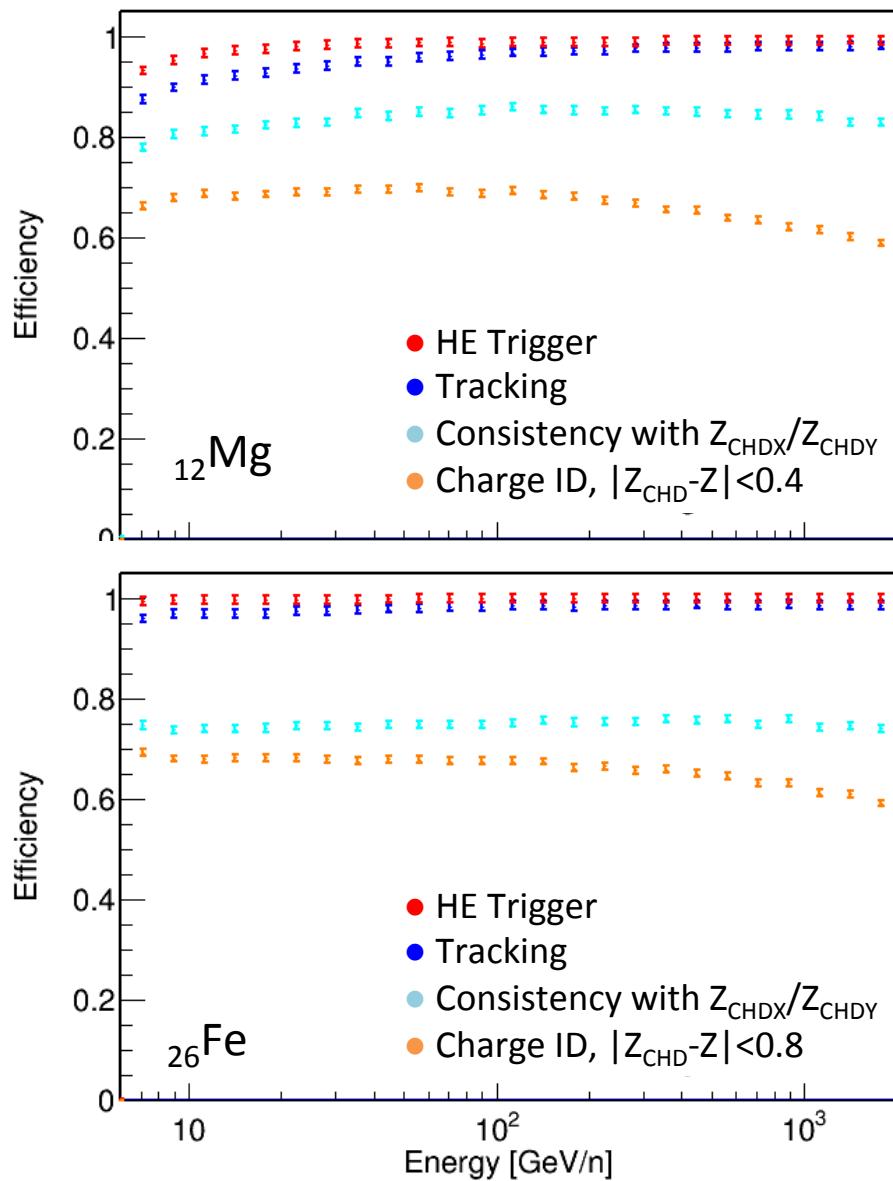
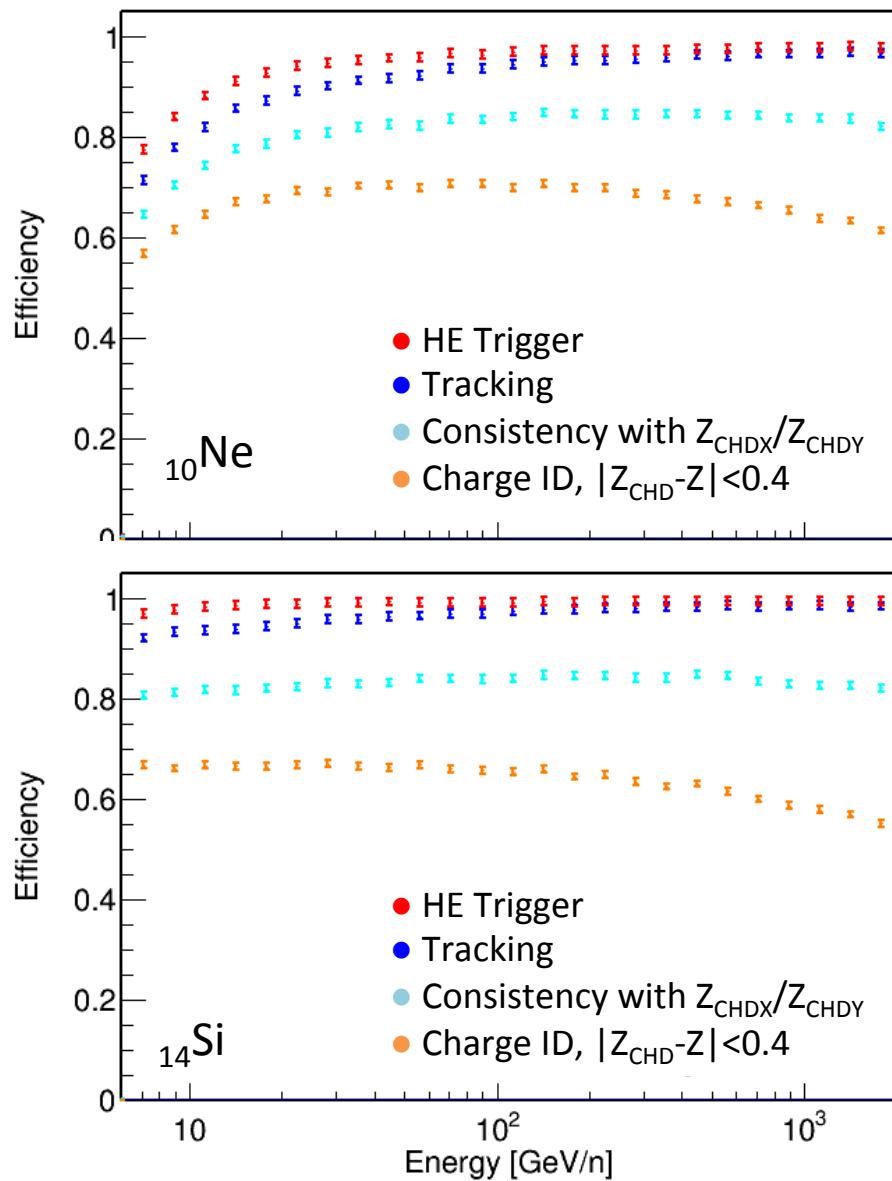
- Characteristics of heavy nuclei analysis
  - trigger efficiency :  $\sim 100\%$
  - particle charge is identified by CHD



# dN/dE and Background Estimation



# Efficiency



# Preliminary Flux of Primary Components

Flux measurements:

$$\Phi(E) = \frac{N(E)}{S\Omega\varepsilon(E)T\Delta E}$$

$N(E)$  : Events in unfolded energy bin

$S\Omega$  : Geometrical acceptance

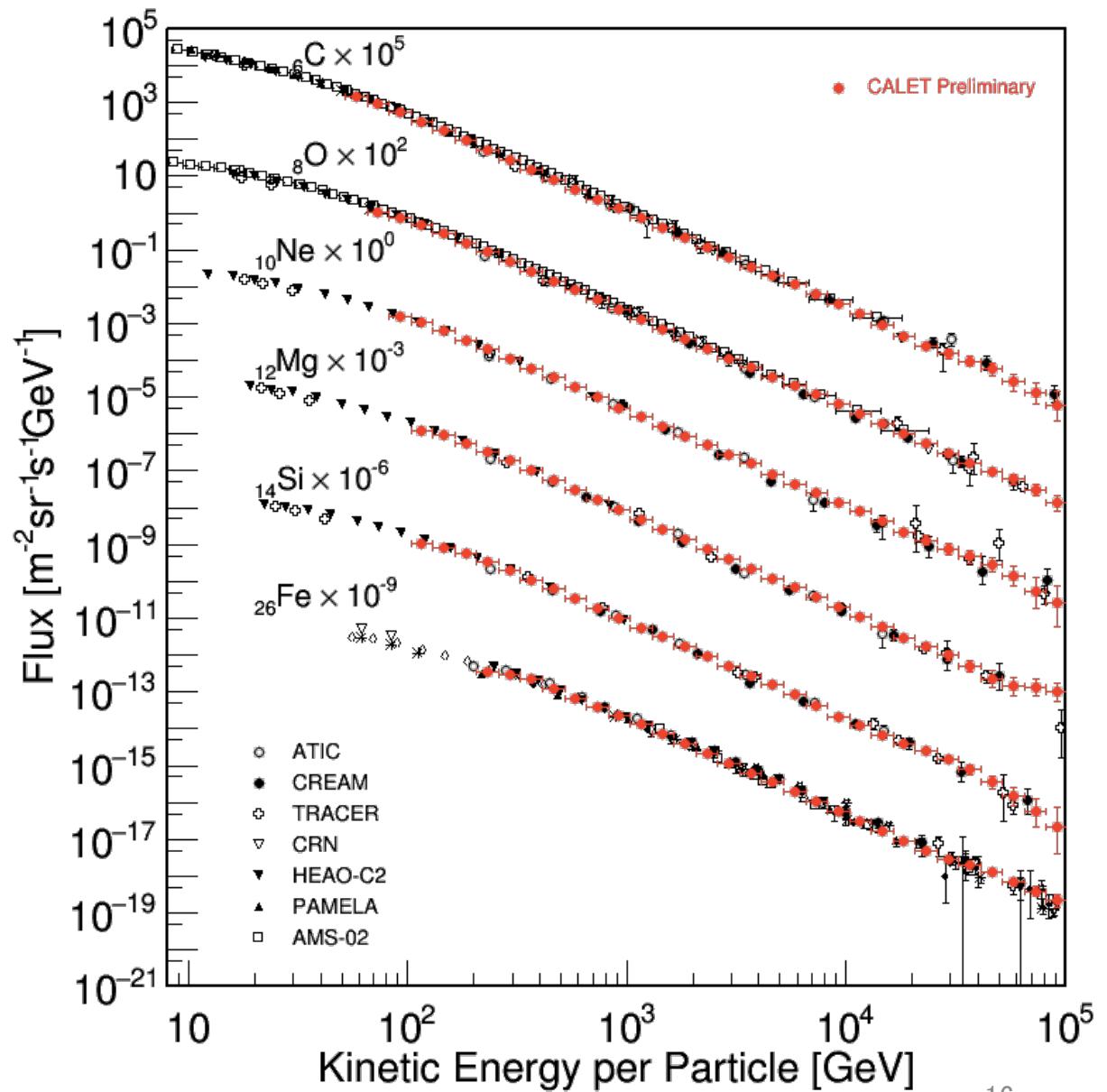
$\varepsilon(E)$  : Efficiency

$T$  : Live Time

$\Delta E$  : Energy bin width

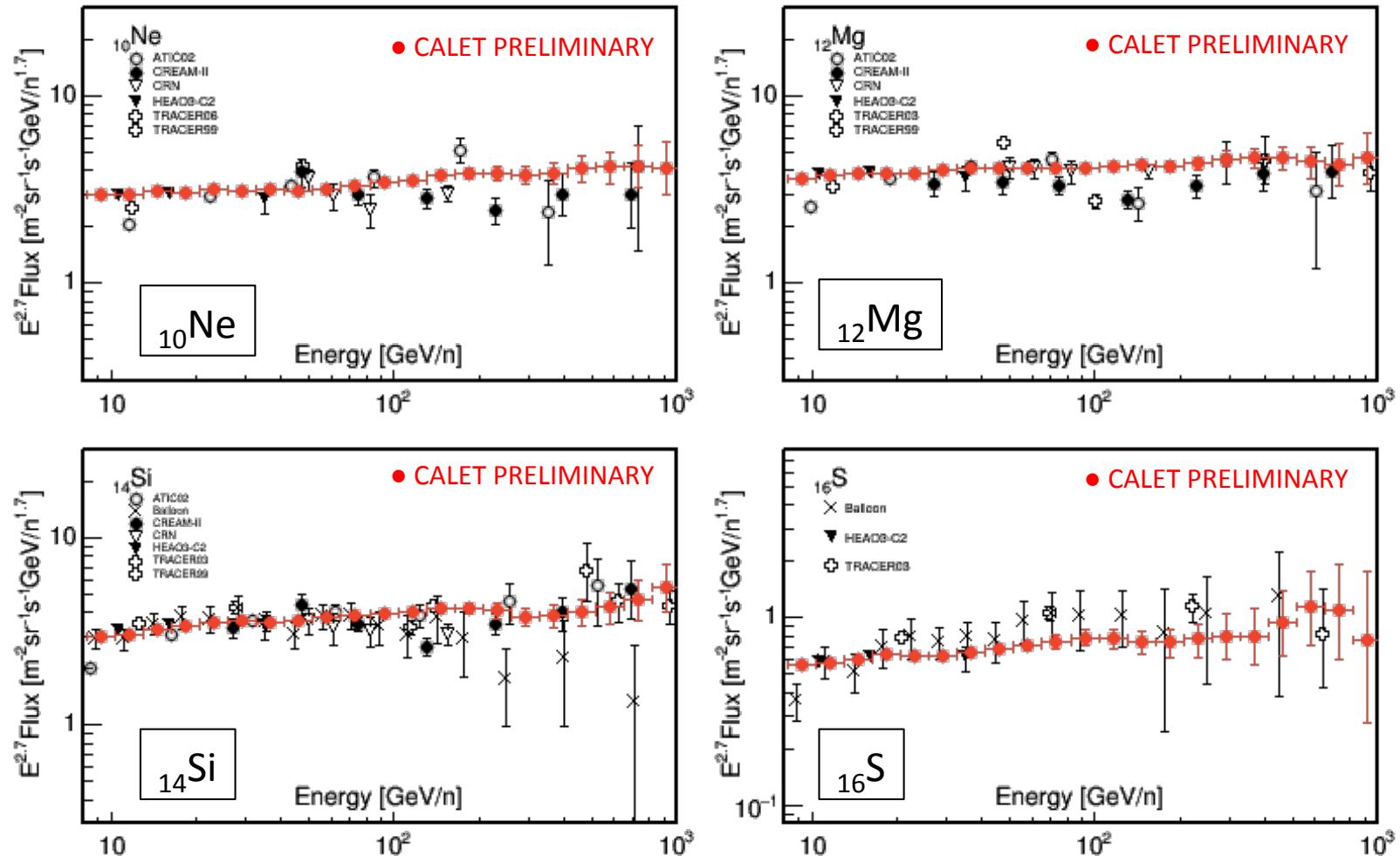
Observation period:

Oct.13 2015 – Oct.31 2018  
(750 days)



# Preliminary Spectra of Nuclei with Even Atomic Number (Z=10-16)

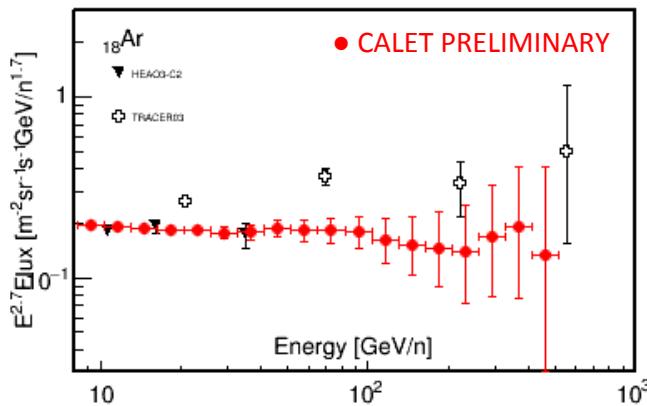
Only statistical error are shown



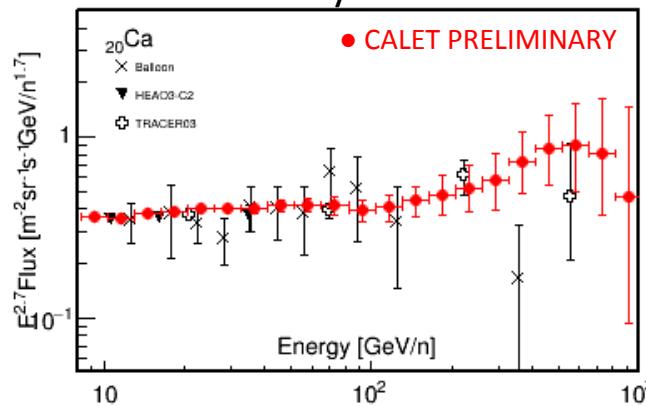
# Preliminary Spectra of Nuclei with Even Atomic Number (Z=18-28)

Only statistical error are shown

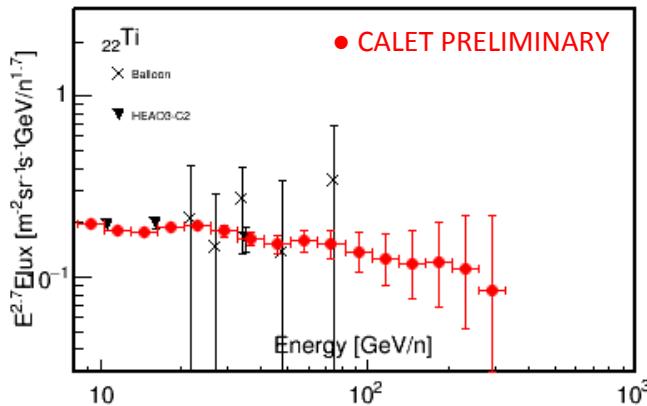
$^{18}\text{Ar}$



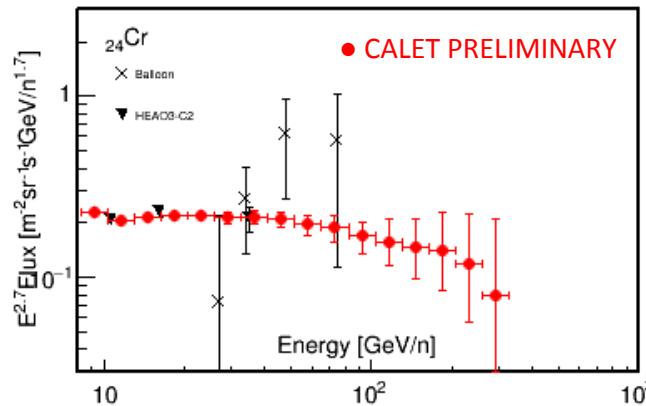
$^{20}\text{Ca}$



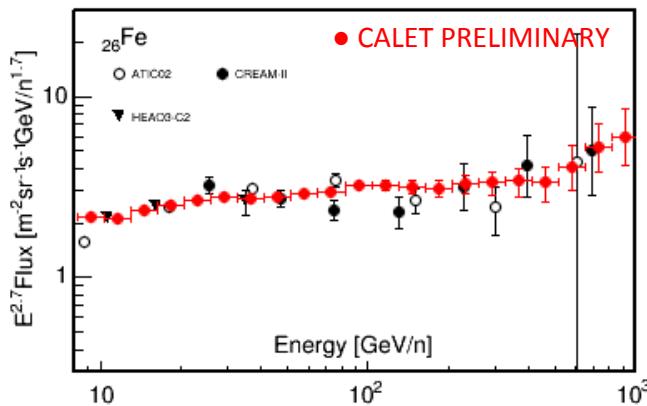
$^{22}\text{Ti}$



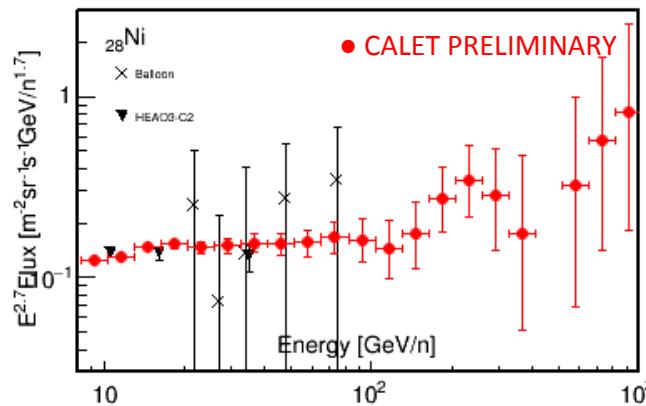
$^{24}\text{Cr}$



$^{26}\text{Fe}$



$^{28}\text{Ni}$



# Conclusion

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- The ability of CALET to measure heavy cosmic-ray nuclei has been successfully demonstrated
  - Charge resolution: 0.15 for carbon, 0.30 for iron
  - Dynamic range for energy measurement:  $1\text{-}10^6 \text{ MIP}$  ( $1\text{GeV} - 1\text{PeV}$ )
- Using data from the 750 days of operation, preliminary analysis of nuclei have successfully carried out
  - B/C ratio up to 200 GeV/n
  - primary cosmic-ray elements up to 100 TeV
- Further studies will provide the excellent energy spectra with high statistics in a wide energy range, and reveal details spectral features